

Telescope design

Steve Padin

The Submillimeter Universe: The CCAT View

12 November 2010

CCAT truss designs

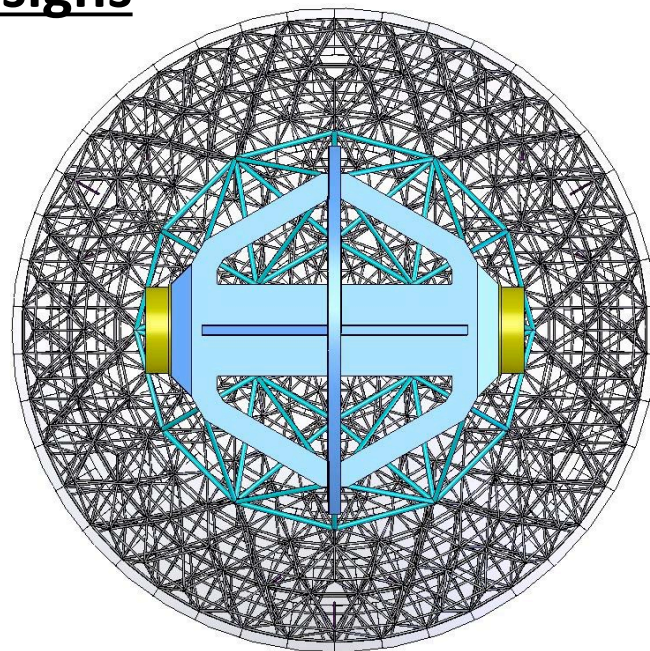
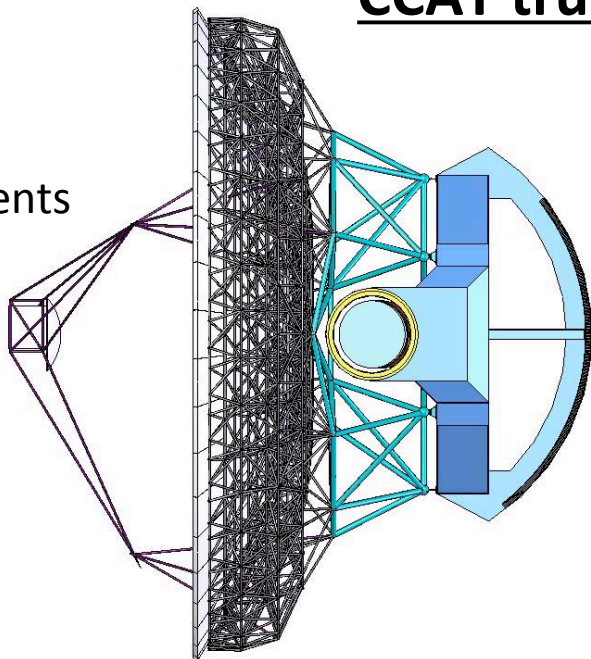
Ring & pillar

30,000 kg truss

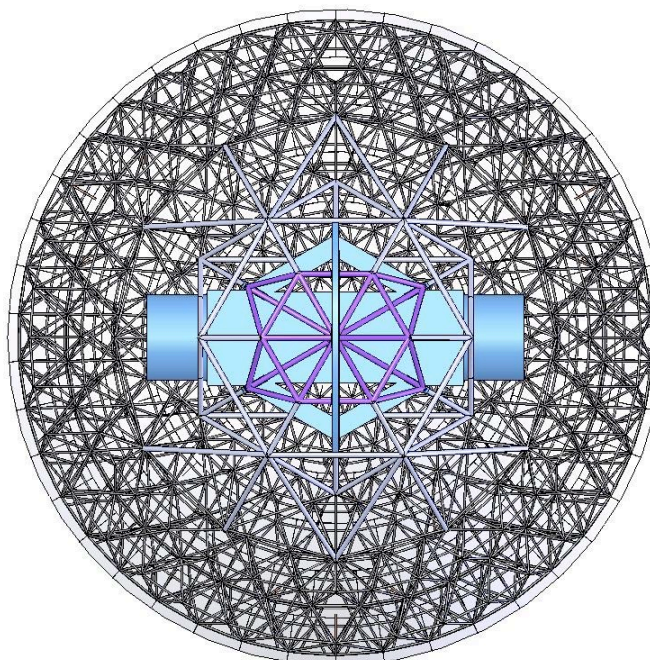
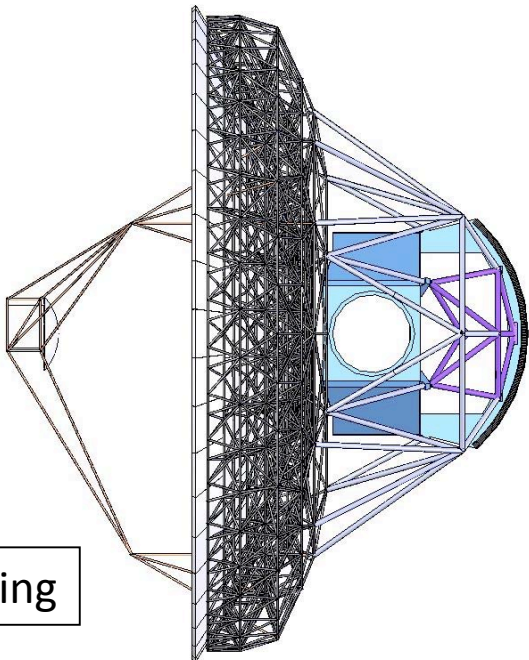
12,000 kg segments

(25 kg m⁻²)

7-8 Hz

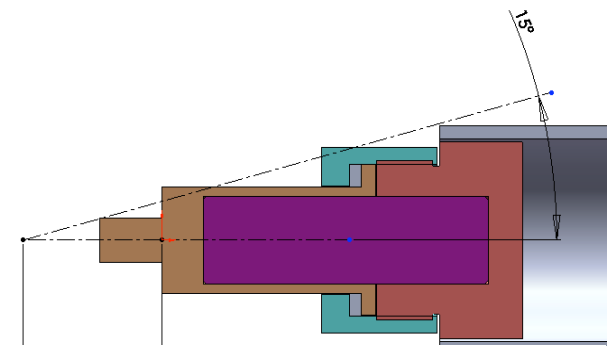
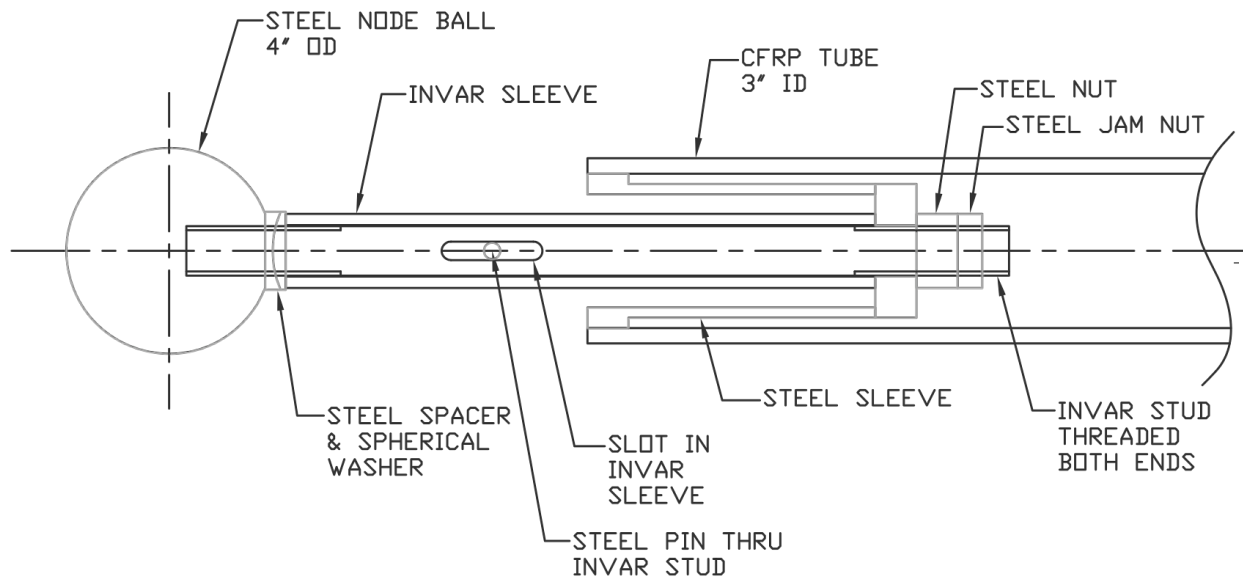


3d

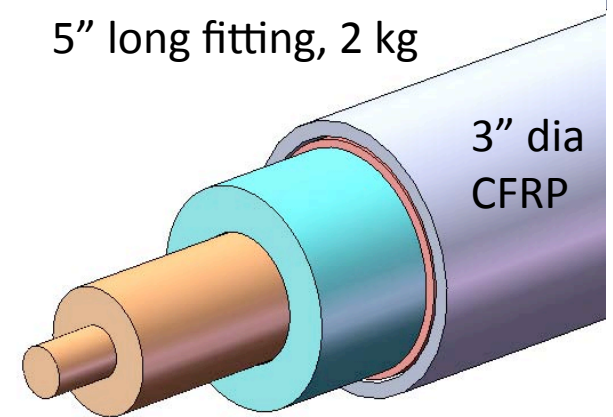


Stutzki Engineering

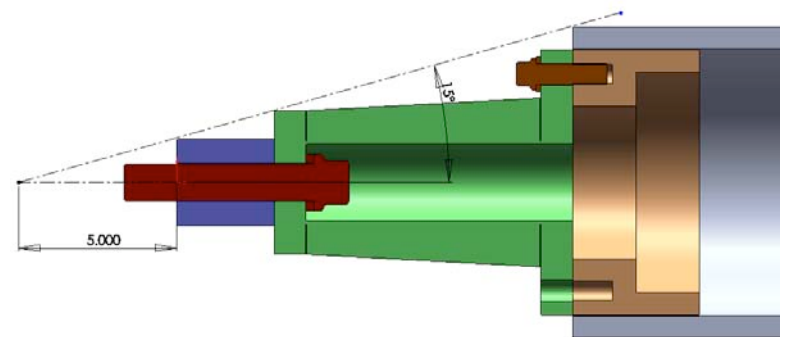
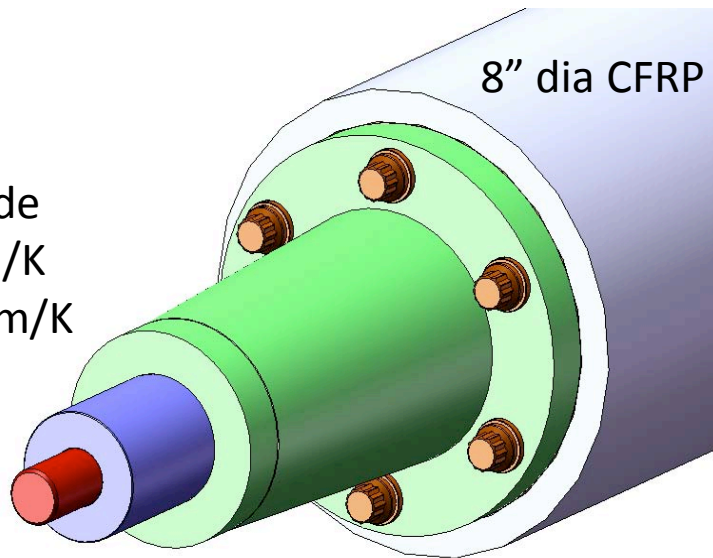
Truss rods



