



An RC Design for the Atacama 25m Sub-Millimeter Telescope

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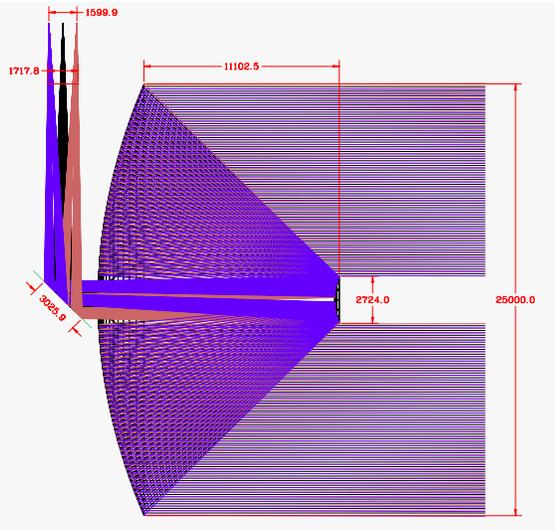
Atacama 25m Symmetric Design Parameters

Design	Ritchey-Chrétien / Nasmyth Focus	
Primary Diameter		25.0 m
Primary f/#		0.6
Primary Focal Length		15.0 m
Back Focal Distance		18.0 m
Secondary Diameter		2.642 m
Secondary Diameter for a 18 ' FOV		2.725 m
Sec-Primary Distance		13.429 m
System Magnification		20.00
Total f/#		12.0
Plate Scale		0.688 "/mm
Size of 18 arcmin FOV	18 '	157.1 cm
Diffraction Limit	@ 200 um	1.980 arcsec
Diffraction Spot Size	@ 200 um	2880.0 um





Atacama 25m Symmetric Design Geometry



On Axis Performance

	Uniform Illumination	Edge Taper -11 dB	
Wavelength:	200.0	200.0	[μm]
Frequency:	1499	1499	[GHz]
HPFW Beam Width:	1.841	2.122	[arcsec]
Aperture Strehl:	100.00	100.00	[%]
Polarization Efficiency:	100.00	100.00	[%]
Beam Efficiency:	79.18	88.74	[%]
Aperture Plane Efficiency:	99.70	76.99	[%]
Side Lobe Level (SLL):	-16.92	-26.24	[dB]
Cross-Pol Level:	-325.86	-327.31	[dB]
Antenna Gain:	-----	110.43	[dB]
Overall Antenna Efficiency:	-----	71.55	[%]



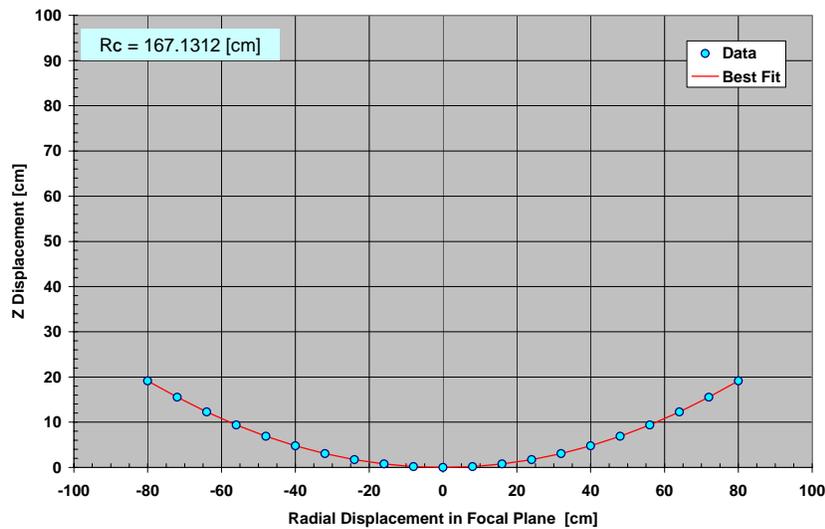


Edge of 18" FOV

	Uniform Illumination	Edge Taper -11 dB	
Wavelength:	200.0	200.0	[μm]
Frequency:	1499	1499	[GHz]
HPFW Beam Width:	1.841	2.139	[arcsec]
Aperture Strehl:	98.72	98.42	[%]
Polarization Efficiency:	99.99	99.99	[%]
Beam Efficiency:	76.06	88.20	[%]
Aperture Plane Efficiency:	98.29	75.30	[%]
Side Lobe Level (SLL):	-16.26	-24.78	[dB]
Cross-Pol Level:	-49.68	-52.55	[dB]
Antenna Gain:	-----	110.36	[dB]
Overall Antenna Efficiency:	-----	70.39	[%]



