



CCAT

Telescope Interface

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Outline

- Introduction
- Re-Imaging Optics at the Nasmyth Foci
- Re-Imaging Optics: Common Optics
- Instrument Layout
- Conclusions



CCAT Telescope Interface Instruments

- Location
 - Nasmith Optics
 - Re-imaging Optics: Common Optics ?
 - Bent Cassegrain Optics
- Metrology Interface
- Practical Aspects
 - Power consumption, weight, volume, cryogenics, telemetry/ control
 - Telescope restrictions: max load



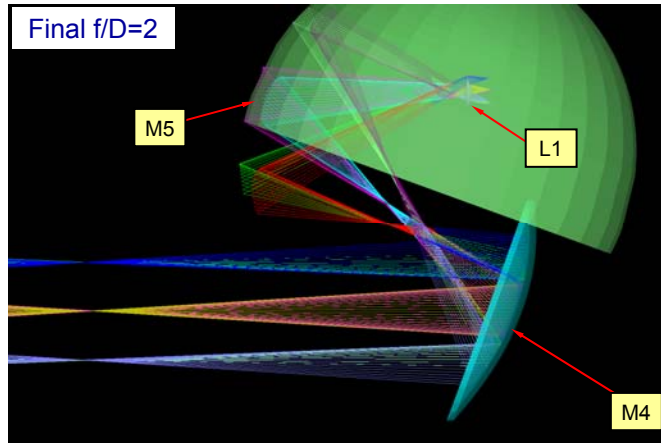
Re-Imaging Optics at the Nasmith Foci

LW-Cam Re-imaging Optics



LW-Cam Re-imaging Optics

for 20' Diameter FOV, $\theta_{inc}=20^\circ$, $f/0.6$

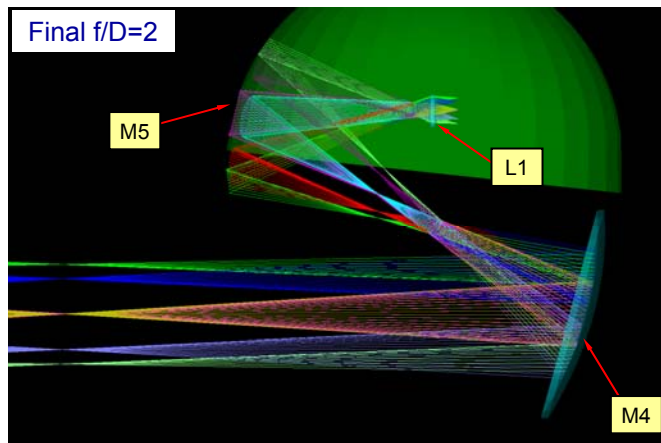


LWCam $f/0.6$, $\theta_{inc}=20^\circ$, $f/2$



LW-Cam Re-imaging Optics

for 20' Diameter FOV, $\theta_{inc}=15^\circ$, $f/0.6$

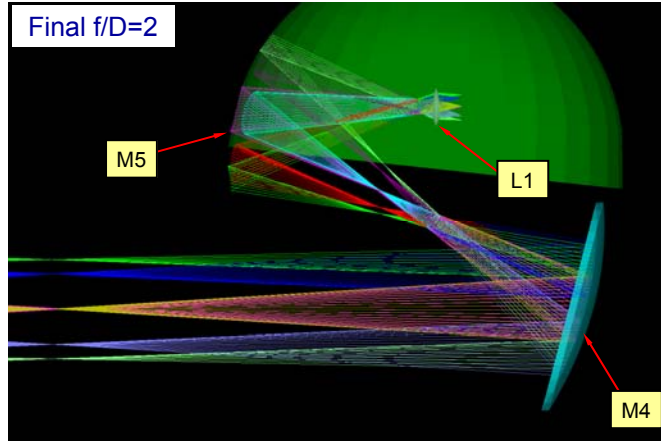


LWCam $f/0.6$, $\theta_{inc}=15^\circ$, $f/2$



LW-Cam Re-imaging Optics

for 20' Diameter FOV , $\theta_{inc}=15^\circ$, $f_1/0.4$

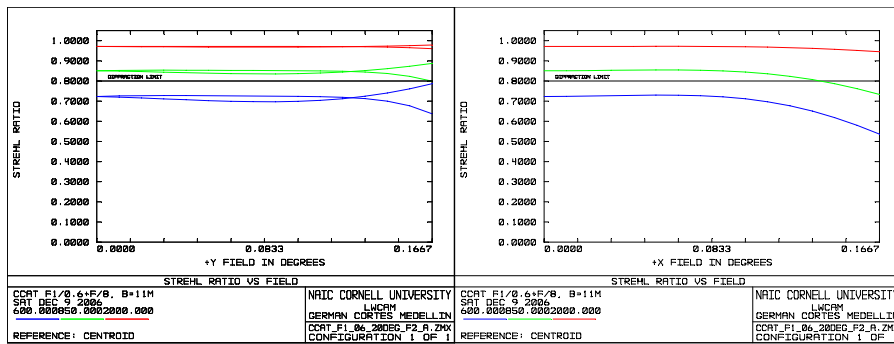


LWCam $f_1/0.4$, $\theta_{inc}=15^\circ$, f/2



LW-Cam Strehl Ratio over FOV

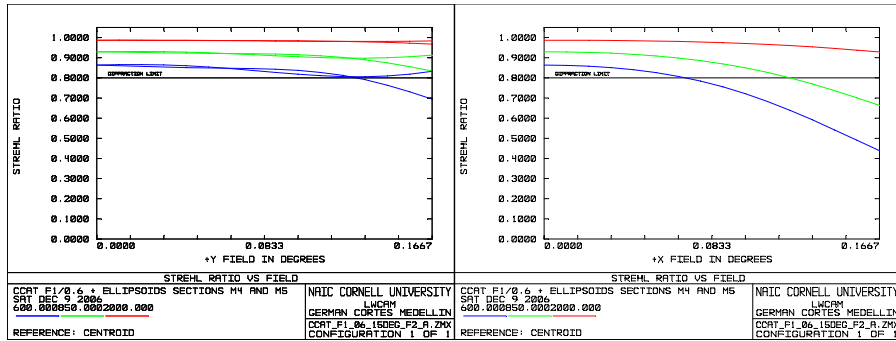
$\theta_{inc} = 20^\circ$, $f_1/0.6$





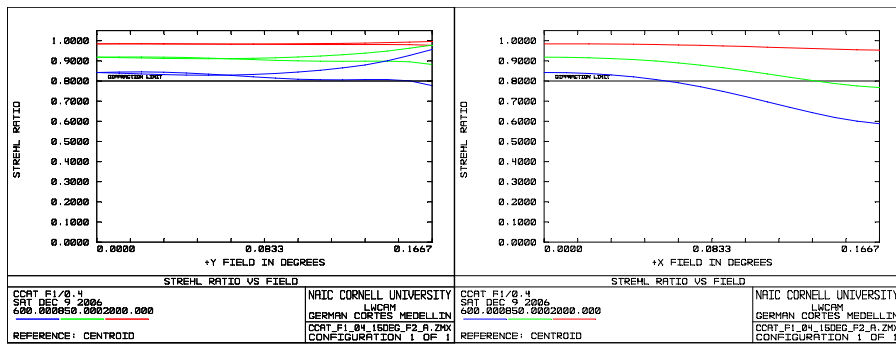
LW-Cam Strehl Ratio over FOV

$$\theta_{inc} = 15^\circ, f_1/0.6$$



LW-Cam Strehl Ratio over FOV

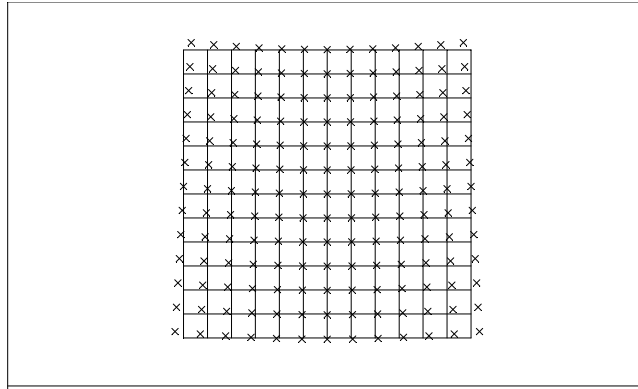
$$\theta_{inc} = 15^\circ, f_1/0.4$$





LW-Cam Distortion over FOV

$$\theta_{inc} = 15^\circ, f_1/0.6$$

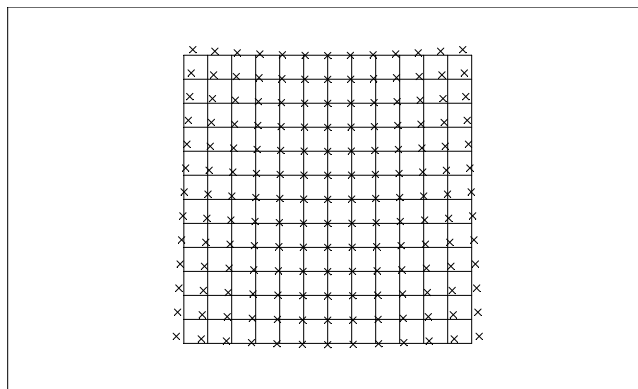


GRID DISTORTION	
CCAT F1/0.6 + ELLIPSOIDS SECTIONS M4 AND M5	NAIC CORNELL UNIVERSITY
SAT DEC 9 2006	LWCAM