5" long fitting, 2 kg



For a rod & node
 $\langle \text{CTE} \rangle < 0.2 \text{ ppm/K}$
 $\Delta \text{CTE} \pm 0.03 \text{ ppm/K}$



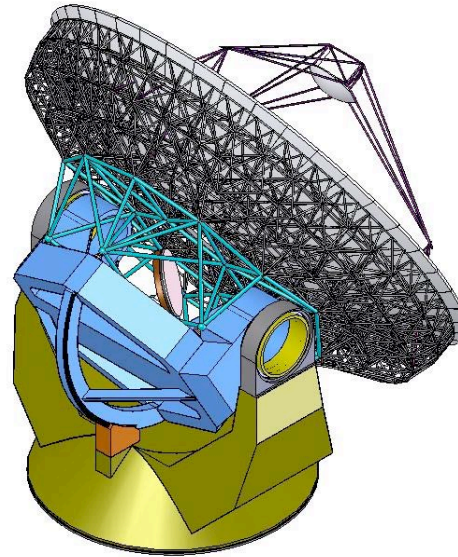
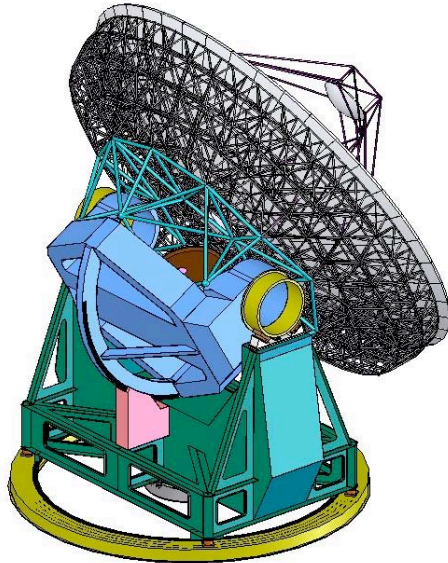
12.5" long fitting, 36 kg

ATK (COI)

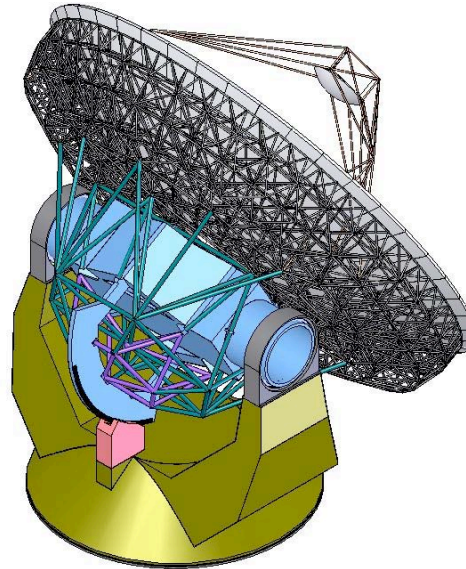
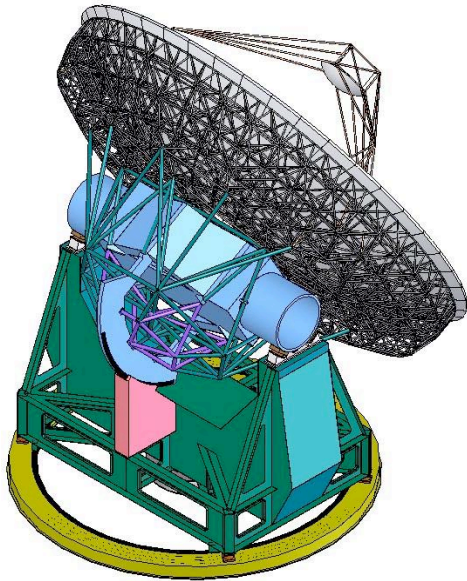
CCAT bearing & truss options

General Dynamics

Ring & pillar



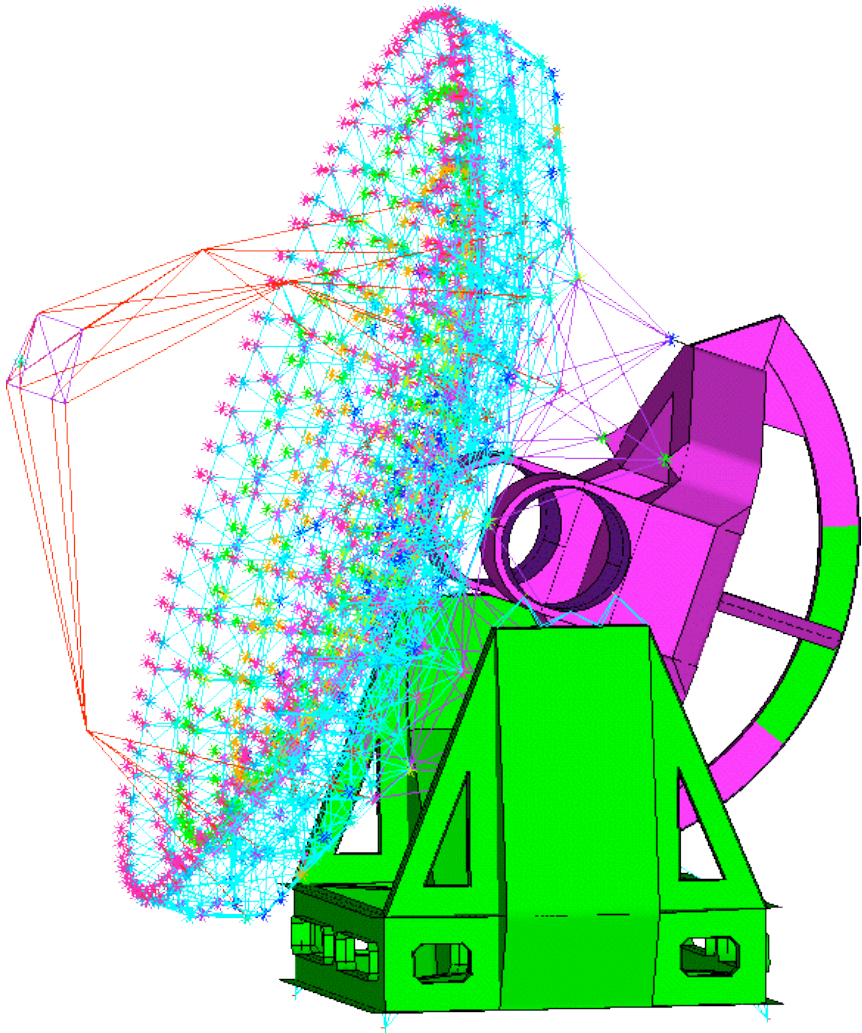
3
d



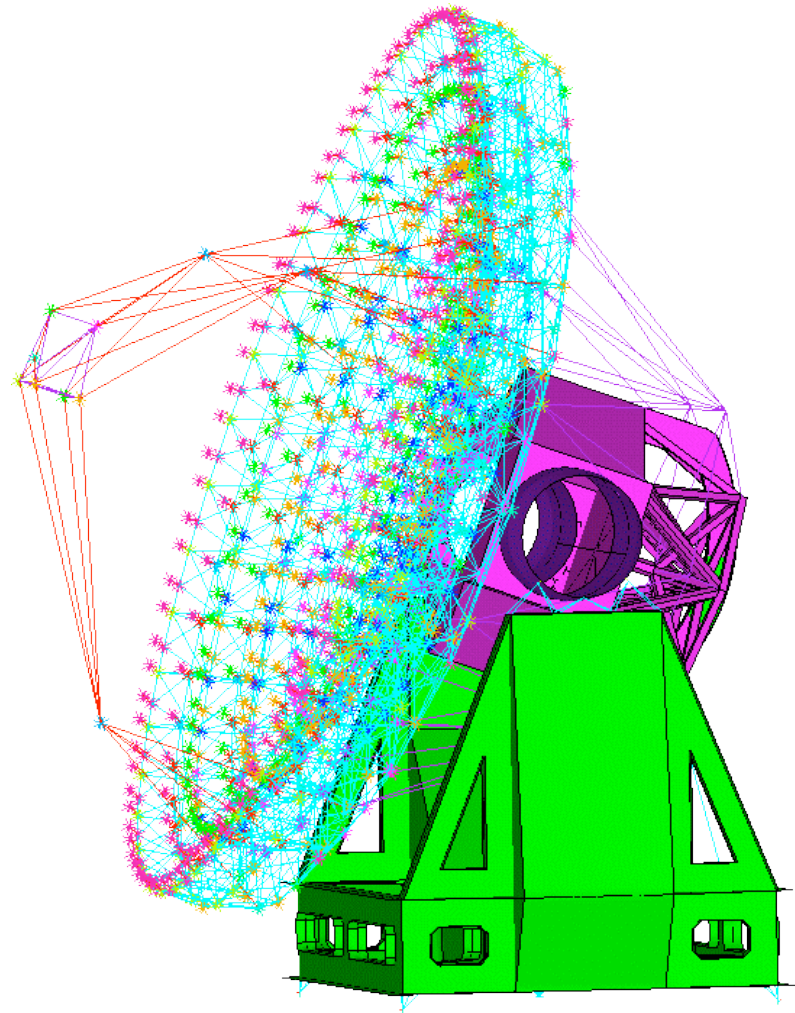
Hydrostatic bearings

Rolling element bearings

Finite element models

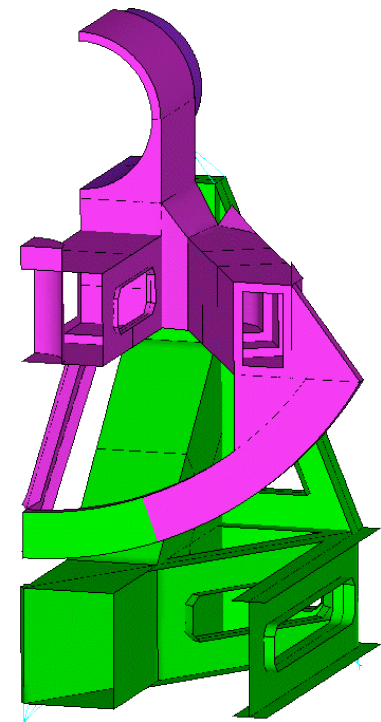
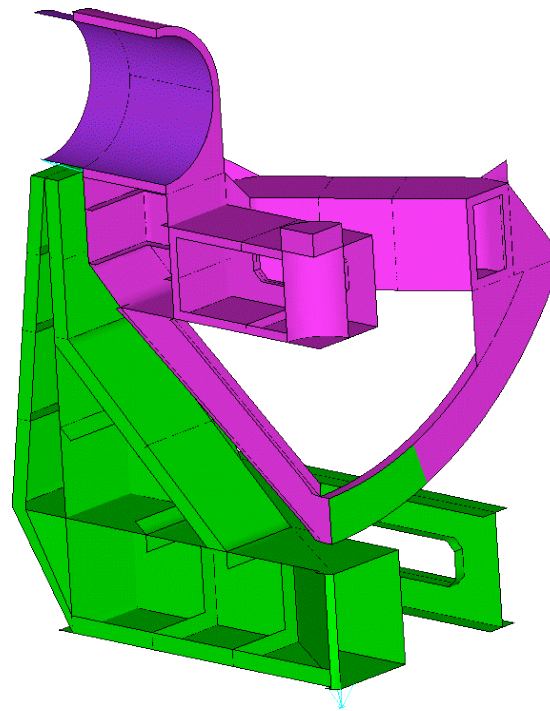
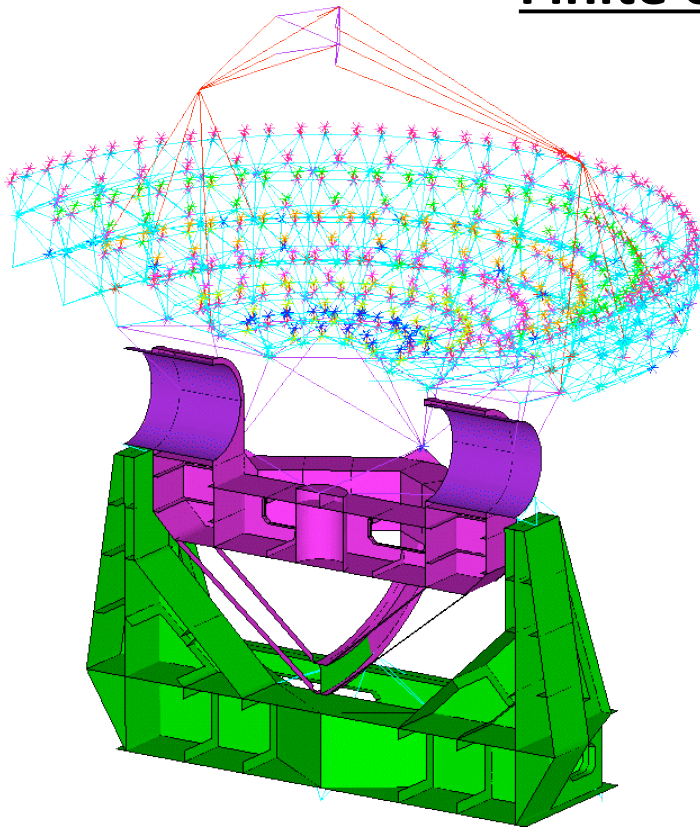


