

# Z-SPEC

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## A Broadband Millimeter-wave Grating Spectrometer

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## Case for broadband millimeter-wave spectroscopy

Redshift measurements of dusty submillimeter galaxies using mid-J CO rotational lines

For  $z > 0.9$ , at least 2 CO transitions available in Z-Spec's bandpass

Does not rely on detection at other frequencies, reducing selection effects

Systematic line surveys of LIRGs, ULIRGs, and nearby starbursts

High-density tracers, CO cooling, optically thin isotopes are simultaneously accessible :  $^{12}\text{CO}$ ,  $^{13}\text{CO}$ , HCN, CN....

Complementary data to lower-J transitions previously obtained at radio frequencies

# Requirements

**Broadband** to cover entire 1mm atmospheric transmission window (190 – 310 GHz)

**Diffraction grating** spectrometer with moderate resolution (R 250 ~ 400)

**Sensitivity** limited by sky & telescope background

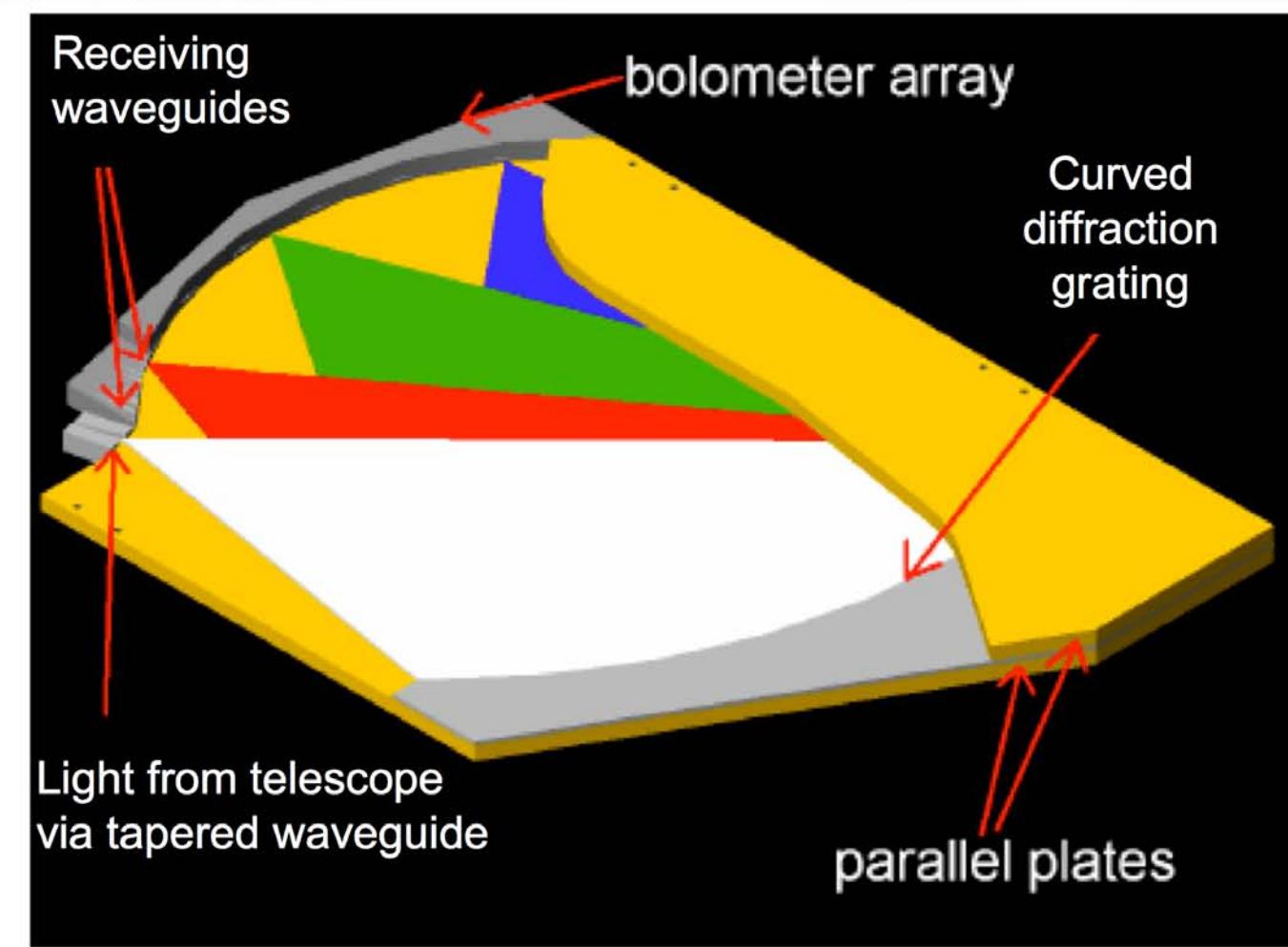
**Cold** → bolometers and entire grating cooled to 70 mK

**Space-efficient** → minimum mechanical envelope

**Concept demonstration** scalable to higher frequencies

D  
E  
S  
I  
G  
N

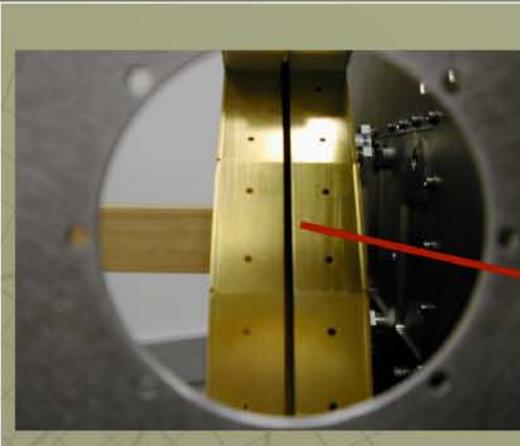
## Curved grating in a 2-D envelope (WaFIRS)