FIELD: 0.2357 W 0.2357 H DEGREES	GERMAN CORTES MEDELLIN
IMAGE: 206.54 W 206.63 H MILLIMETERS	CCAT_F1_06_15DEG_F2_A.ZMX
MAXIMUM DISTORTION: 5.2222%	CONFIGURATION 1 OF 1
SCALE: 1.000X, WAVELENGTH: 600.0000 μm	



LW-Cam Distortion over FOV

$$\theta_{inc} = 15^\circ, f_1/0.4$$

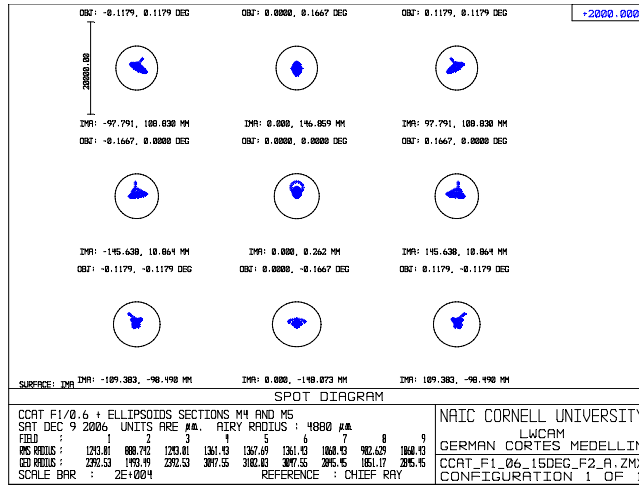


GRID DISTORTION	
CCAT F1/0.4	NAIC CORNELL UNIVERSITY
SAT DEC 9 2006	LWCAM
FIELD: 0.2357 W 0.2357 H DEGREES	GERMAN CORTES MEDELLIN
IMAGE: 205.44 W 205.43 H MILLIMETERS	CCAT_F1_04_15DEG_F2_A.ZMX
MAXIMUM DISTORTION: -5.2141%	CONFIGURATION 1 OF 1
SCALE: 1.000X, WAVELENGTH: 600.0000 μm	



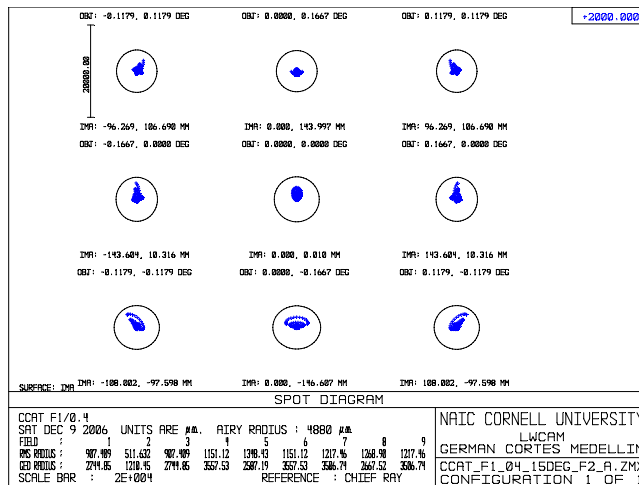
LW-Cam Spot Diagram

$\theta_{inc} = 15^\circ, f_1/0.6$



LW-Cam Spot Diagram

$\theta_{inc} = 15^\circ, f_1/0.4$





Common Optics SW-Cam and LW-Cam



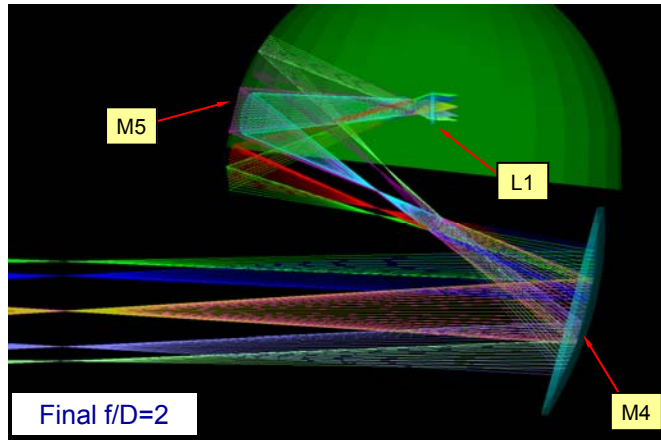
Common Optics SW-Cam and LW-Cam

- Only LW-Cam M4 same, or
- Both LW-Cam Re-imaging M4 and M5
- +Different Lens System
- Optimized for SW-Cam f/4.8 for 20' \varnothing FOV



LW-Cam Optics

for 20' Diameter FOV, $\theta_{inc}=15^\circ$, $f/0.6$

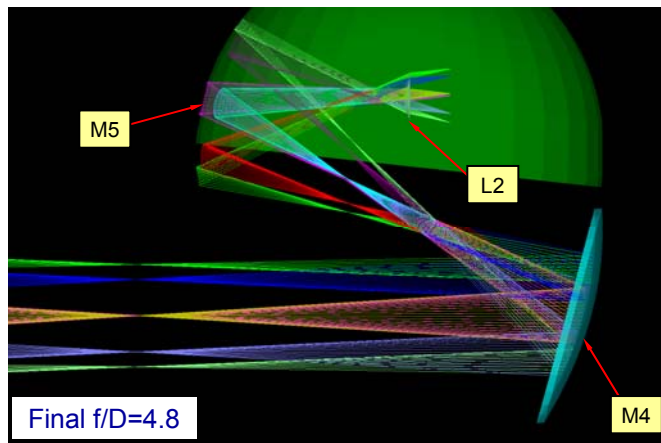


LWCam $f/0.6$, $\theta_{inc}=15^\circ$, $f/2$



SW-Cam Optics

for 20' Diameter FOV, $\theta_{inc}=15^\circ$, $f/0.6$



LWCam $f/0.6$, $\theta_{inc}=15^\circ$, $f/4.8$



Common Optics SW-Cam and LW-Cam

- Only LW-Cam M4 same, or
- Both LW-Cam Re-imaging M4 and M5
- +Different Lens System
- Optimized for SW-Cam $f/4.8$ for $20' \varnothing$ FOV
- Initial Results: Promising
- Possible improvements:
 - More degrees of freedom reflectors
 - Additional lens
 - Further reduction of θ_{inc} to 12° or 11°
 - Reduced FOV

Work in Progress...