Ring & pillar

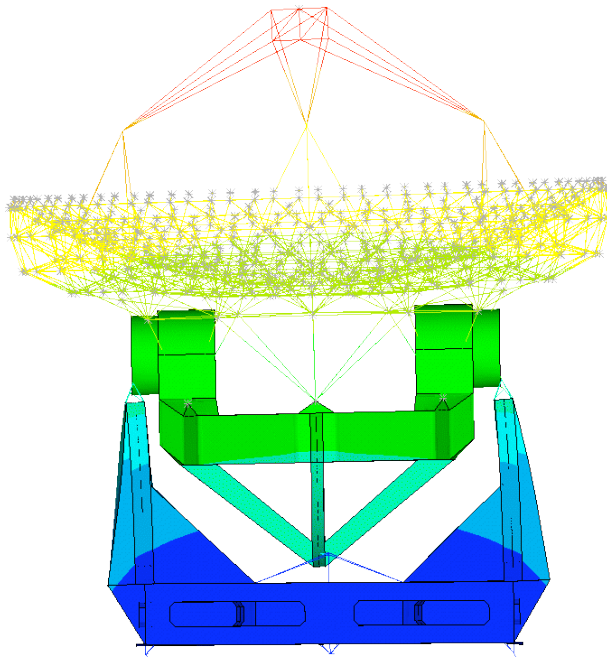


3d

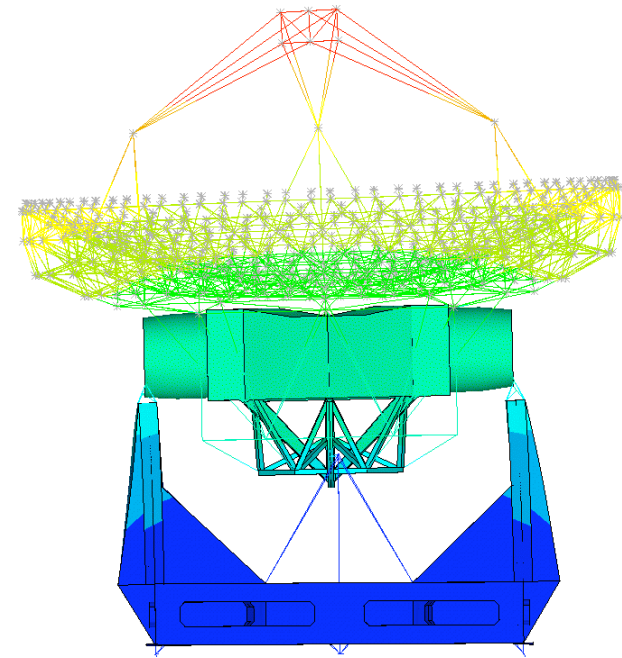
Finite element model details



FEA results



Ring & pillar 3.5 Hz



3d 3.3 Hz

Analytical model
μm rms

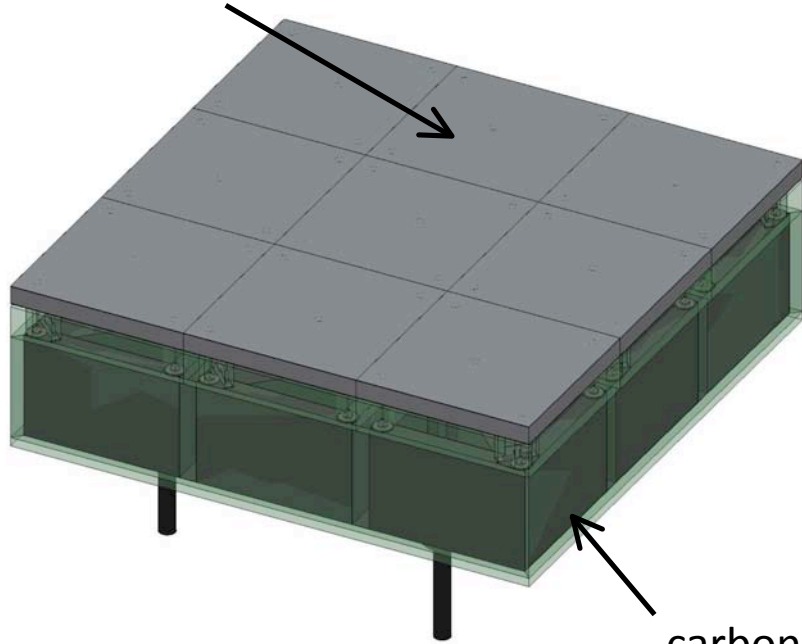
Load case	zenith	EL=20°
	μm rms	μm rms
Gravity	287	276
0.4°s ⁻² acceleration	1.7	1.3
20K soak	3.8	5.0
1K vertical	0.2	0.7
1K horizontal	0.9	0.9

0.7
3.1
3.3

Load case	zenith	EL=20°
	μm rms	μm rms
gravity	360	187
0.4°s ⁻² acceleration	2.3	1.5
20K soak	69.2(!)	69.2
1K vertical	0.7	1.7
1K horizontal	1.8	1.8