# D E S I G N

## WaFIRS module in Z-Spec

**Stackable**  
can  
multiplex  
spatially  
or in  
different  
frequency  
bands



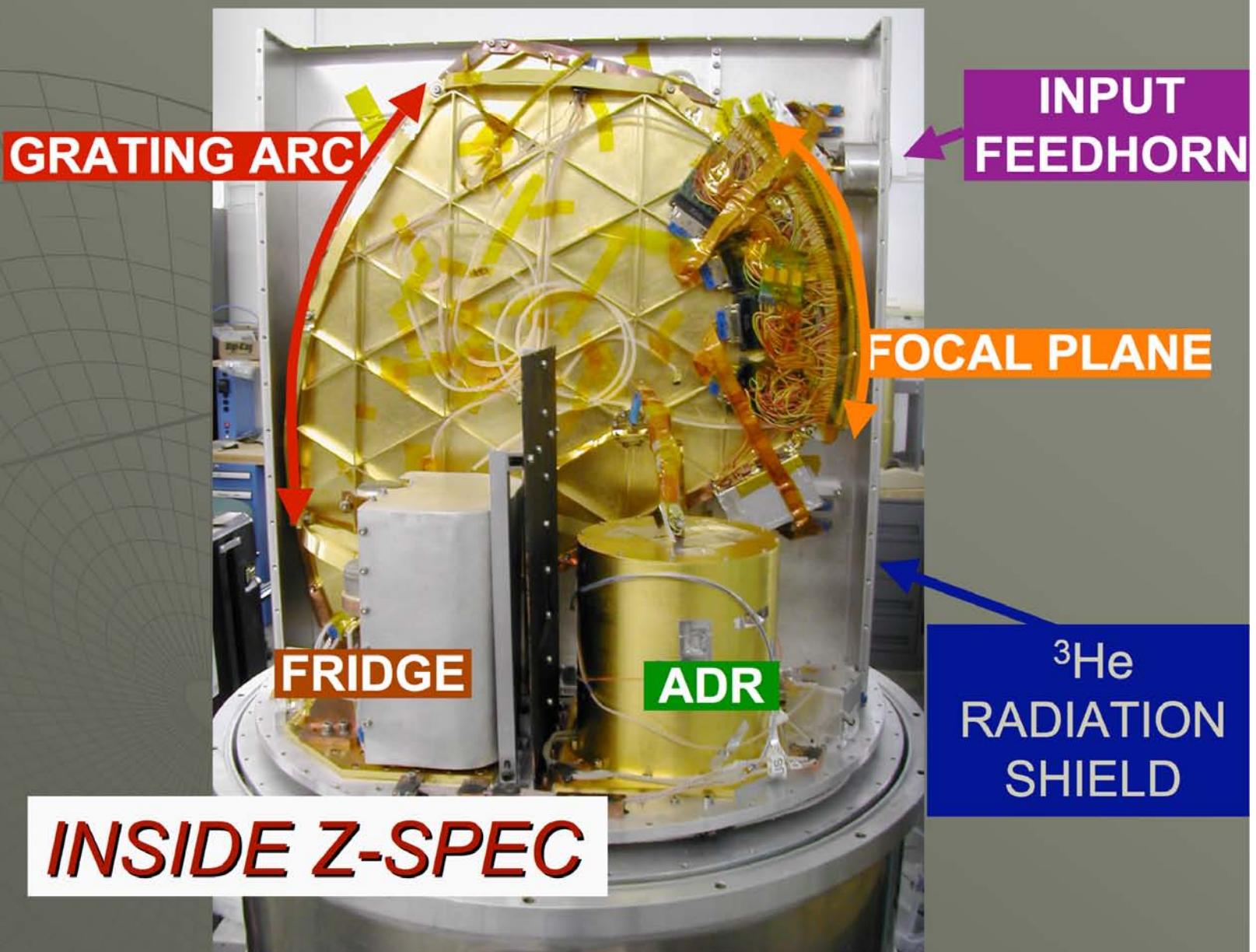
2.5 mm plate  
separation



480 facets

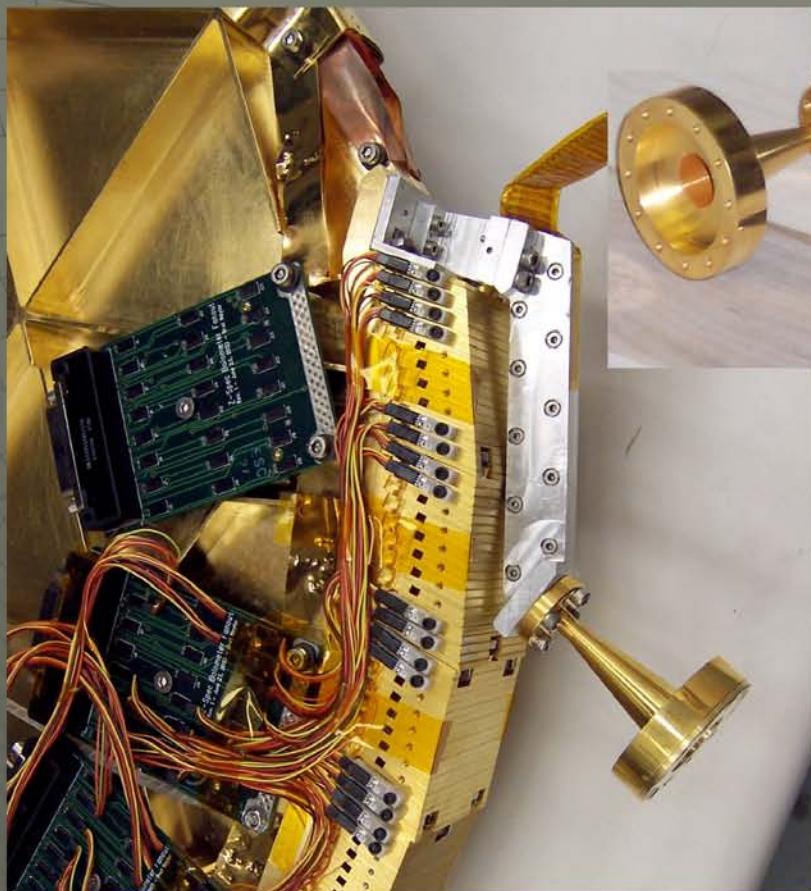
**Compact**  
**(62 cm x 48 cm x 3.3  
cm)**

# D E S I G N

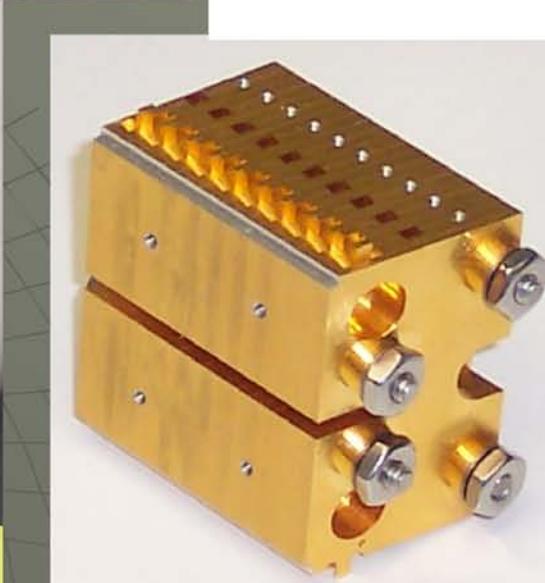