Best Focal Surface Geometry (200 μm)

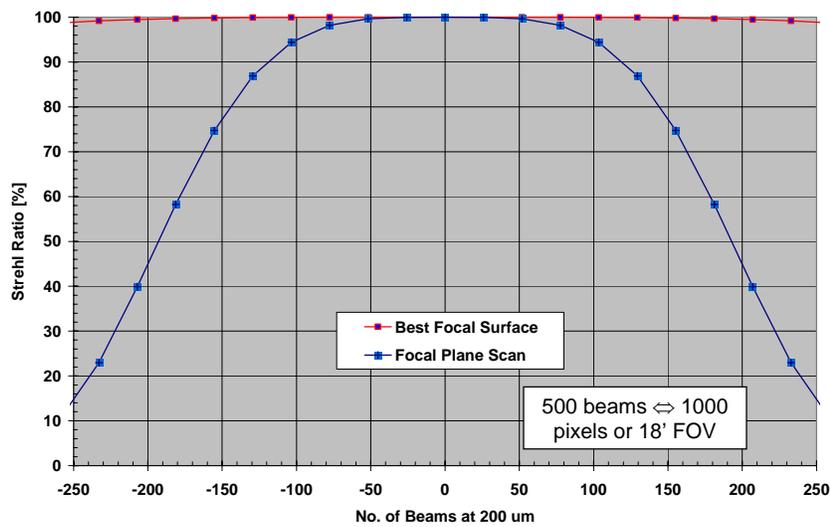




Strehl Ratio Variations in FOV



Strehl Ratio vs. No of Beams at 200 μm





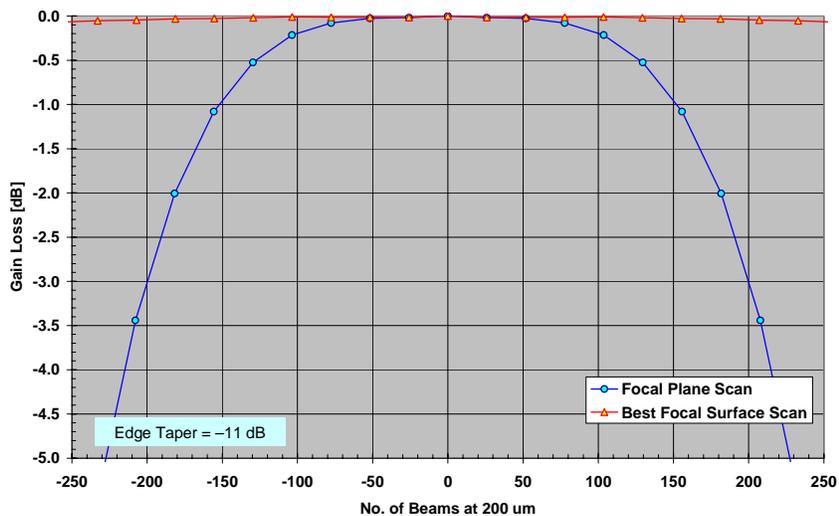
Overall Antenna Efficiency Variations in FOV