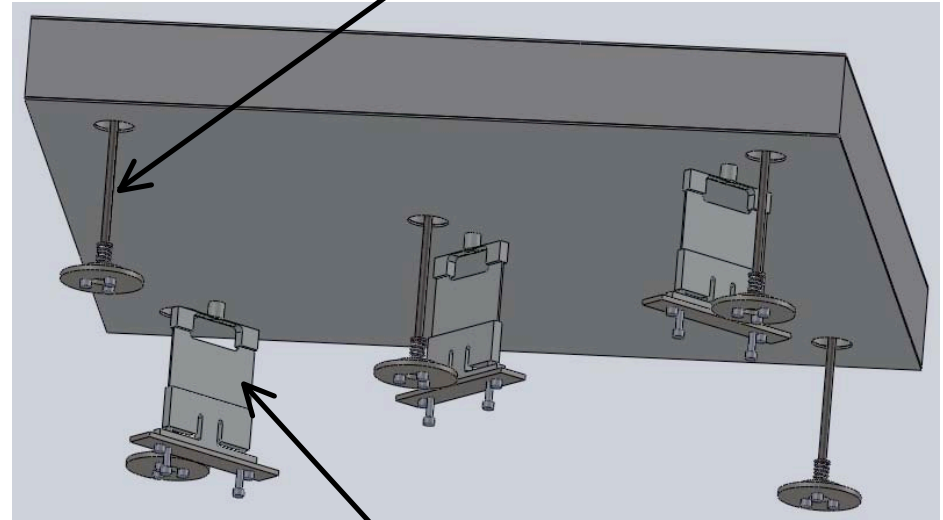
CCAT segment concept

aluminum tiles

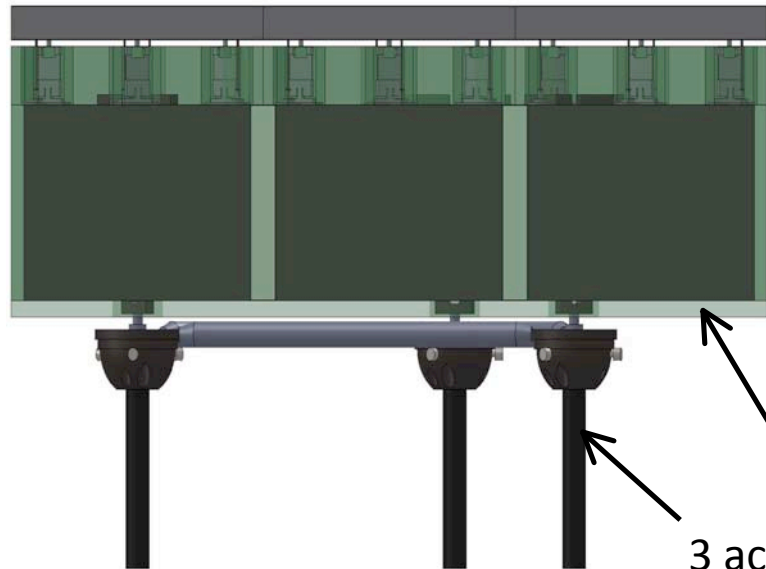


carbon-fiber frame

5 vertical adjusters

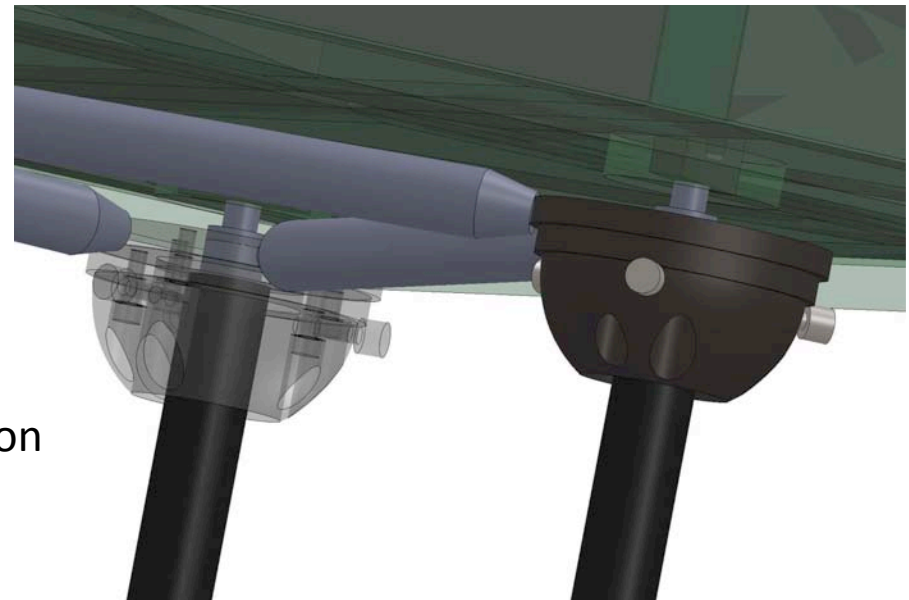


3 horizontal adjusters



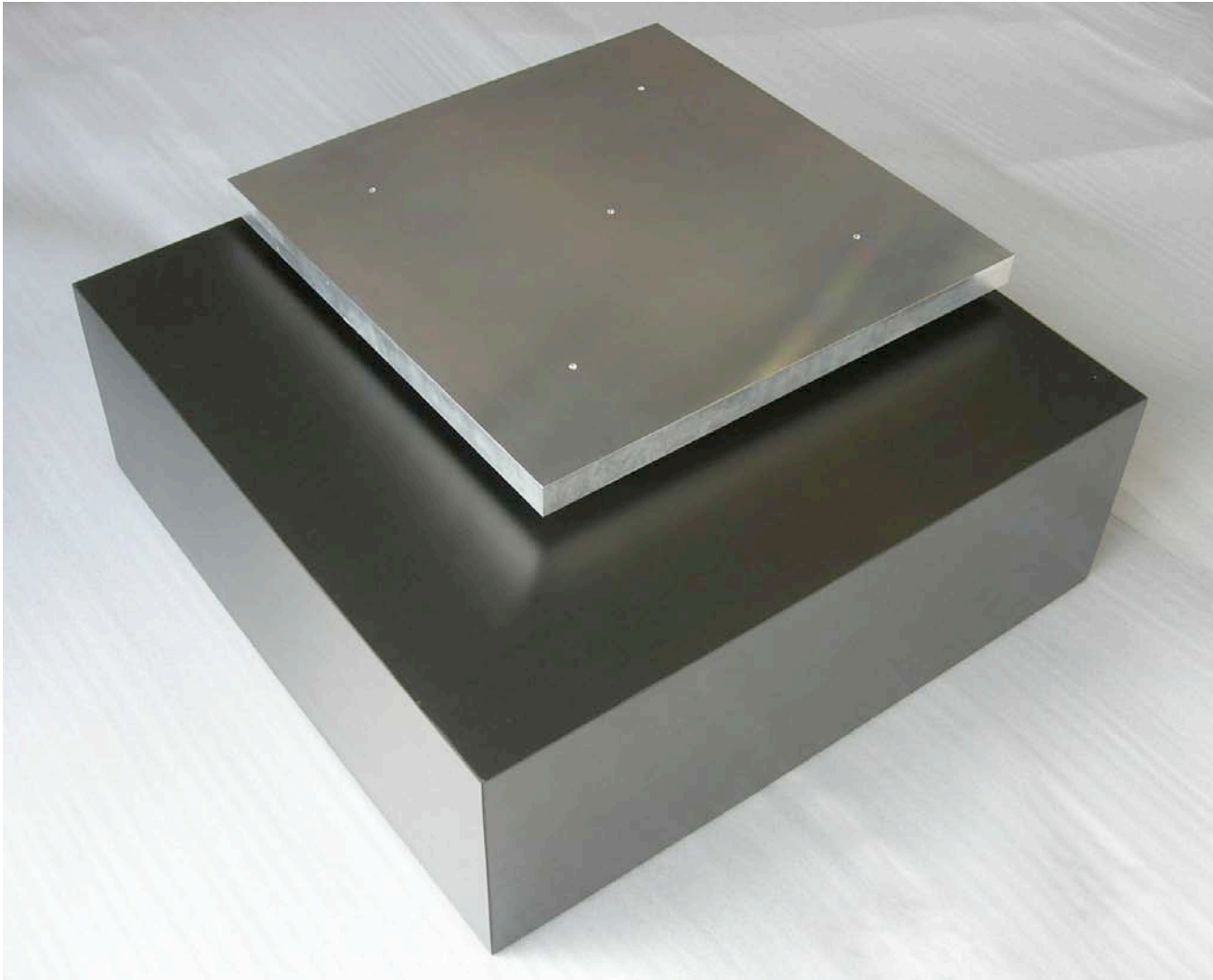
insulation

3 actuators



Prototype tile & subframe

Vertex Antennentechnik

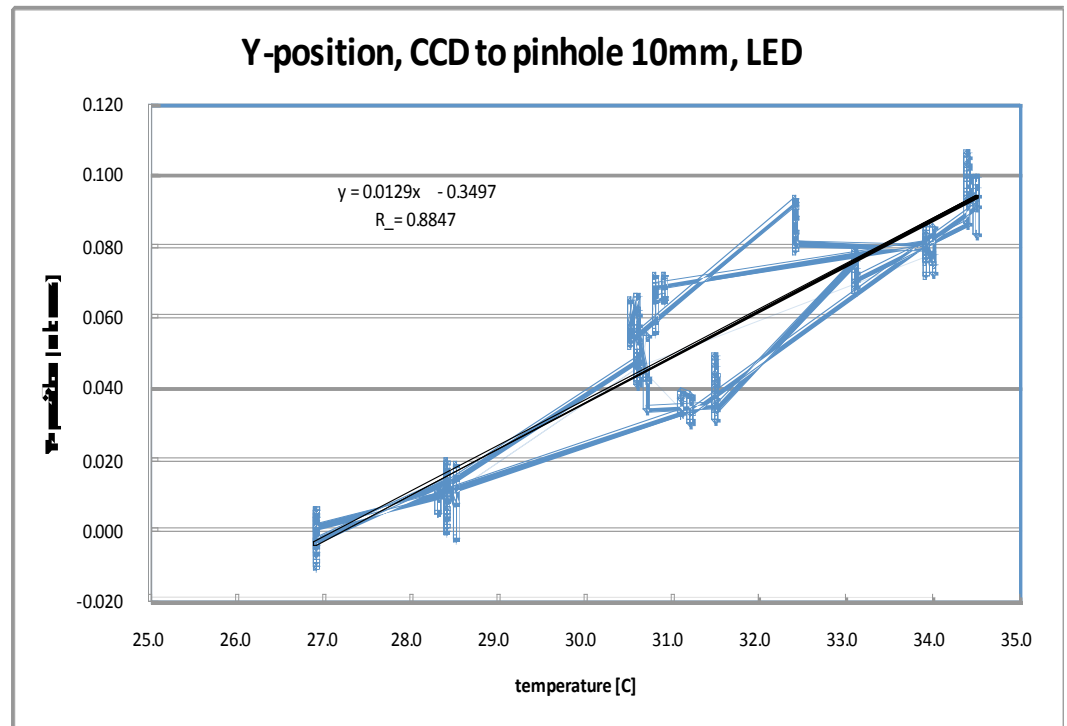
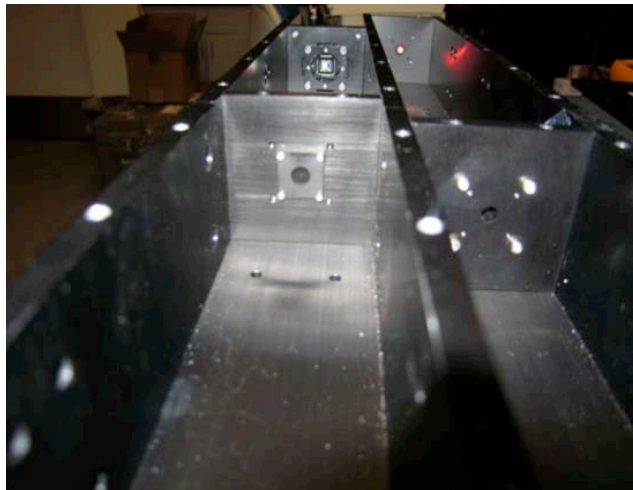
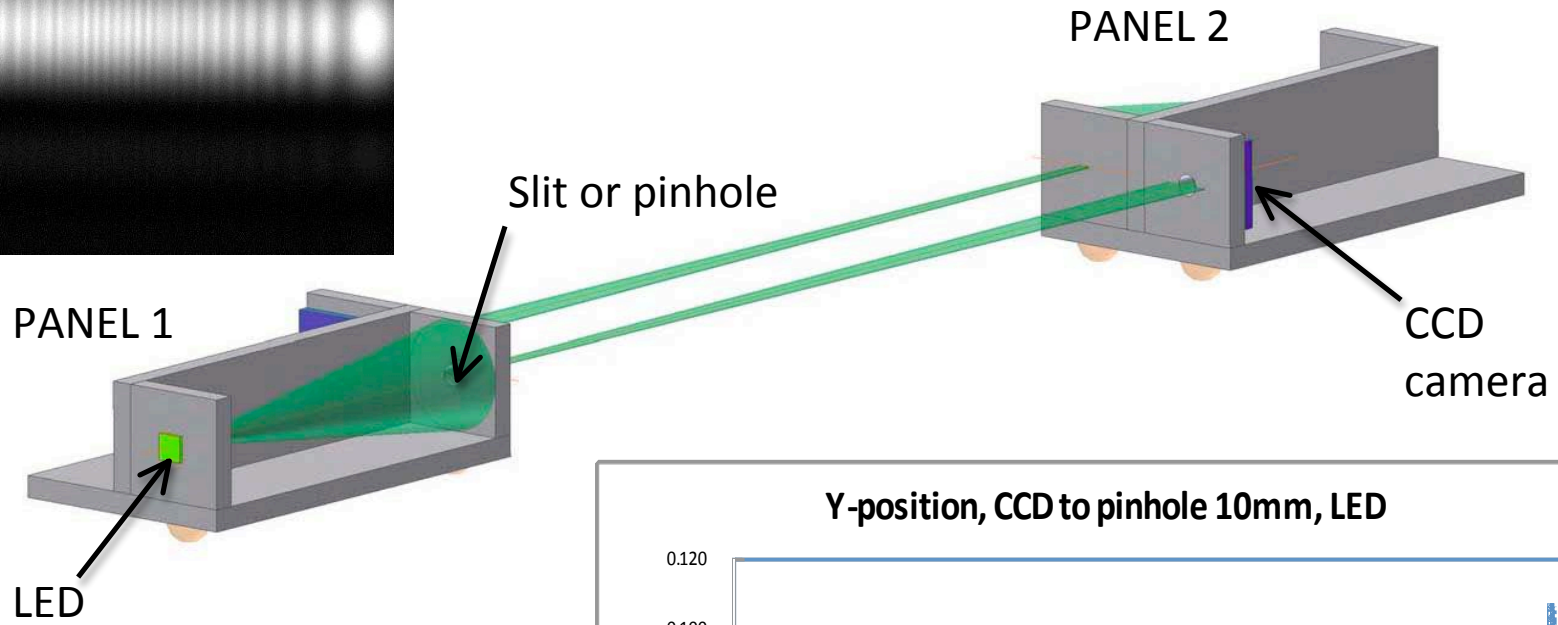
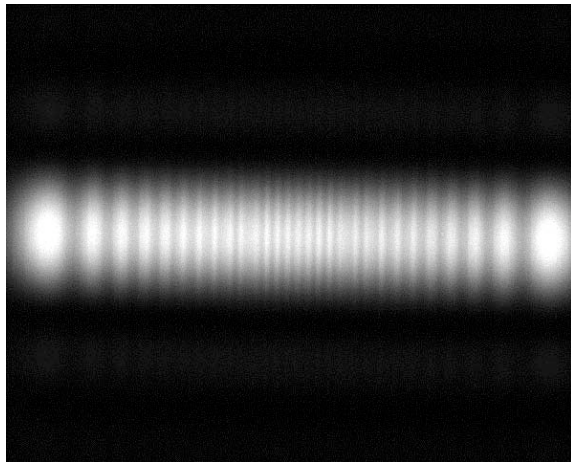