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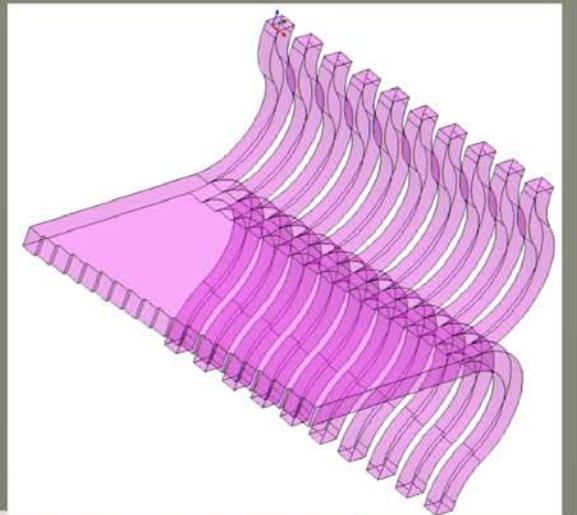
## Coupling Structures



**INPUT:**  
*Corrugated Feedhorn +  
Rectangular Waveguide*



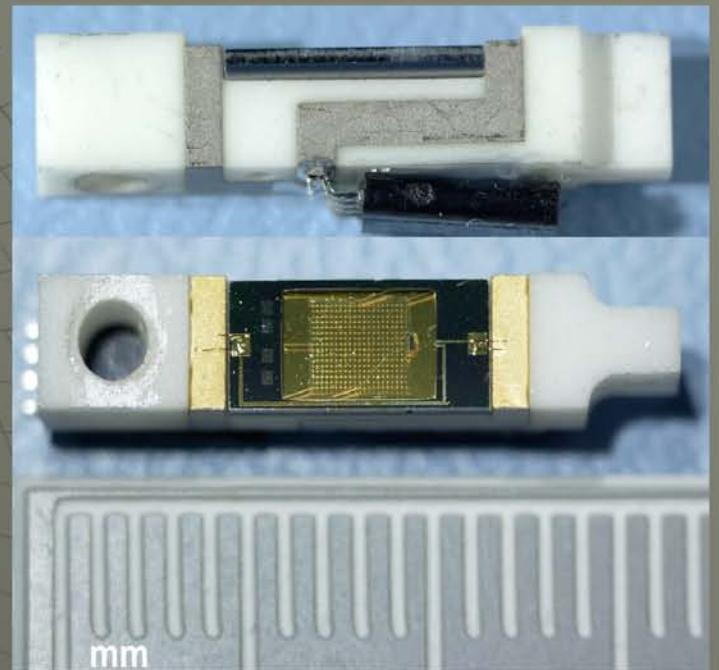
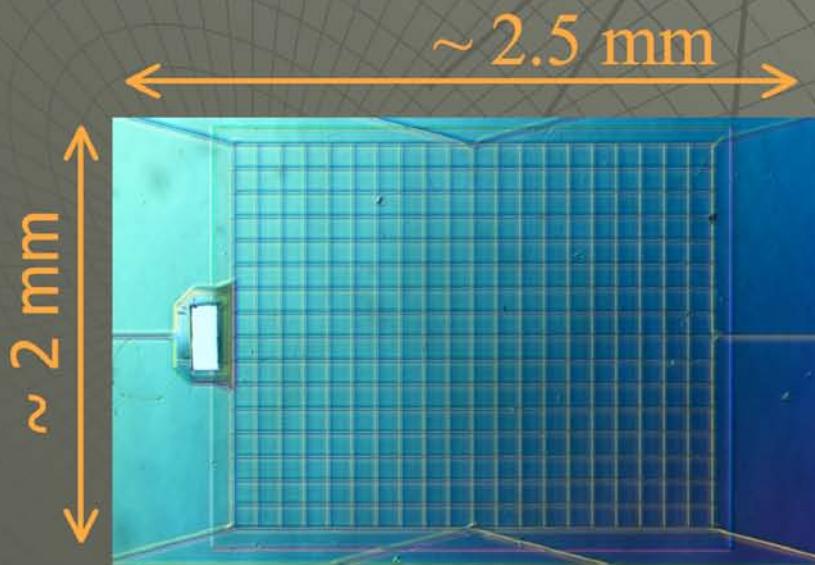
**OUTPUT:**  
*Waveguide Bendblocks*