Nasmith Instrument Layout

(First Round)



Considerations and Limitations

- Required FOV, wavelength and f/D
- Size
- Weight
- Associated Moment of Inertia
- Power
- Cryogenics
- Accessibility
- Telemetry/Control
- Others



CCAT Instrumentation

Instrument	Wavelength μm	FOV arc-min	Final f/D
SW Cam	650 – 200	5' x 5'	4.8
SCUBA-2	850 – 450	3'(5') x 3'(5')	2.4?
LW Cam	2000 – 740	20' x 20'	2
Heterodyne Camera	450 – 350	1' x 1'	8
Spectrometer(Singl/Mult-Object)	450 – 350	2' x 2'	:
ZEUS (echelle spectrometer)	450 - 200	:	2.5
Z-Spec (wguide spectrometer)	1600 - 1000	:	:
Polarimeters	1000 - 200	5' x 5'	:
ALMA Instrumentation	:	:	8



CCAT Instrumentation

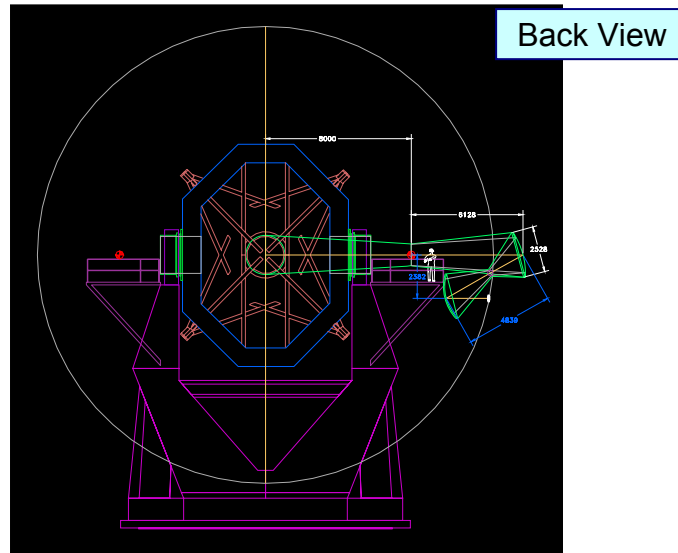
Instrument	Weight Tons	Vol m ³	Pow kW	Cryogen Operation Liters/day	Cryogen Pre-cool Liters
SW Cam	:	0.5	:	:	:
SCUBA-2	3.5	5.0	45	~5	600
LW Cam	1.0	:	:	:	:
Heterodyne Camera	0.9	1	20	:	:
Spectrometer(Singl/Mult-Object)	2	4.5	5-22	:	:
ZEUS (echelle spectrometer)	:	:	:	:	:
Z-Spec (wguide spectrometer)	:	:	:	:	:
Polarimeters	:	:	:	:	:
ALMA Instrumentation	:	:	:	:	:



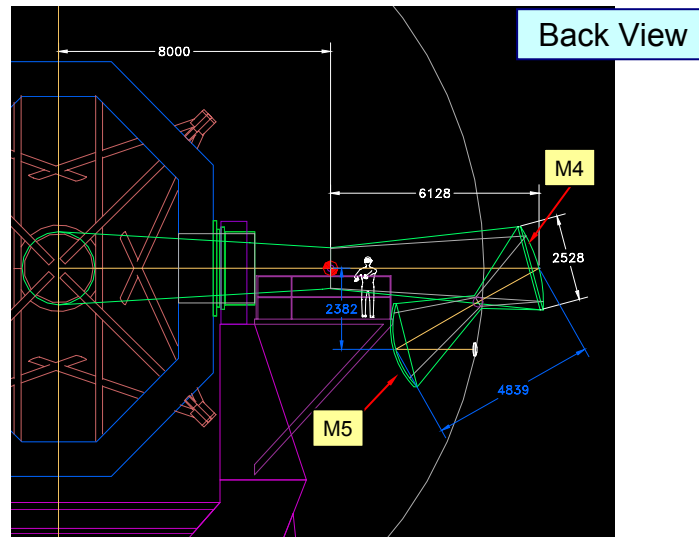
LW-Cam Location



LW-Cam Location I

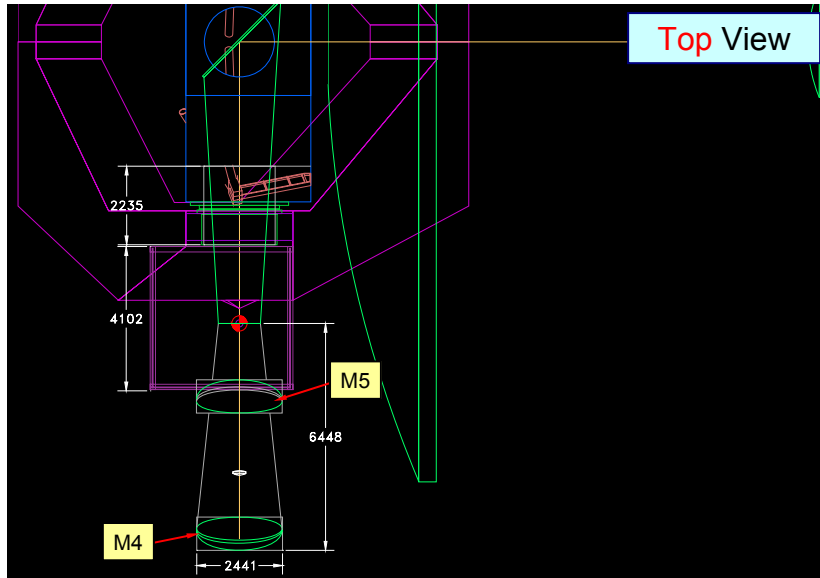


LW-Cam Location I

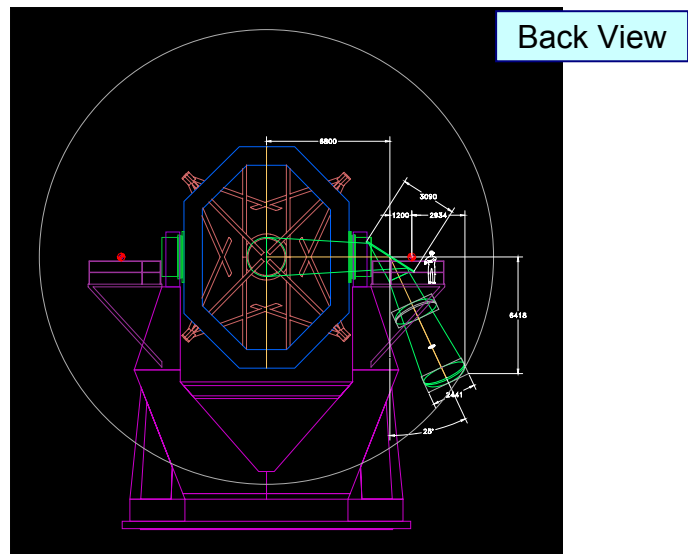




LW-Cam Location I

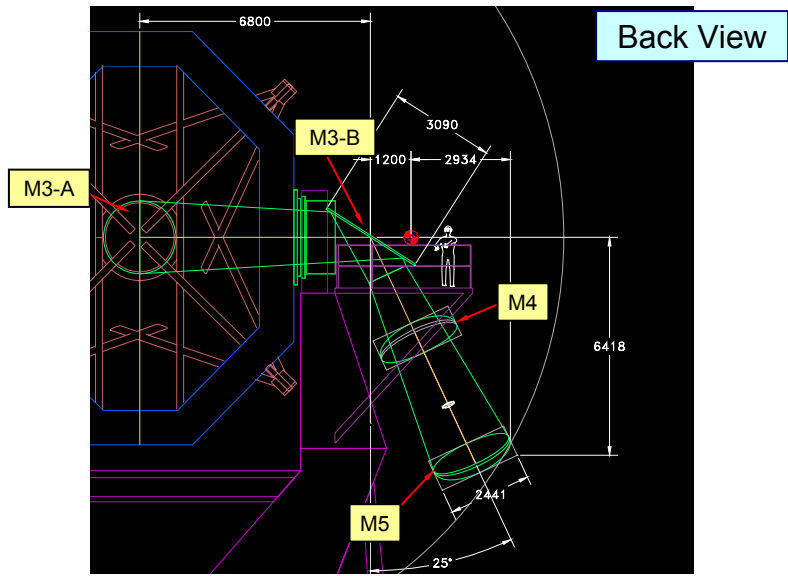


LW-Cam Location II

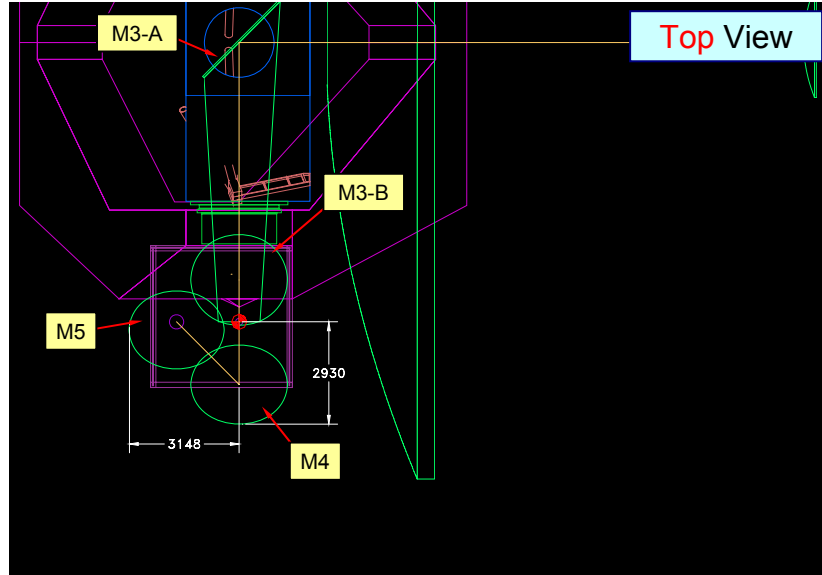




LW-Cam Location II

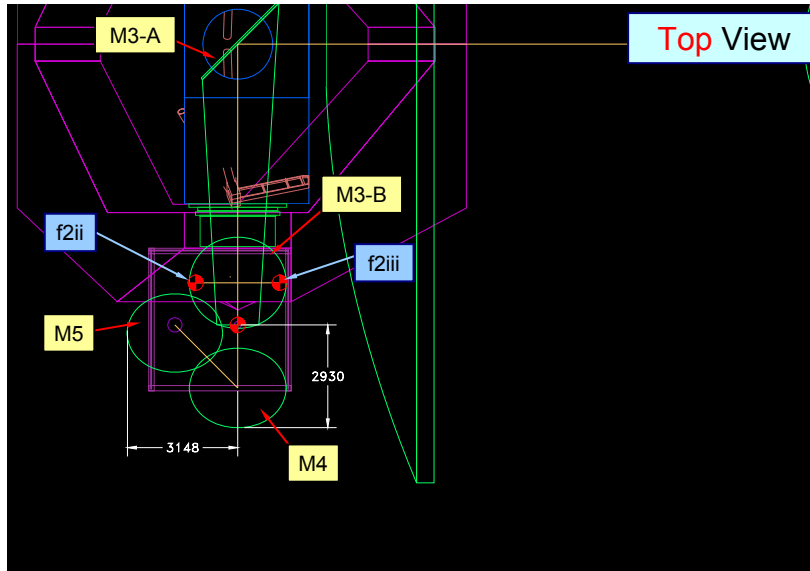


LW-Cam Location II

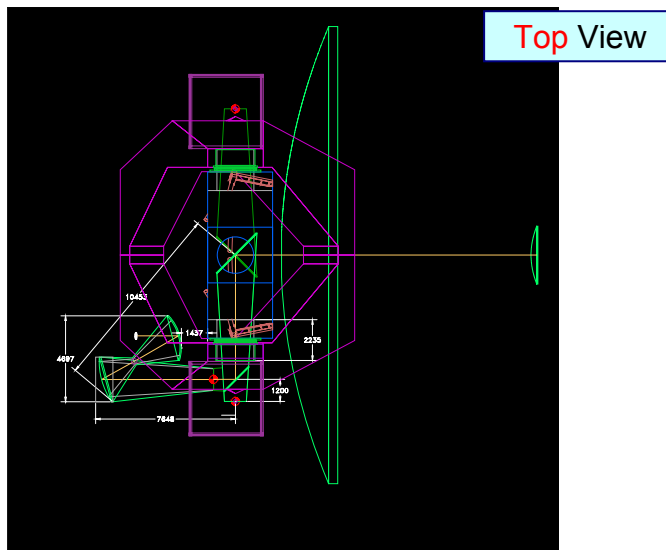




LW-Cam Location II

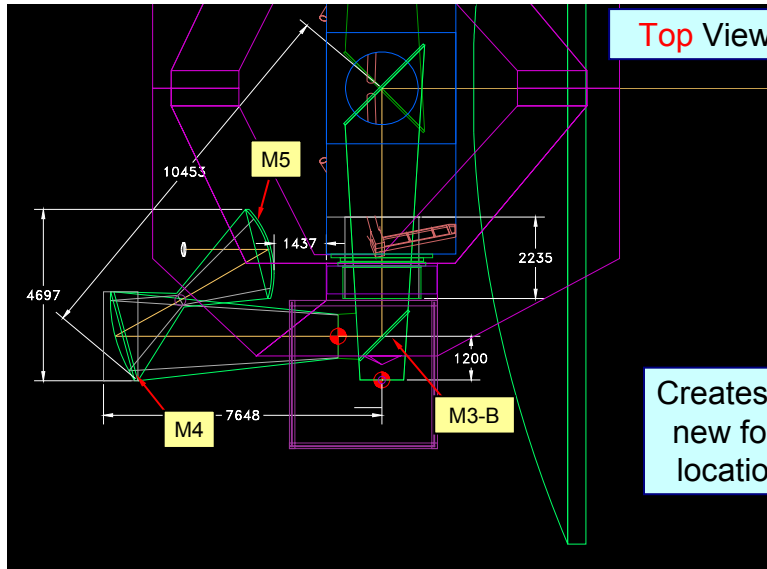


LW-Cam Location III

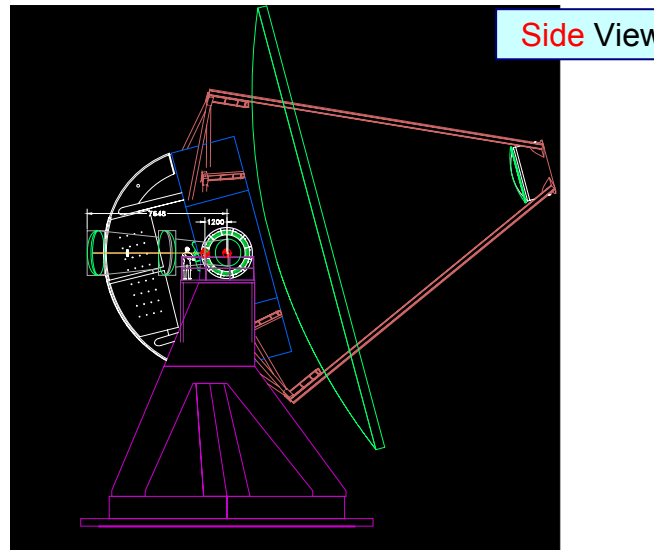




LW-Cam Location III

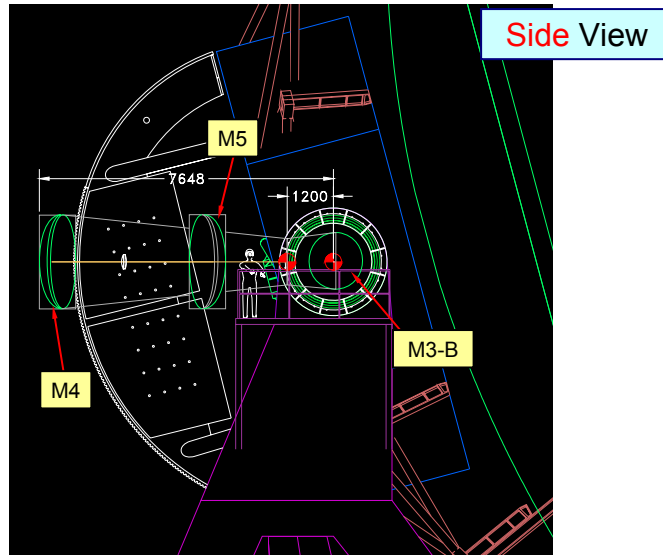


LW-Cam Location III

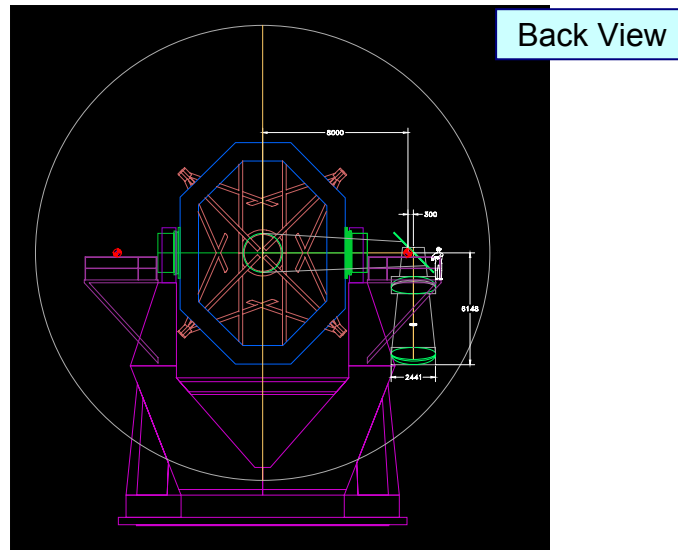




LW-Cam Location III

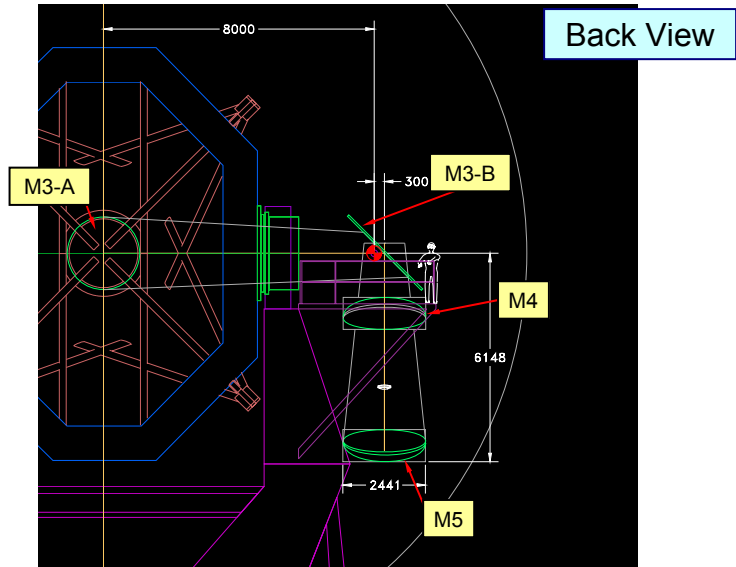


LW-Cam Location IV

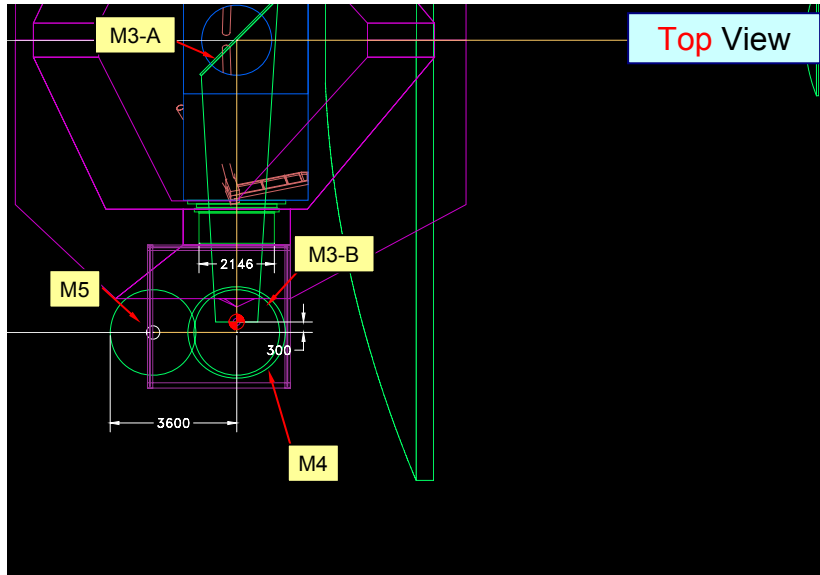




LW-Cam Location IV

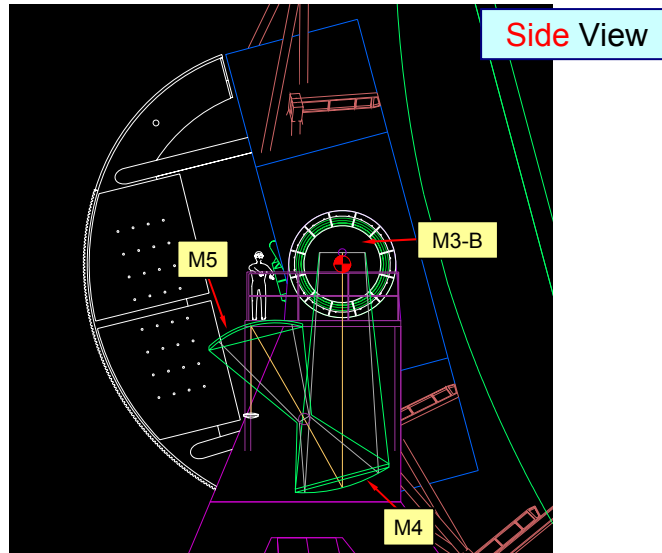


LW-Cam Location IV





LW-Cam Location IV