Tile surface error $2.5 \mu\text{m rms}$

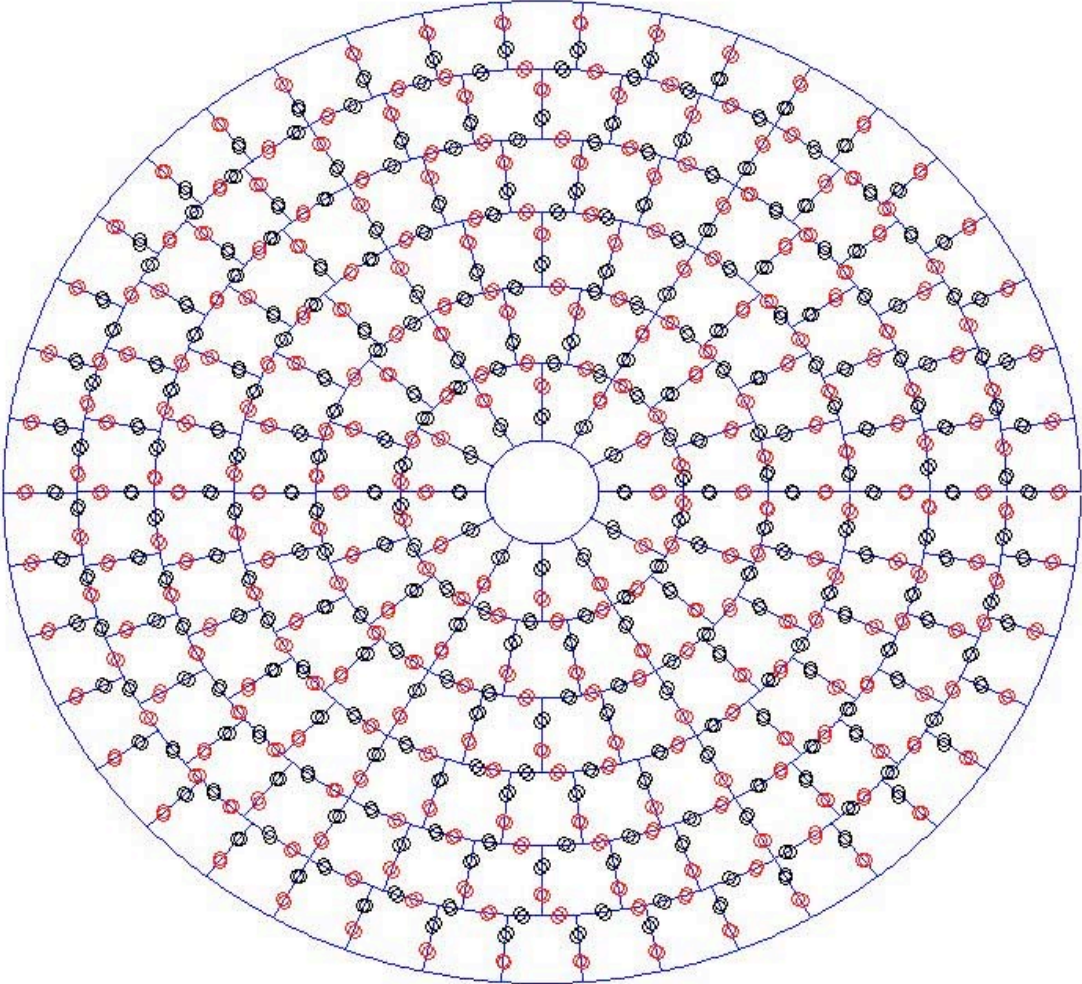
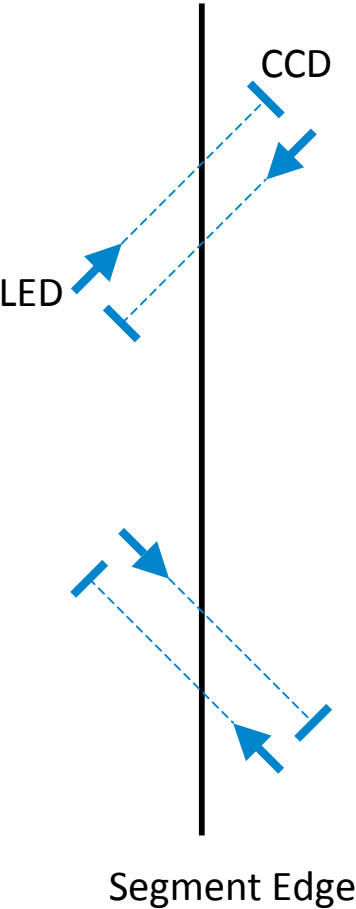
Tile + subframe areal density 19 kg m^{-2}

Optical edge sensor

D. Woody, Caltech



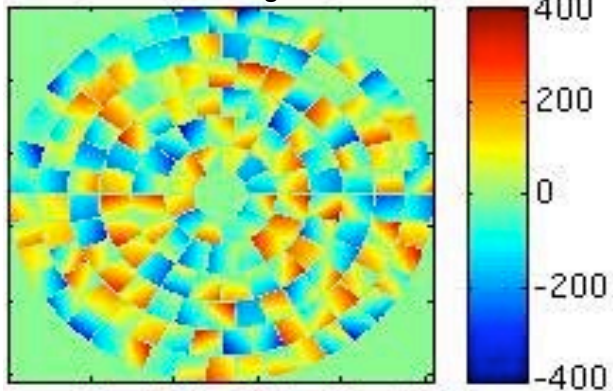
Edge sensor placement



Primary surface control

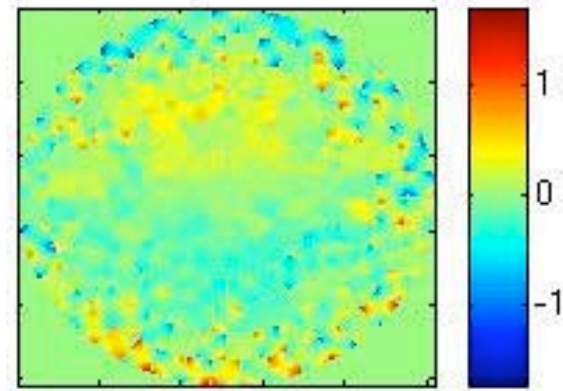
J. Lou, JPL

CCAT WF Error with randomly
Perturbed segments



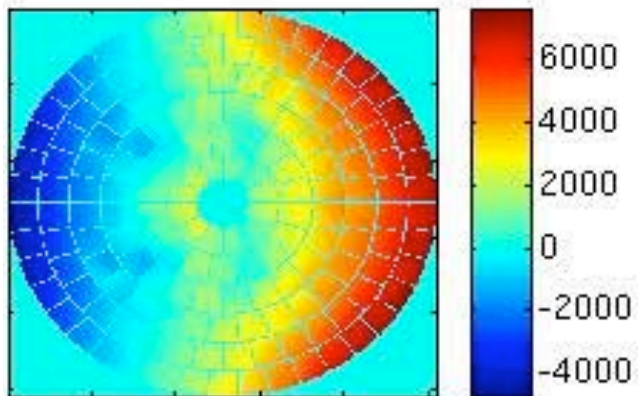
RMS WFE = 143.5 μm

CCAT Controlled WF Error with
Global Modes Removed



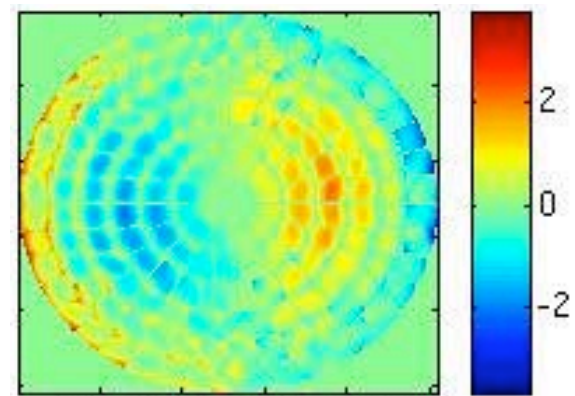
RMS WFE = 0.28 μm

CCAT WF Error with
1g load



RMS WFE = 2904 μm

CCAT Controlled WF Error with
Global Modes Removed

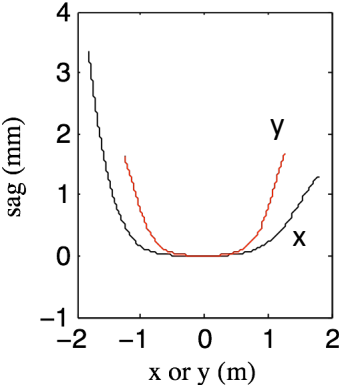
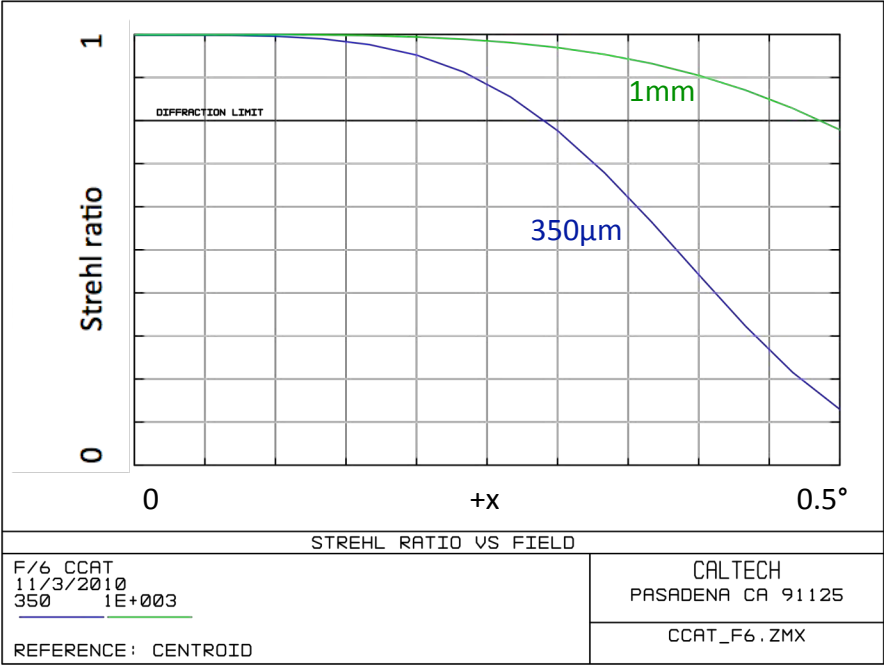
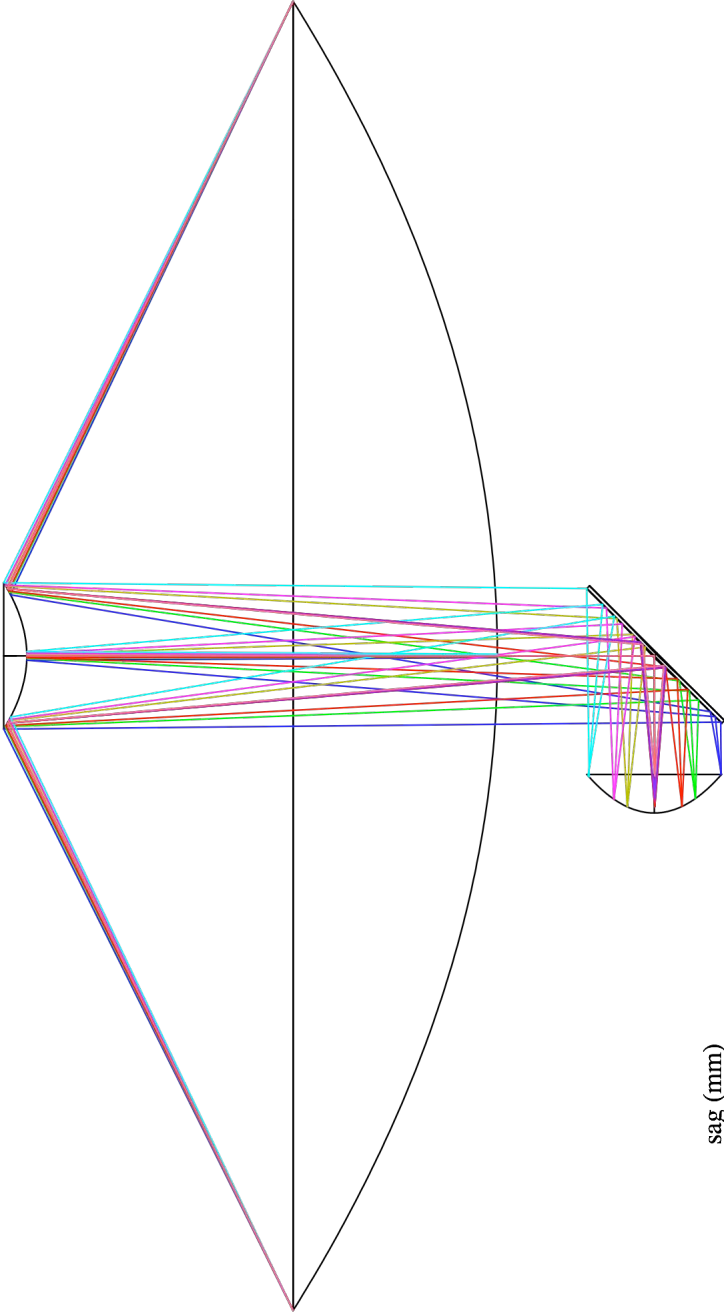