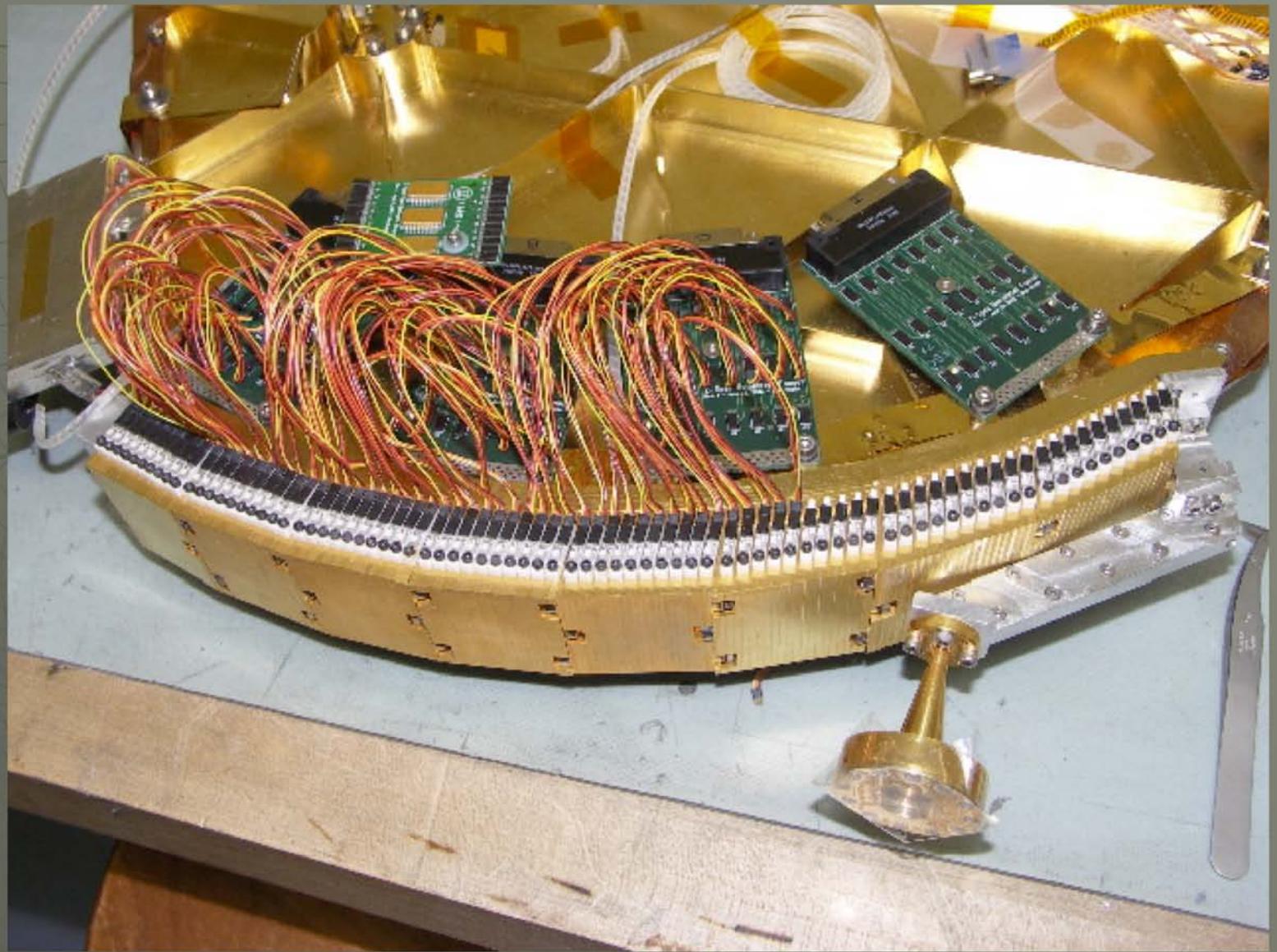
# D E S I G N

## Detectors

- 160  $\text{Si}_3\text{N}_4$  micromesh bolometers with NTD germanium thermistors at < 100 mK
- Bolometers individually mounted
- Design NEP  $\sim 4 \times 10^{-18} \text{ W Hz}^{-1/2}$