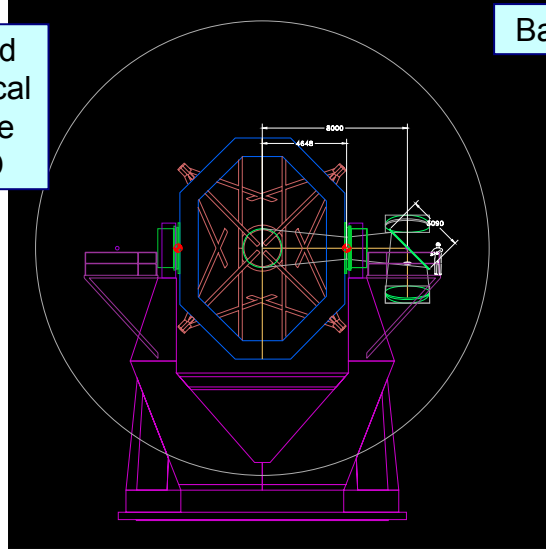
LW-Cam Other Possibility V

Reduced Back Focal
Distance B 11m→8m,
and f/D f/8→f/6.5



LW-Cam Other Possibility V

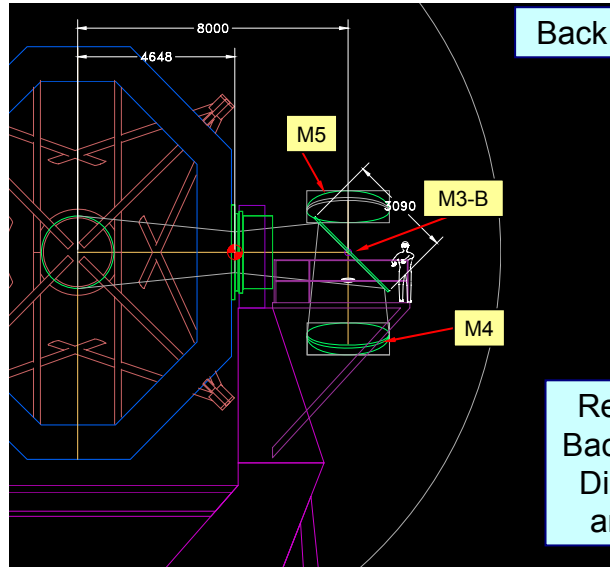
Reduced Back Focal Distance and f/D



Back View



LW-Cam Other Possibility V



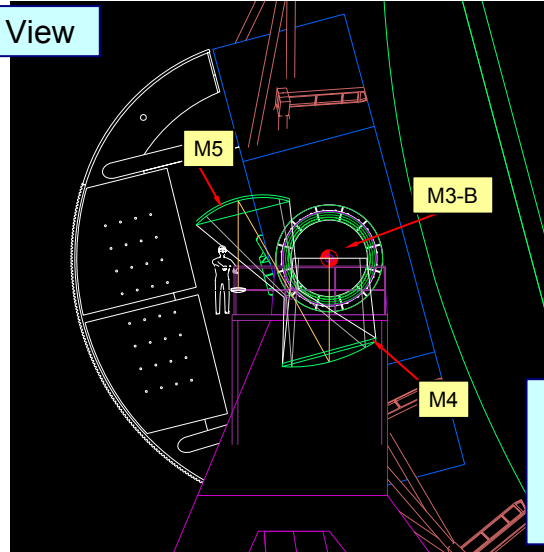
Back View

Reduced Back Focal Distance and f/D



LW-Cam Other Possibility V

Side View



Reduced
Back Focal
Distance
and f/D



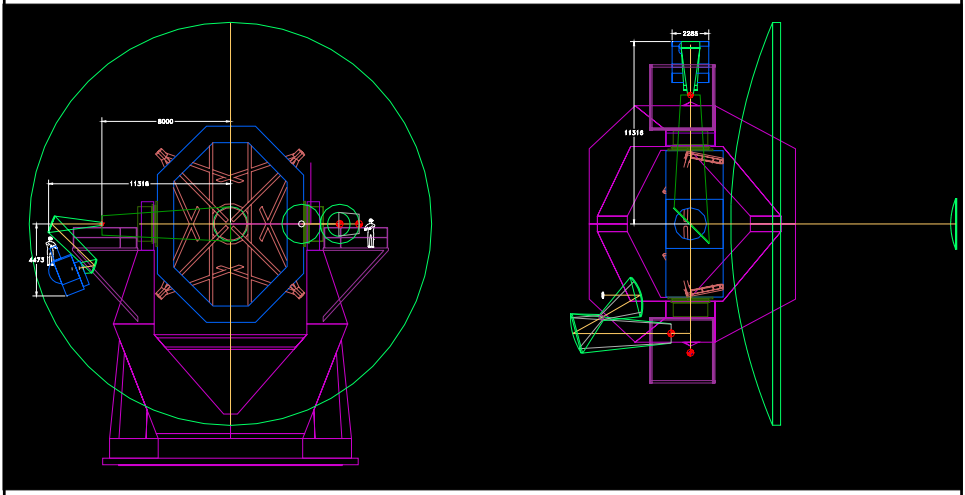
SCUBA-2 Location



SCUBA-2 Location I

Back View

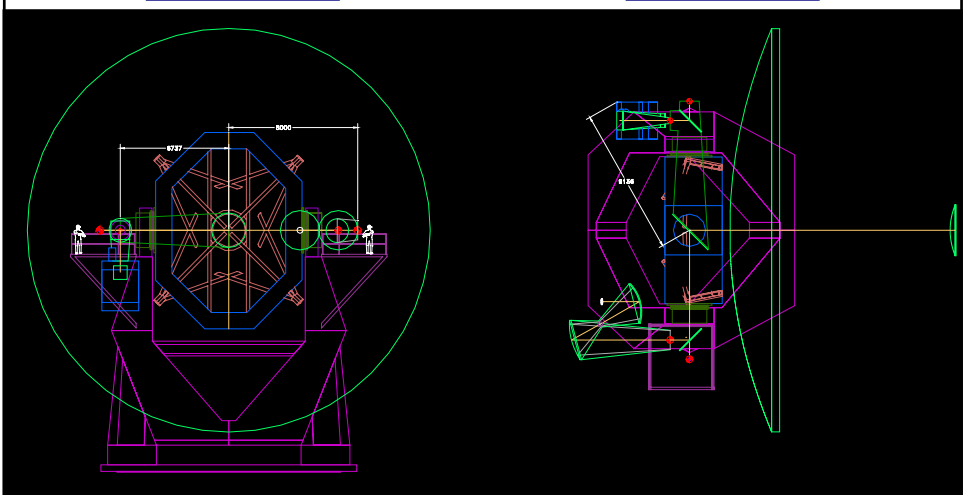
Top View



SCUBA-2 Location II

Back View

Top View

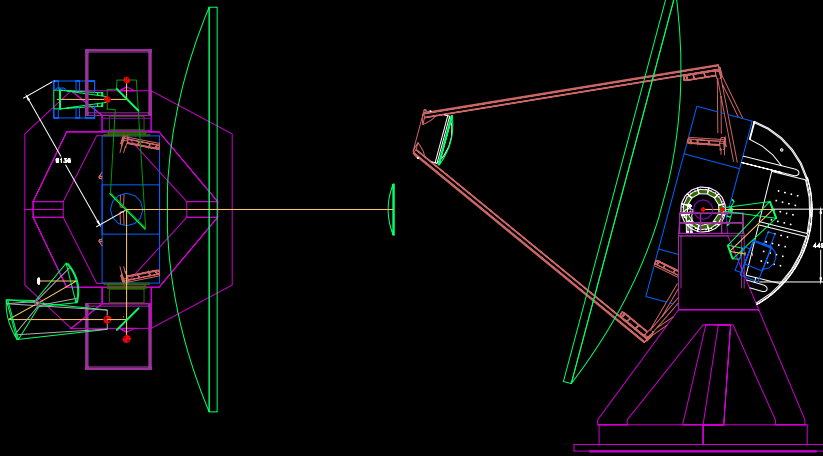




SCUBA-2 Location II

Top View

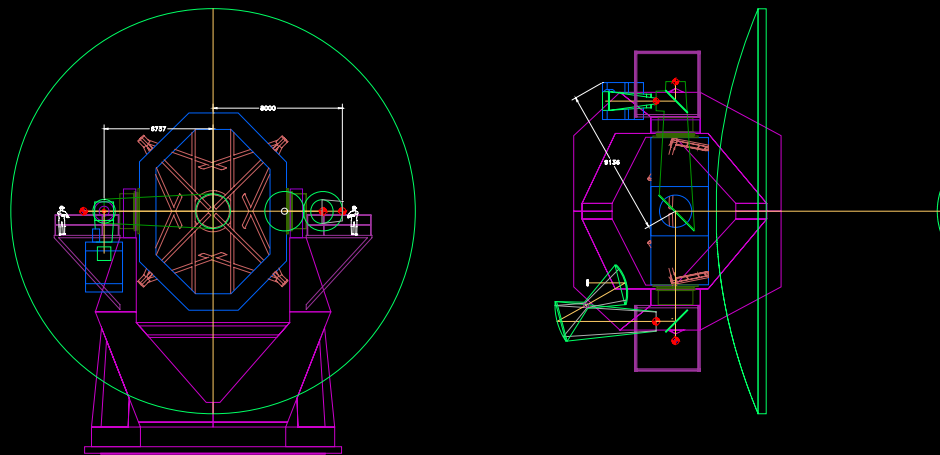
Side View



SW-Cam + SCUBA-2 Location I

Back View

Top View

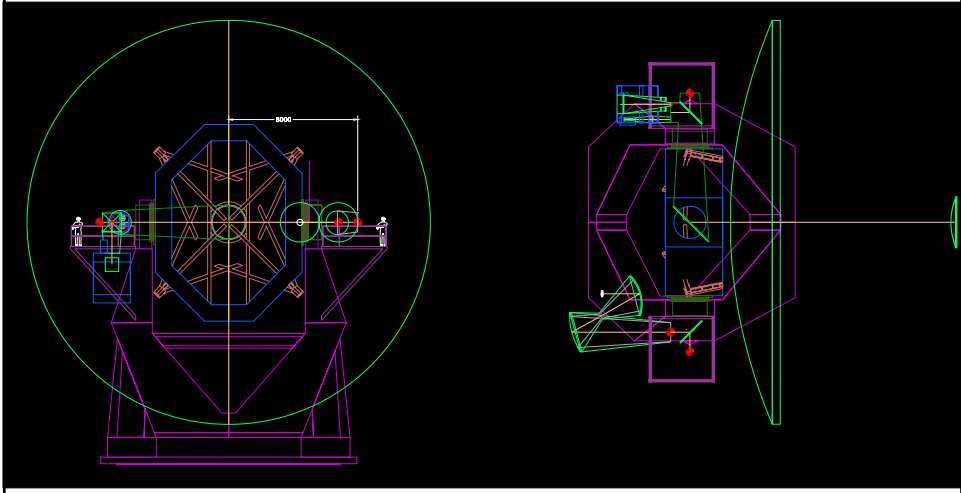




SW-Cam + SCUBA-2 Location I

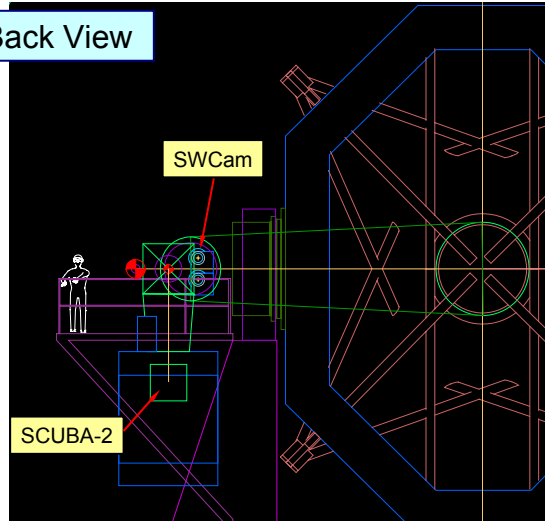