RMS WFE = 0.77 μm

no sensor noise

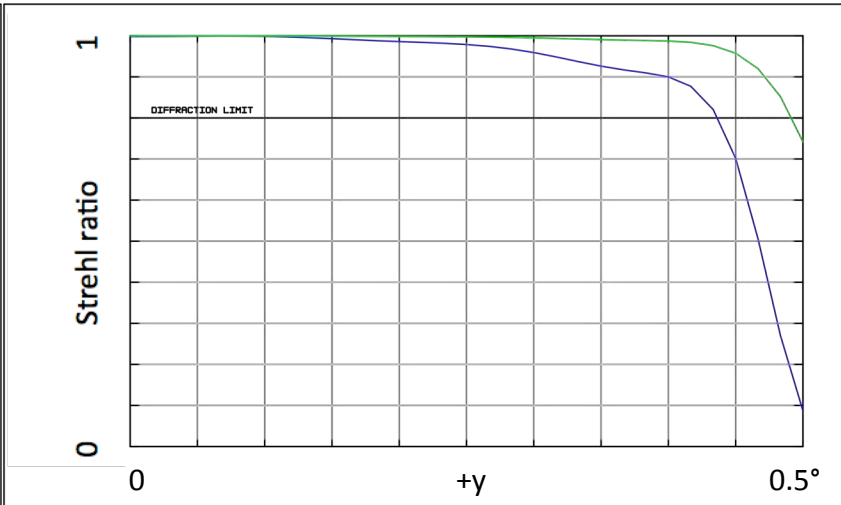
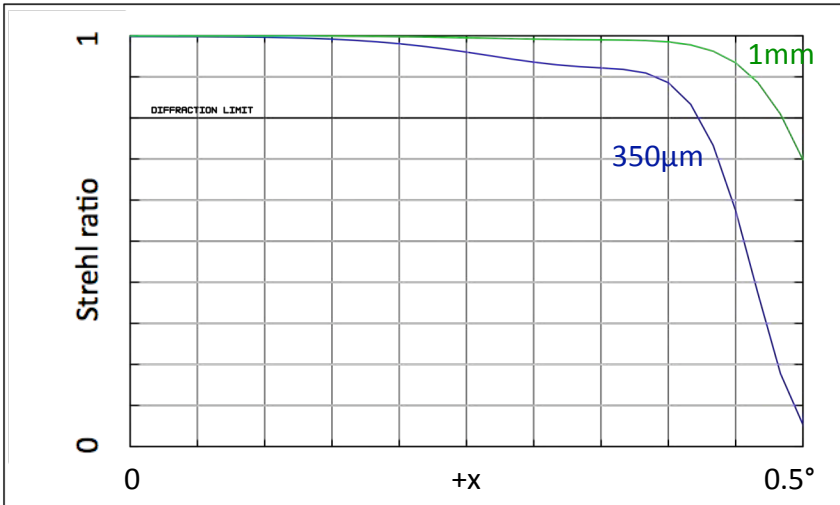
CCAT optics

Strehl ratio vs. field angle for basic RC telescope



Tertiary profile to maximize FoV

Strehl ratio vs. field angle with tertiary corrector

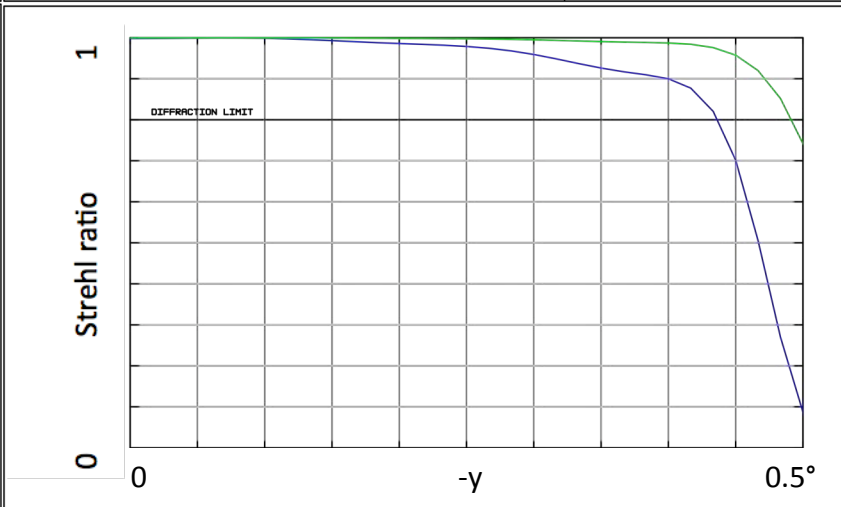
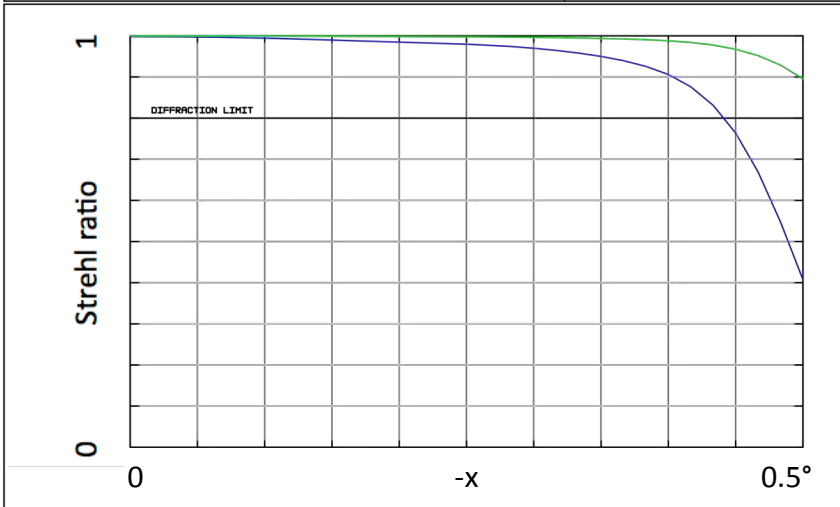