# D E S I G N



CCAT 05/08 - L. Earle

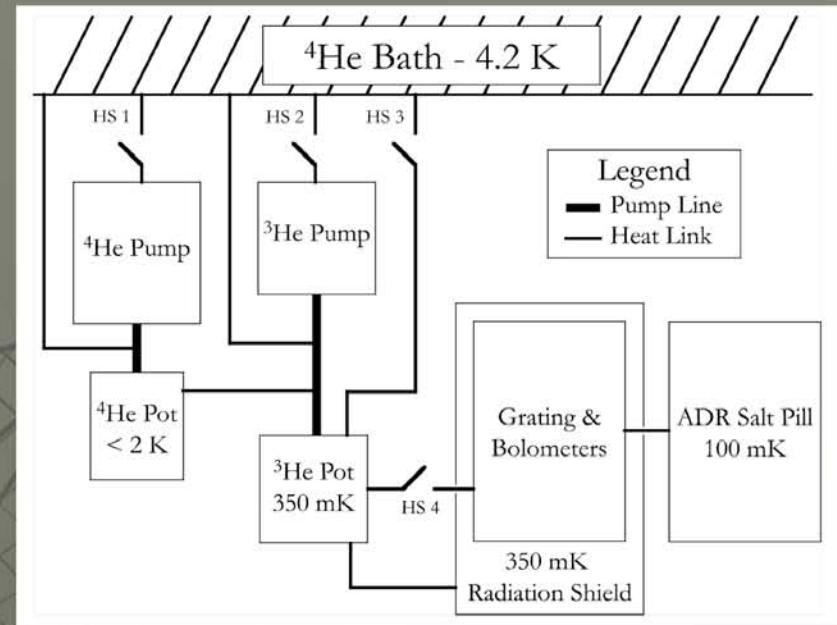
# D E S I G N

## Cryogenics



**3He/4He Fridge**

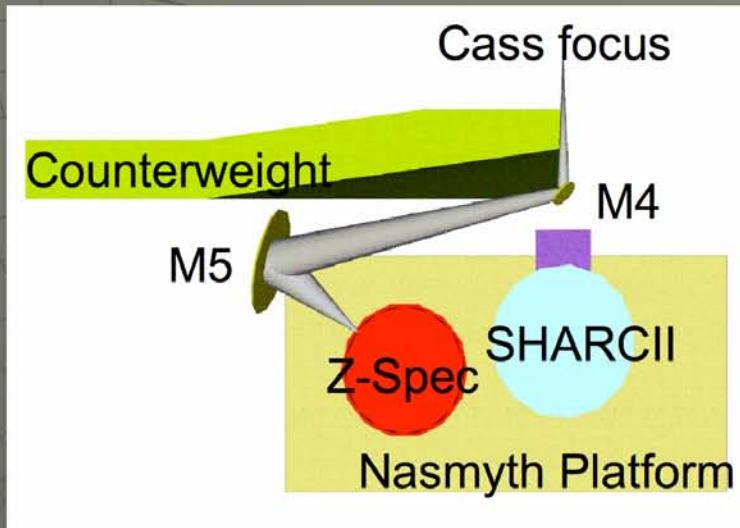
- Entire grating + detectors to 60 – 90 mK
- 2-stage cooling is fully automated
- No pumping on  $^4\text{He}$  bath
- ADR Hold time 16 ~ 20 hours
- $^4\text{He}$  fill every 48 hours