Back View

Top View



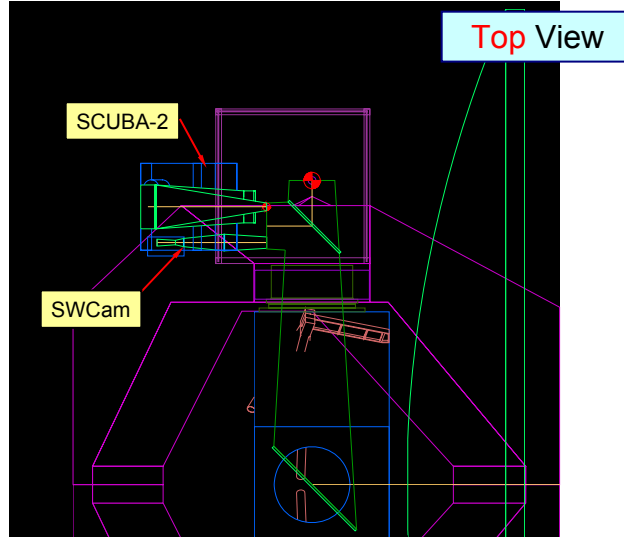
SW-Cam + SCUBA-2 Location I

Back View

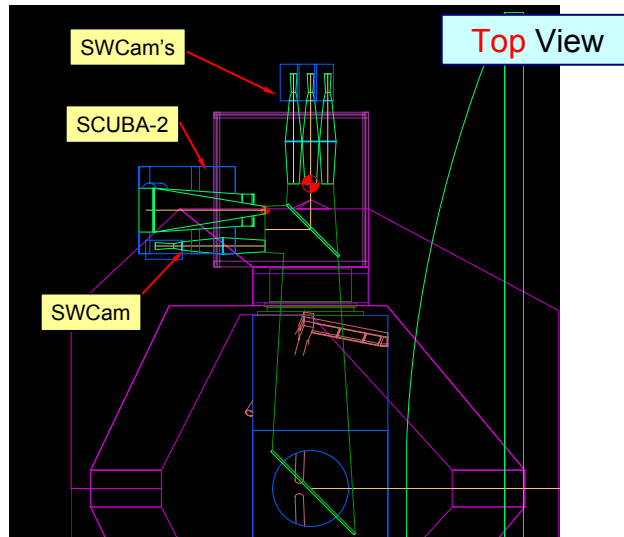


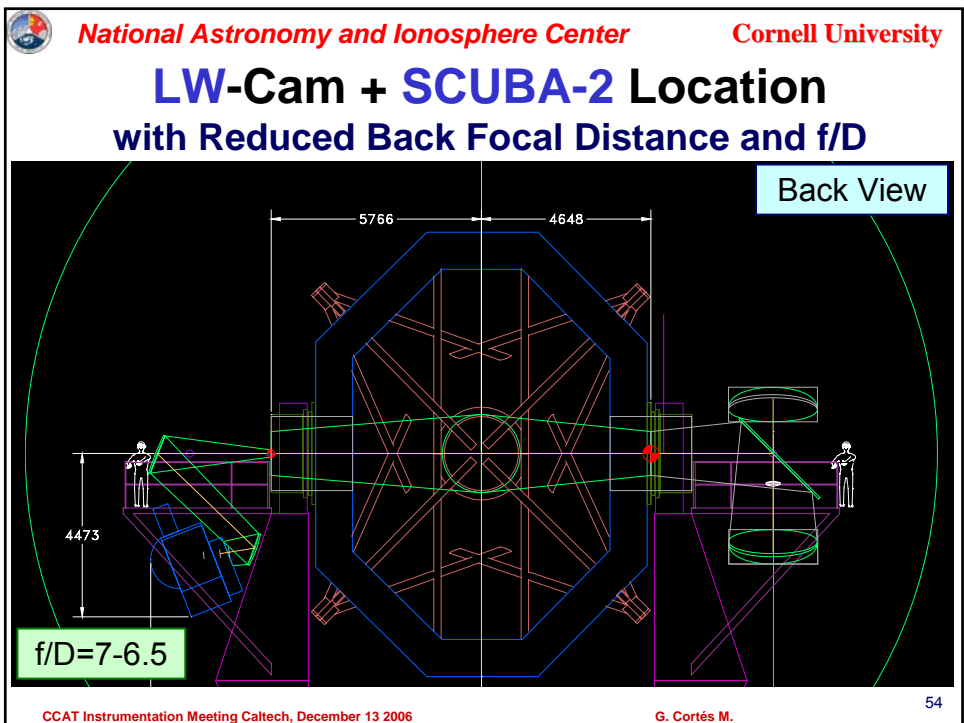
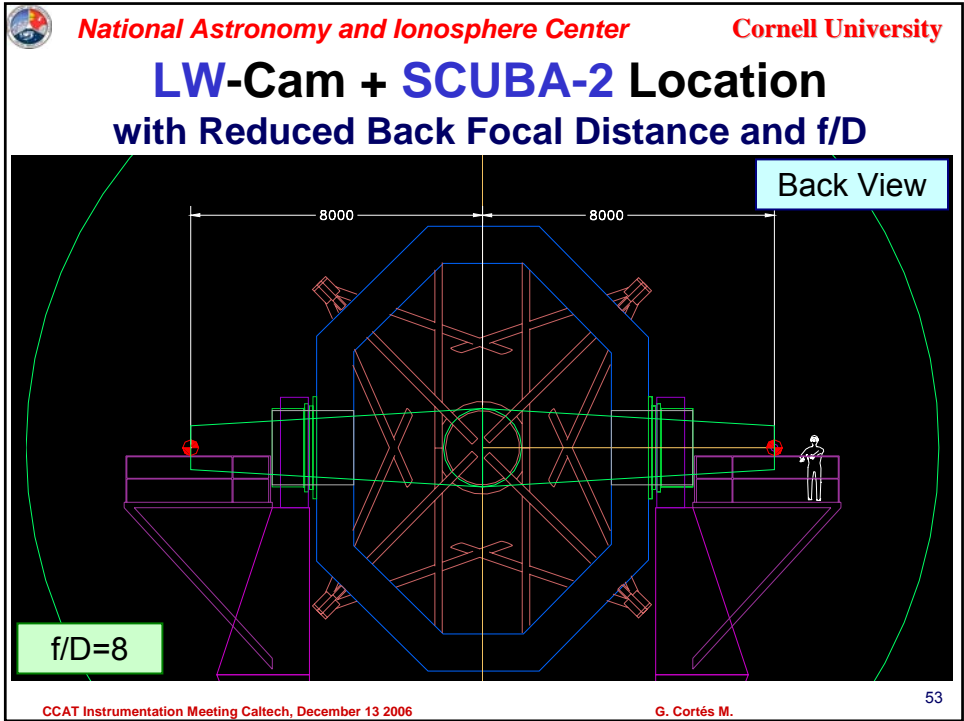


SW-Cam + SCUBA-2 Location I



SW-Cam + SCUBA-2 Location II







Reduced Back Focal Distance and f/D Advantages

- Better optical match for low f/D instruments
- Bent-Cassegrain becomes more compact
- Reduced curvature of Re-imaging reflectors will improve image performance (Strehl ratio, distortion, etc) over a large bandwidth.
- More compact Nasmith optics: improves antenna stiffness, moment of inertia, reduces load on the azimuth drivers

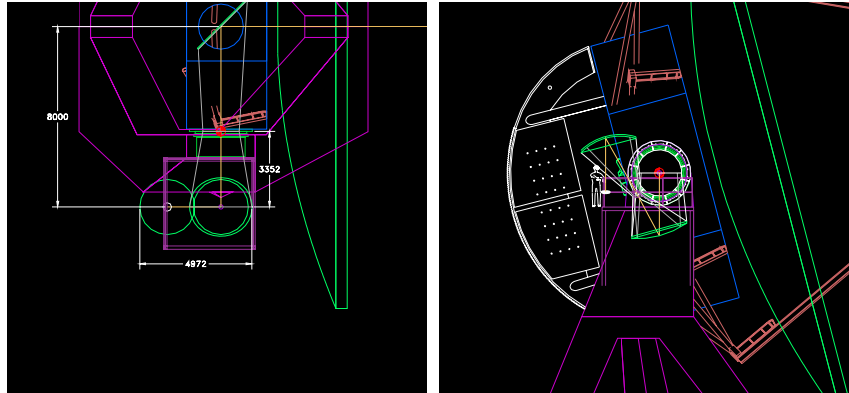


Reduced Back Focal Distance and f/D Possible Disadvantages

- May limit the number of big instruments.
- Although... Doesn't precludes to have additional secondary focal (F_{2i}) points at the Nasmith platforms as the f/8 system