STREHL RATIO VS FIELD

F/6 CCAT WITH CORRECTOR 11/3/2010 350 1E+003	CALTECH PASADENA CA 91125
REFERENCE: CENTROID	CCAT_F6_TERTIARY_CORR.ZMX

STREHL RATIO VS FIELD

F/6 CCAT WITH CORRECTOR 11/3/2010 350 1E+003	CALTECH PASADENA CA 91125
REFERENCE: CENTROID	CCAT_F6_TERTIARY_CORR.ZMX



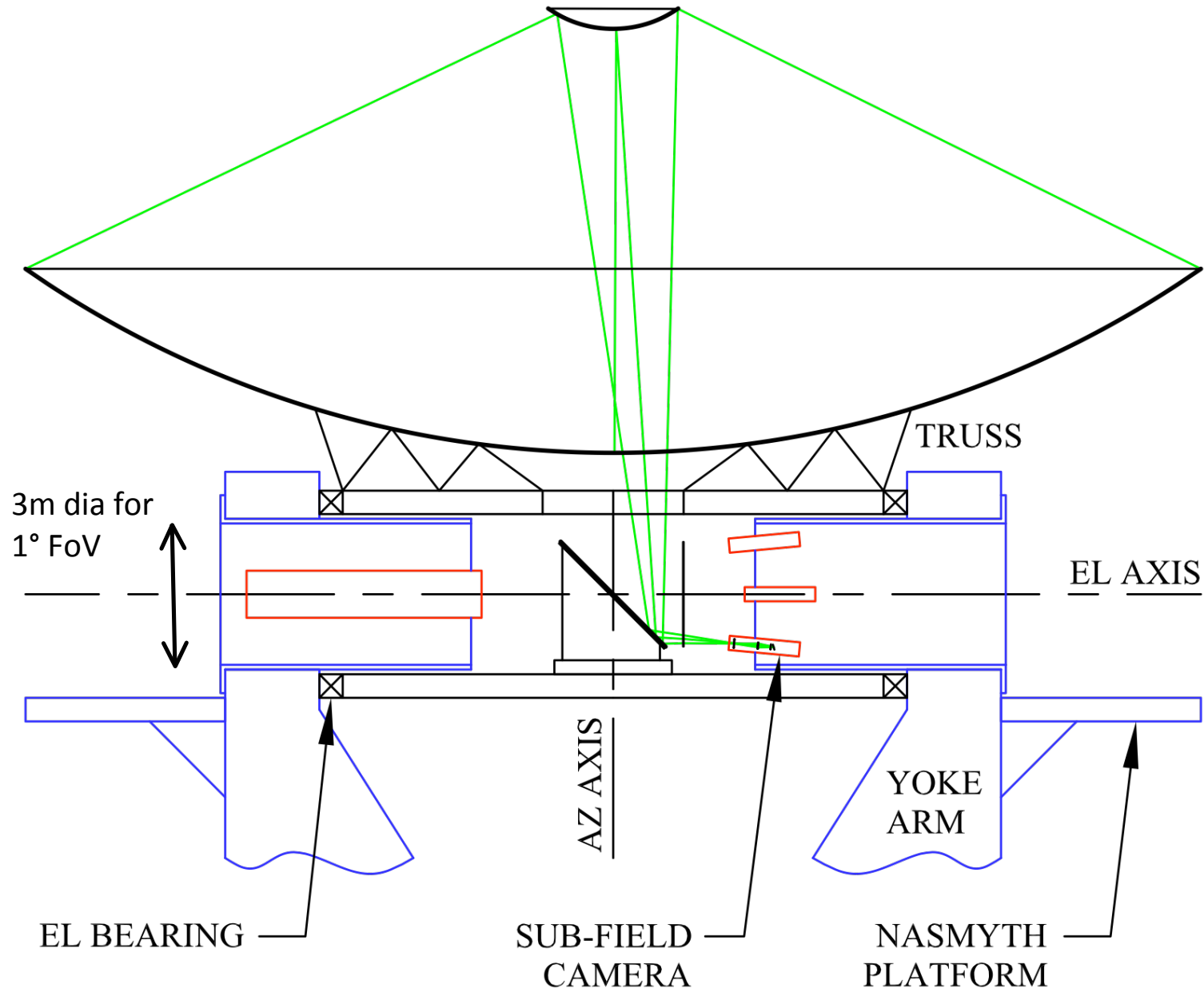
STREHL RATIO VS FIELD

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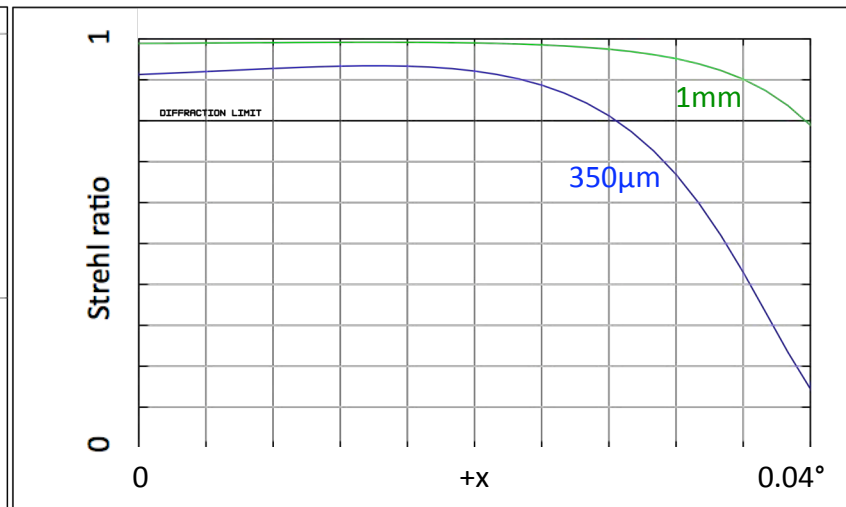
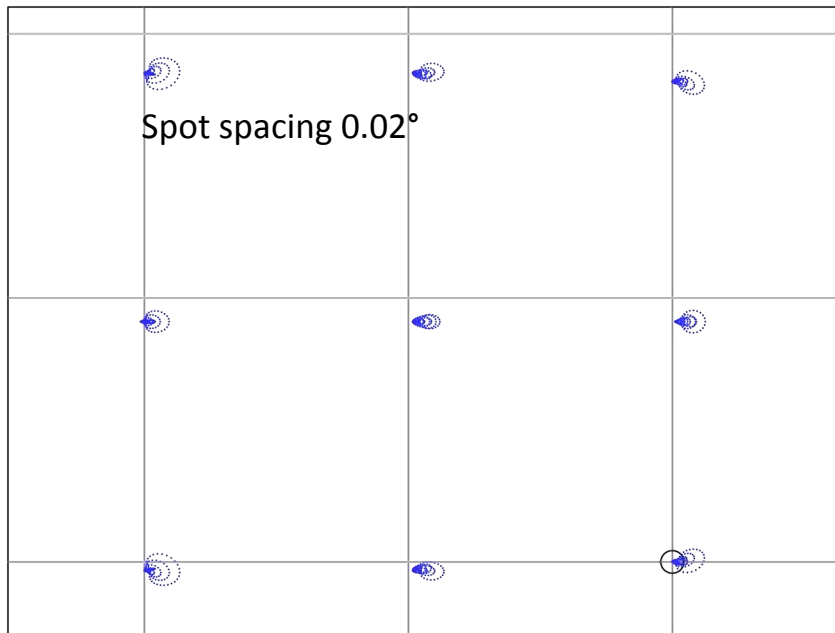
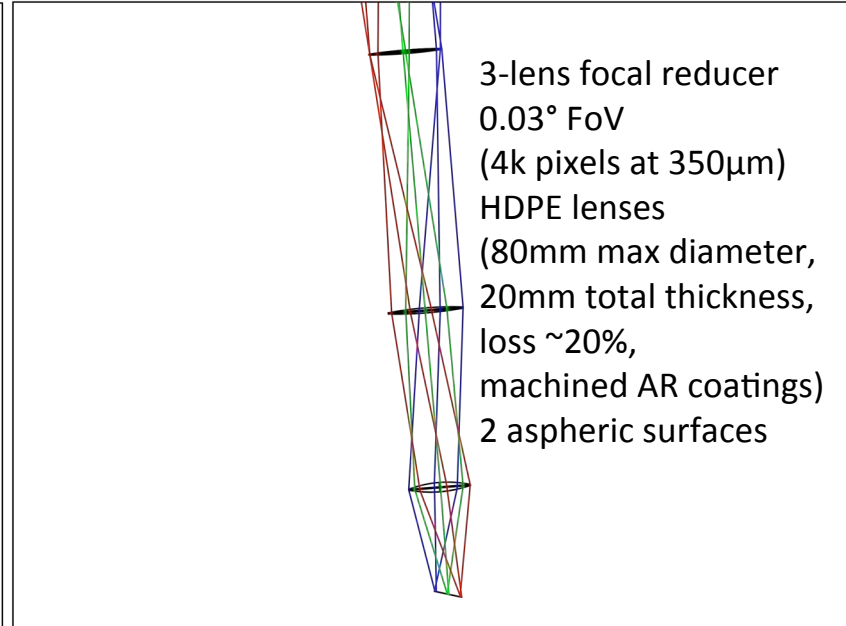
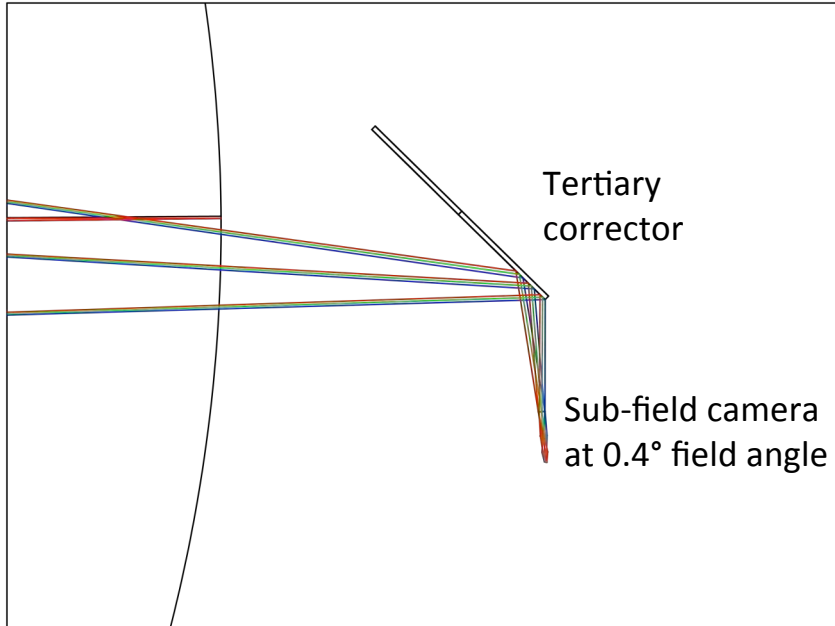
STREHL RATIO VS FIELD

F/6 CCAT WITH CORRECTOR 11/3/2010 350 1E+003	CALTECH PASADENA CA 91125
REFERENCE: CENTROID	CCAT_F6_TERTIARY_CORR.ZMX

Instrument interface



Camera relay with small HDPE lenses



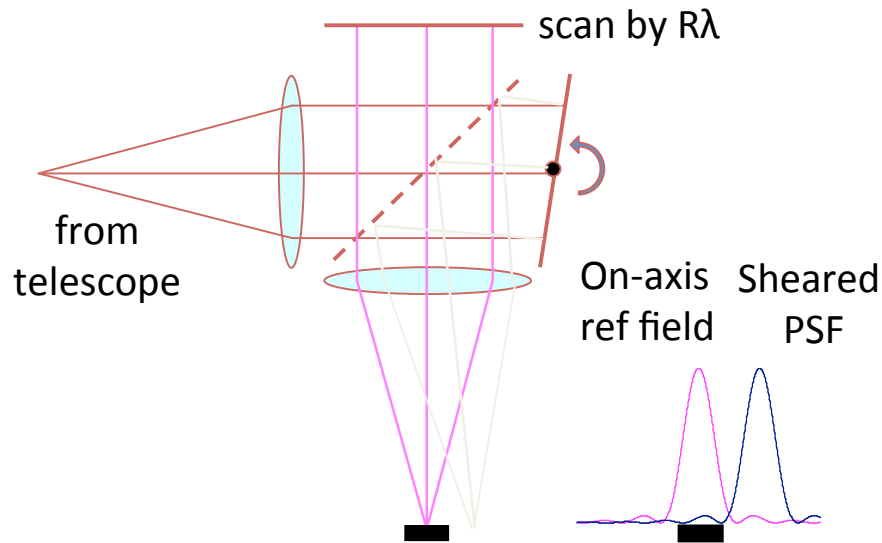
STREHL RATIO VS FIELD

F/6 CCAT WITH TERTIARY CORRECTOR 11/5/2010 350 1E+003	CALTECH PASADENA CA 91125
REFERENCE: CENTROID	CCAT_F6_TERTIARY CORR_3BCLENS_OFFSET.ZMX