**ADR**

# D E S I G N

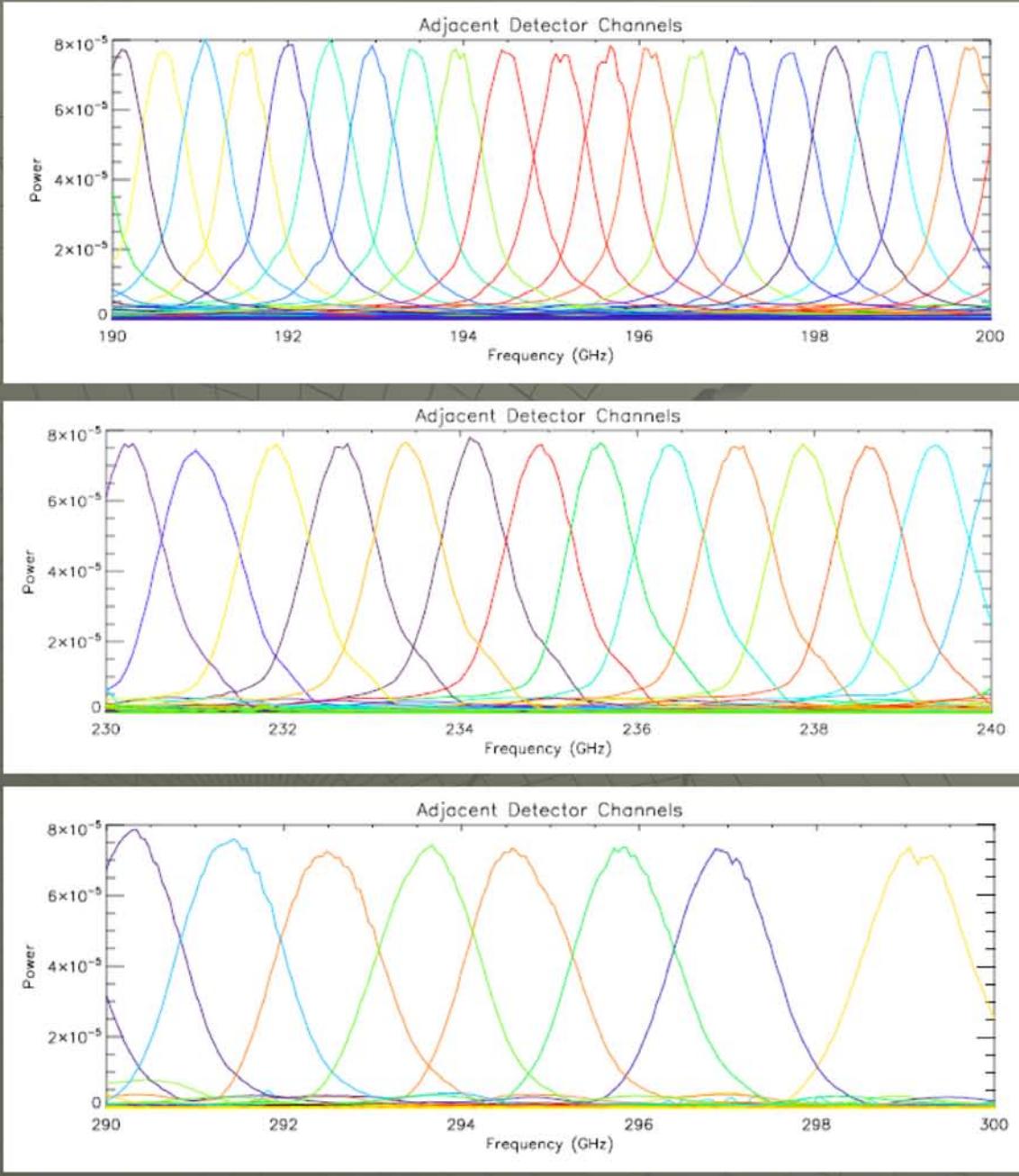
## At the CSO – Right Nasmyth port



ellipsoidal M5 converts  
f/12 from telescope to f/3  
spillover efficiency limited  
by secondary taper  
(89 ~ 91% theoretical)