Reduced Back Focal Distance and f/D Additional focal planes



Reduced Back Focal Distance and f/D Possible Disadvantages

- May limit the number of big instruments.
- Although, a f/8 system offers the possibility of direct access to multiple folded Nasmith focal points,
- Does not preclude the possibility of having additional indirect Nasmith focal points.
- Other possibilities include to position the instruments on a linear or circulating track to alternatively have access to the Nasmith focal region.



Conclusions

- Obtained an optical design for the re-imaging optics of the LW Cam based on off-axis ellipsoidal sections with a single lens system to provide the and $f/D=2$ at the image plane.
- The calculated Strehl ratio is better than 80% over a $20' \text{ } \varnothing$ FOV from $850 \text{ } \mu\text{m}$ to 2mm and reduced FOV performance down to $600 \text{ } \mu\text{m}$.
- The angle of incidence $\theta_{\text{inc}}=15^\circ$, works well at $f_1/0.6$ and $f_1/0.4$
- Fill distortion is about 5.2% for both $f_1/0.6$ and $f_1/0.4$
- The use of the same reflector system for both, LW- Cam and the SW-Cam is plausible, work is still in progress...



Conclusions Cont...

- We have an initial exploration of different possibilities of instrument locations at the Nasmith focal points based on the known dimensions of the instruments: the LW-Cam, the SW-Cam and SCUBA-2.
- A system $f/8$ offers the possibility of three or even four additional third focal points at each of the Nasmith platforms.
- Nevertheless, the bent Cassegrain optics for an $f/8$ system requires several bents to fit the optics inside of the telescope hub.
- We have also initially explored the possibility of a reduced f/D system between 6.5 and 7, which offers the advantage of a better optical match to the instruments, and a more compact optics.



End