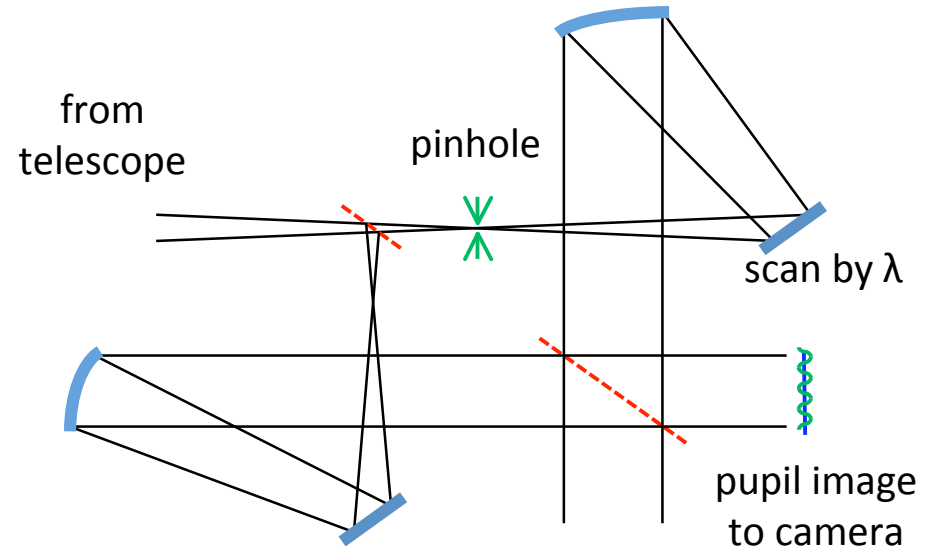
Alignment WFS

G. Serabyn, JPL

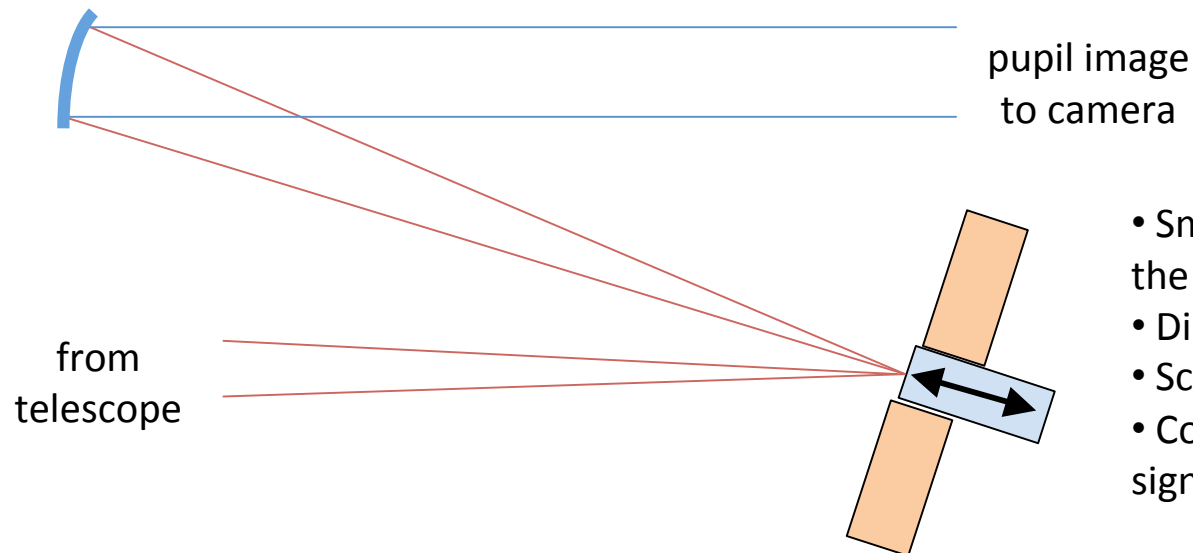
CSO shearing interferometer



Point diffraction interferometer



Phase contrast interferometer



- Small scanning mirror in the center of a large mirror
- Diameter matches PSF core
- Scan by $\sim \lambda$
- Common path for ref & signal gives good stability

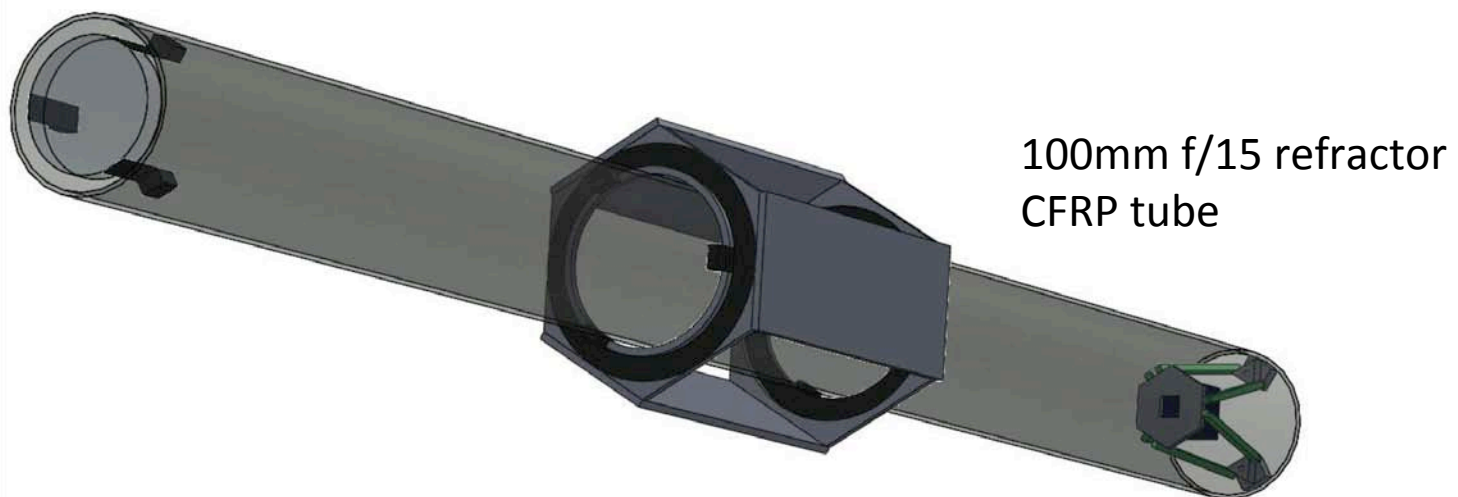
CCAT performance

Contribution	HWFE ($\mu\text{m rms}$)	PE (arcsec rms)	Notes
Aberrations	3.65	0.00	Ritchey-Chretien design
Primary open-loop	7.21	0.03	CFRP truss
Primary closed-loop	7.48	0.52	Steel truss
Secondary	6.28	0.20	
Tertiary	4.45	0.10	
Instrument	0.05	0.04	
Mount	0.00	0.09	No HWFE from mount
Alignment	2.19	0.10	Regular pointing with science camera, occasional wavefront measurements with WFS
Telescope RSS open loop	11.37	0.26	
Telescope RSS closed-loop	11.55	0.58	
Telescope requirement	10.00	0.35	<50% increase in integration time, PE<1/10th beam, from CCAT-TM-48
Atmosphere	5.74	0.23	1st quartile

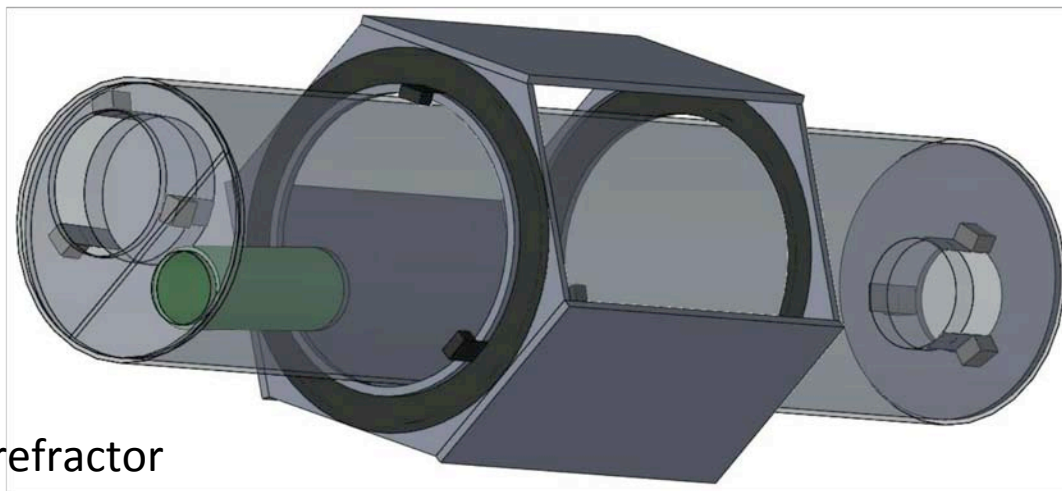
Observatory parameters

Parameter	Value	Units	Notes
Wavelength	350	μm	
Mean outside wind speed	6	m/s	3rd quartile
Wind speed for pressure	8.484	m/s	$2^{1/2}v_{\text{outside}}$, see CCAT-TM-56
Density of air	0.7	kg m^{-3}	At 5600 m altitude
Scan acceleration	0.4	deg s^{-2}	0.4 deg s^{-2} if $\lambda < 620\mu\text{m}$, else 2 deg s^{-2} , from CCAT-TM-48
rms temp gradient in dome	1	K	From TMT CFD
Soak temp change	20	K	Diurnal & longer
Flux density of pointing source	0.1	Jy	For >1 source/ deg^2 , $S < 0.3\text{Jy}$ at $\lambda = 350\mu\text{m}$ and $S < 40\text{mJy}$ at $\lambda = 850\mu\text{m}$, see Fig. 4.9 in feasibility study
Pointing integration time	120	s	< a few min for reasonable observing efficiency
Field angle	0.08	deg	

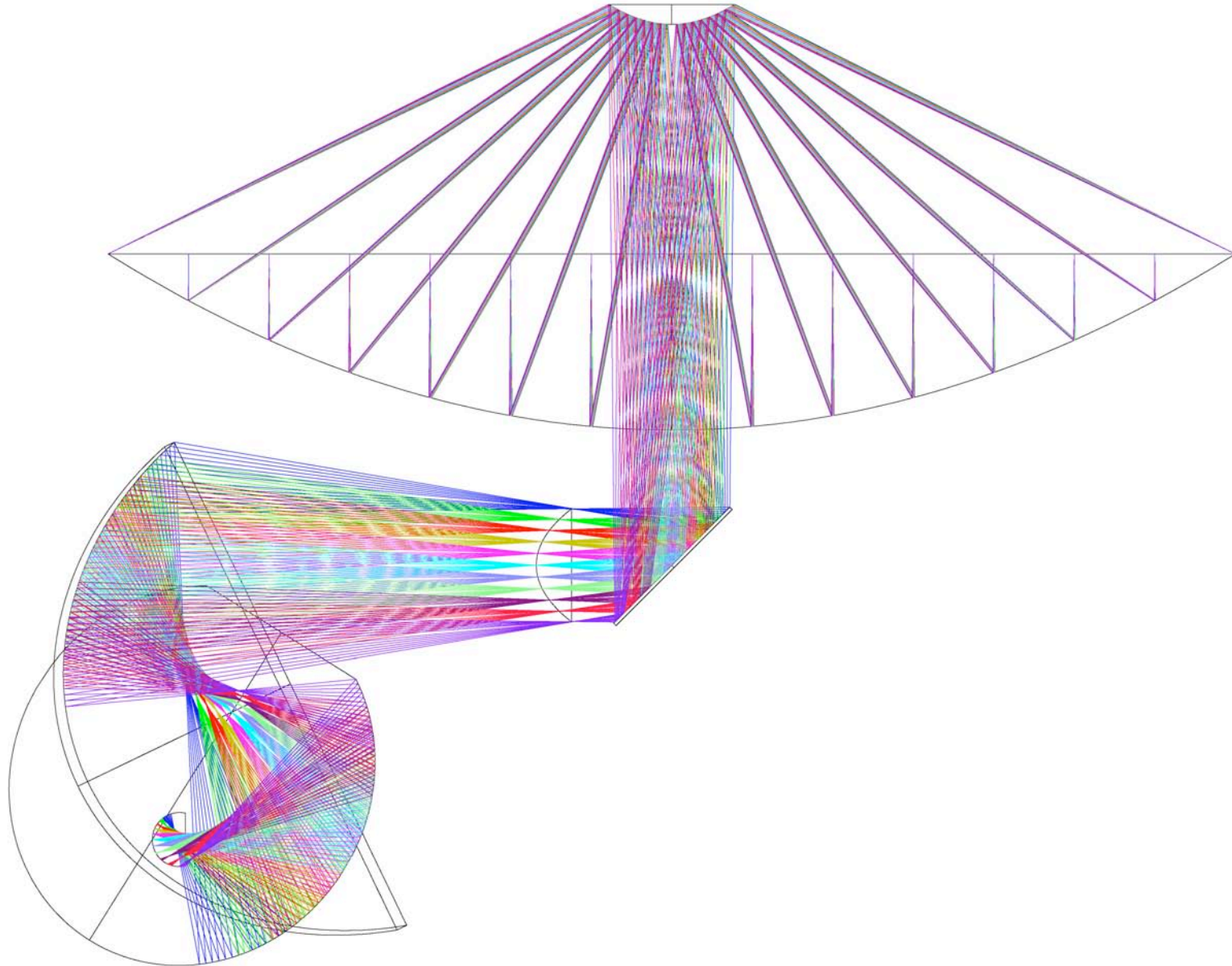
Pointing telescope concepts



Folded refractor



Reflective relay for 1° FoV



0.5° FoV camera optics