# PERFORMANCE



CCAT 05/08 - L. Earle

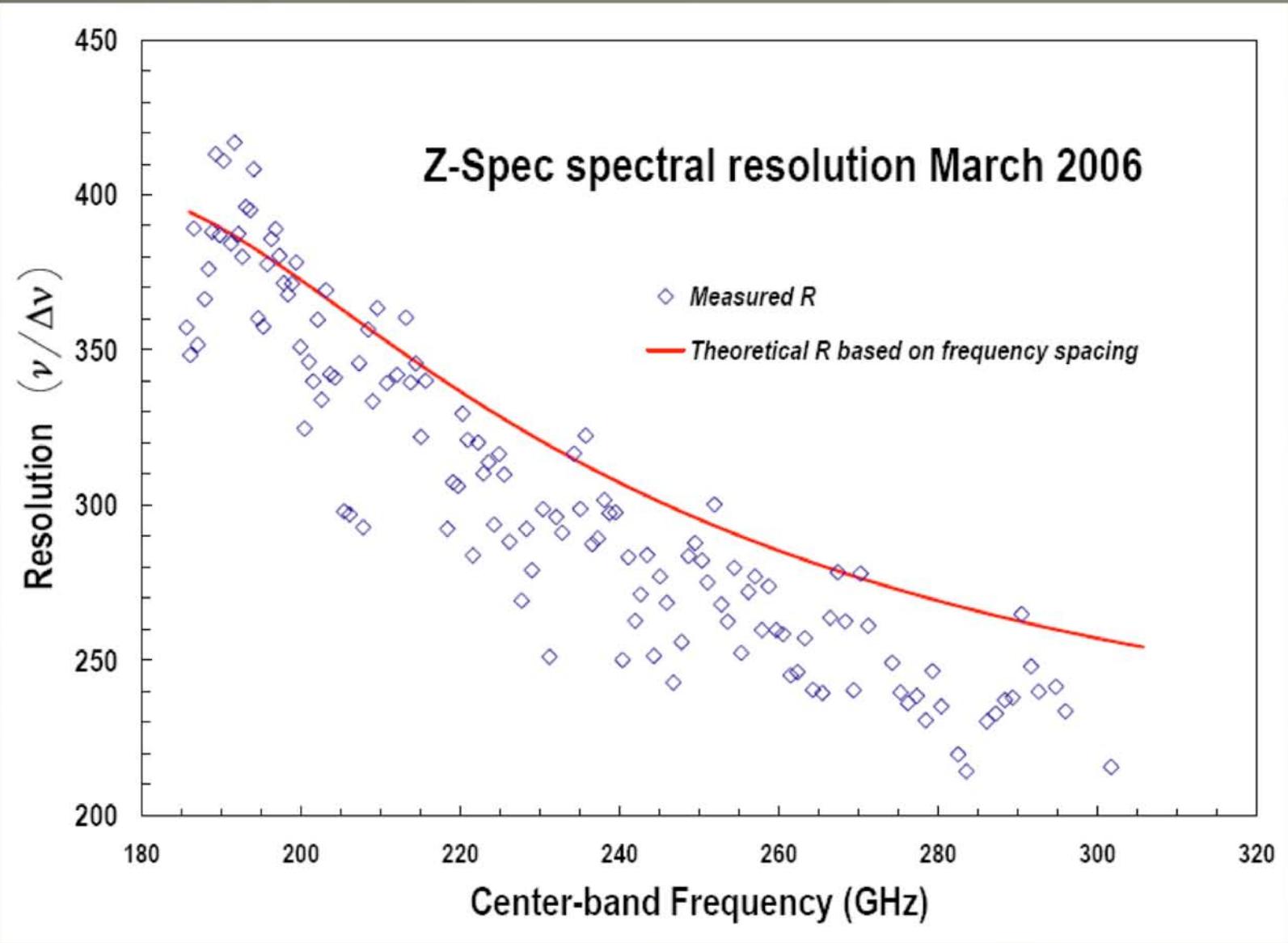
Spectral Response measured with FTS

Thorough coverage, not oversampled

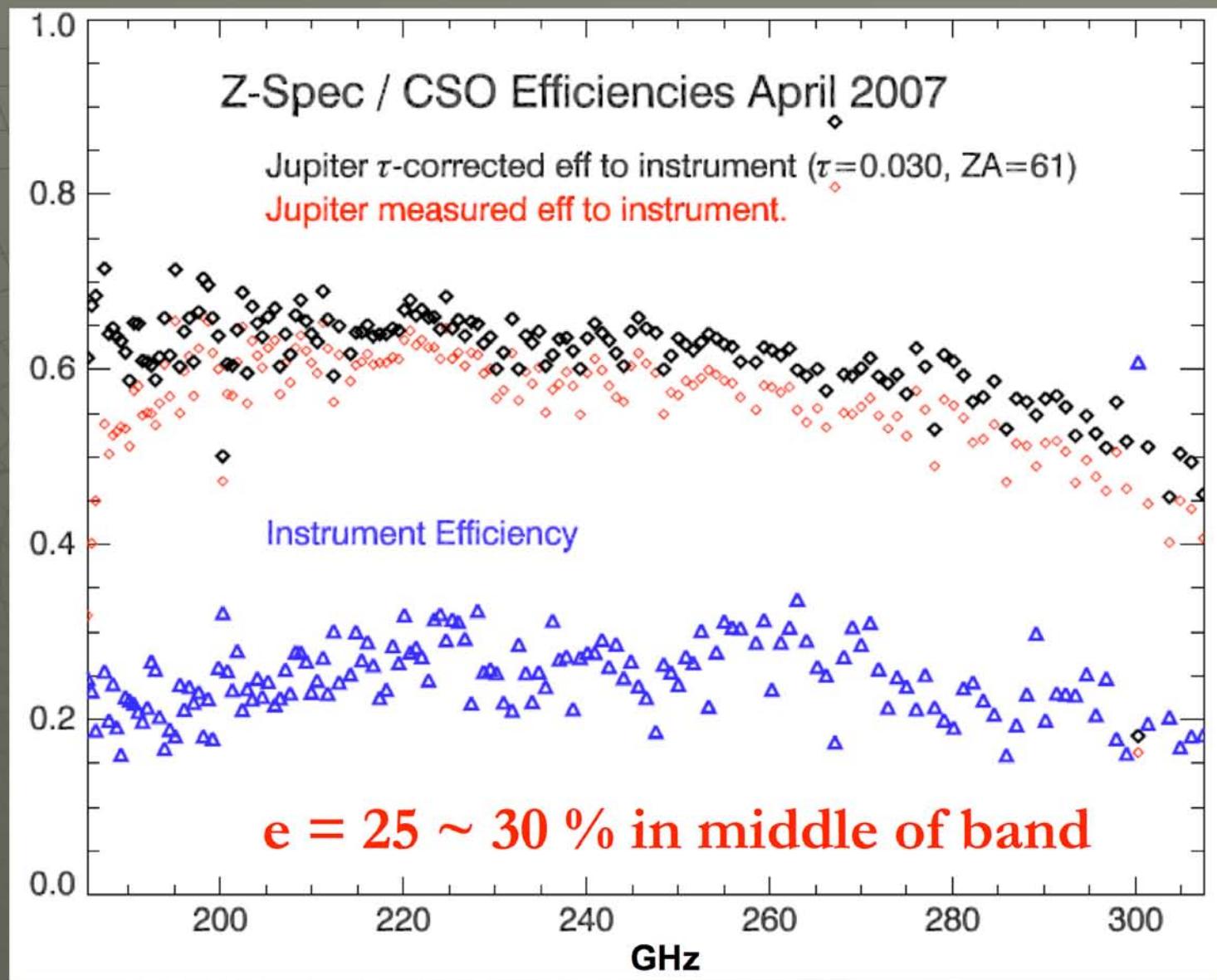
185 – 305 GHz  
(FTS resolution ~ 100 MHz)

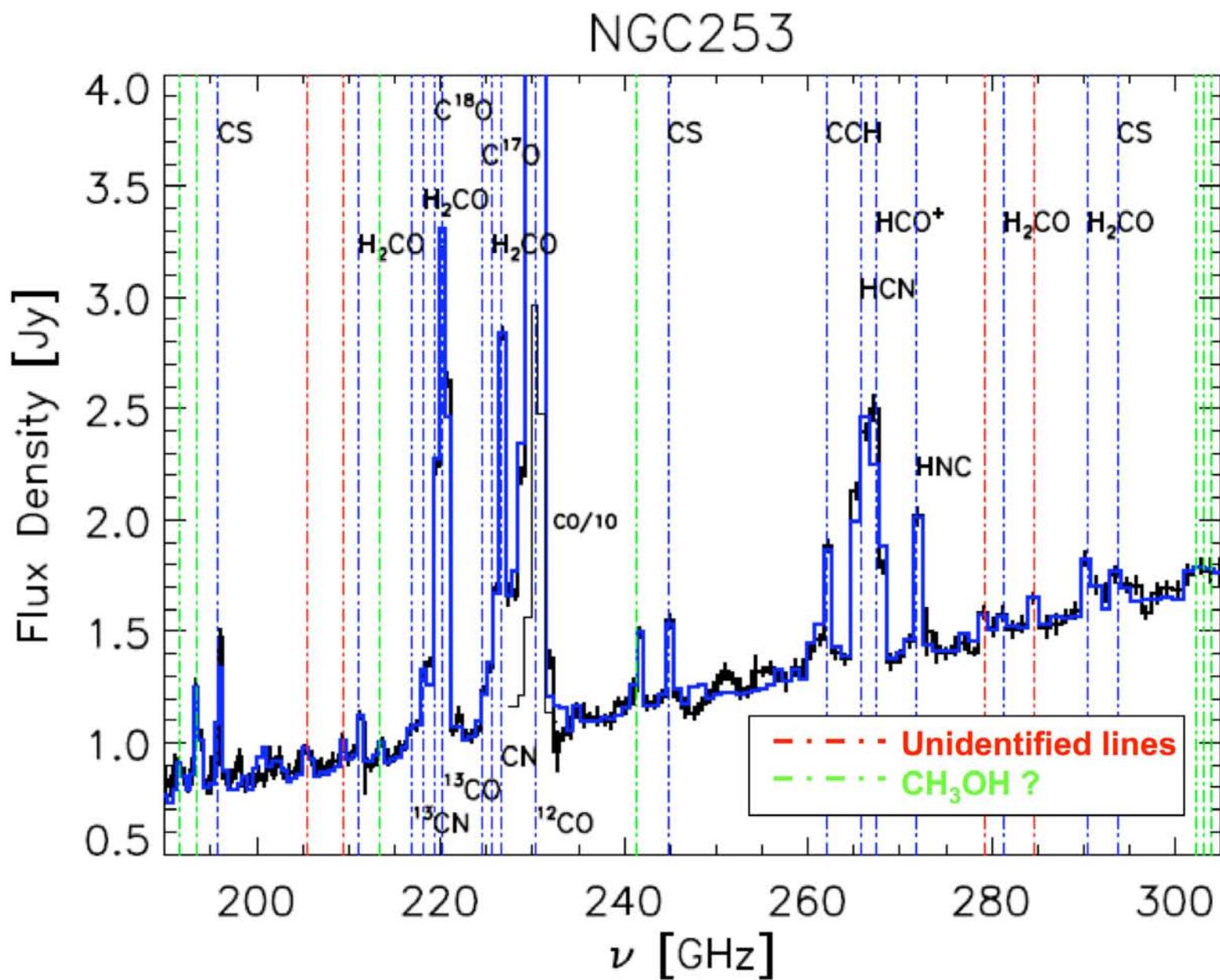
Peak height normalized to one

# PERFORMANCE



# PERFORMANCE





# PERFORMANCE

Spectral Lines in NGC253					
Species	Transition	Freq [GHz]	Flux [Jy km s <sup>-1</sup> ]	Flux [K km s <sup>-1</sup> ]	SNR
CO	J=2→1	230.35	41645.4	1281.4	158.9
<sup>13</sup> CO	J=2→1	220.22	4080.3	125.5	68.8
C <sup>18</sup> O	J=2→1	219.38	1055.0	32.5	22.0
C <sup>17</sup> O	J=2→1	224.53	58.2	1.8	1.9
CN	J=2→1	226.66	2418.8	74.4	60.2
<sup>13</sup> CN	J=2→1	216.87	191.3	5.9	3.8
CCH	J=3→2	262.01	857.6	26.4	19.8
HCN	J=3→2	265.67	1992.1	61.3	28.2
HCO <sup>+</sup>	J=3→2	267.34	1987.9	61.2	22.8
HNC	J=3→2	271.76	875.2	26.9	15.6
CS	J=4→3	195.80	785.8	24.2	17.3
CS	J=5→4	244.74	548.1	16.9	11.5
CS	J=6→5	293.68	380.7	11.7	5.1
H <sub>2</sub> CO	J <sub>K<sub>a</sub>,K<sub>e</sub></sub> =3 <sub>1,3</sub> →2 <sub>1,2</sub>	211.04	349.0	10.7	10.6
H <sub>2</sub> CO	J <sub>K<sub>a</sub>,K<sub>e</sub></sub> =3 <sub>0,3</sub> →2 <sub>0,2</sub>	218.05	434.7	13.4	10.0
H <sub>2</sub> CO	J <sub>K<sub>a</sub>,K<sub>e</sub></sub> =3 <sub>1,2</sub> →2 <sub>1,1</sub>	225.52	325.2	10.0	8.0
H <sub>2</sub> CO	J <sub>K<sub>a</sub>,K<sub>e</sub></sub> =4 <sub>1,4</sub> →3 <sub>1,3</sub>	281.30	189.3	5.8	2.9
H <sub>2</sub> CO	J <sub>K<sub>a</sub>,K<sub>e</sub></sub> =4 <sub>0,4</sub> →3 <sub>0,3</sub>	290.39	599.3	18.4	7.9

3.5 hrs at the  
CSO in  
November 07

Chopping at  
1.4 – 1.6 Hz  
with 90"  
throw

$\tau_{225 \text{ GHz}}$   
range .1 ~ .2

# PERFORMANCE