with -11 dB Edge Taper Illumination



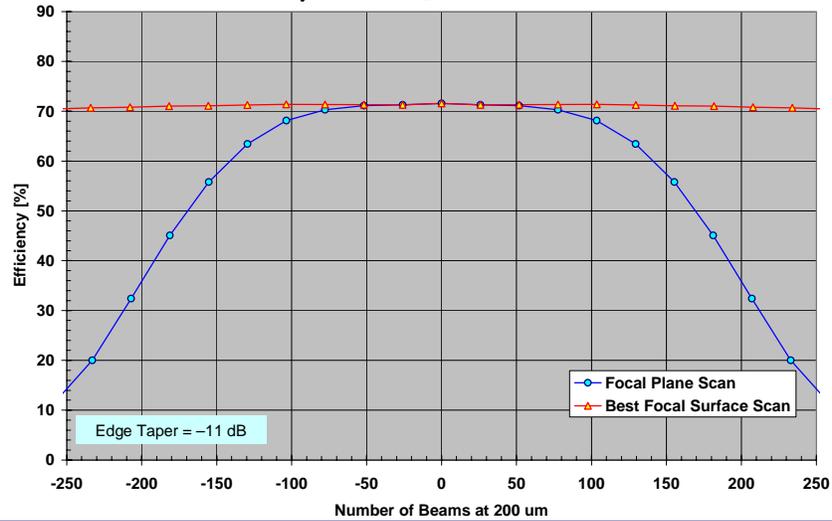
Gain Loss vs. No. of Beams at 200 μm

(Edge Taper -11 dB)





Overall Antenna Efficiency vs. No. of Beams at 200 μm (Edge Taper -11 dB)

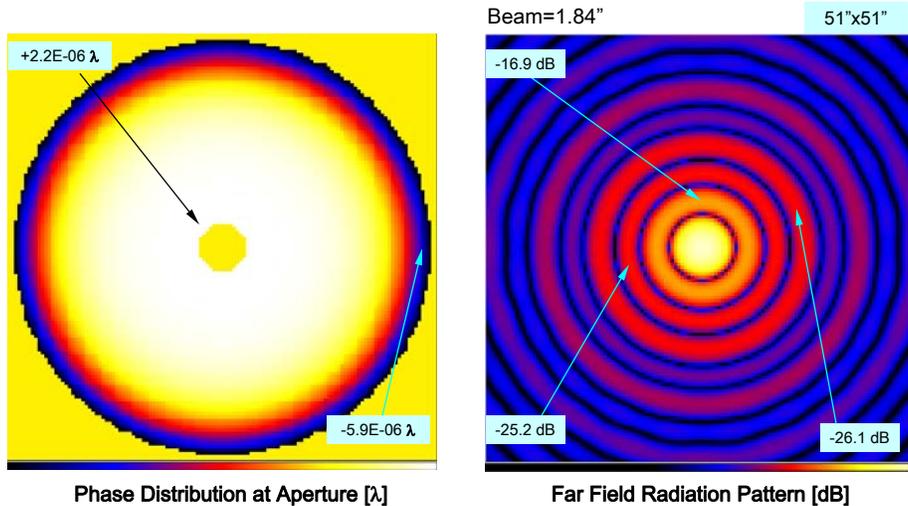


Aperture Phase Distribution and Far Field Beam Patterns

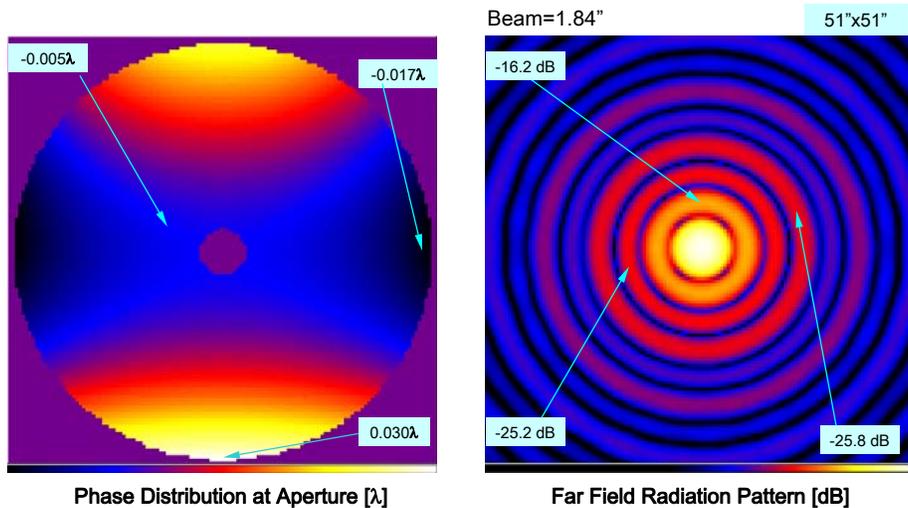
with Uniform Illumination



Aperture Phase Distribution and Co-Polar Beam Pattern at Center of FOV (at 200 μm , Uniform Illumination)



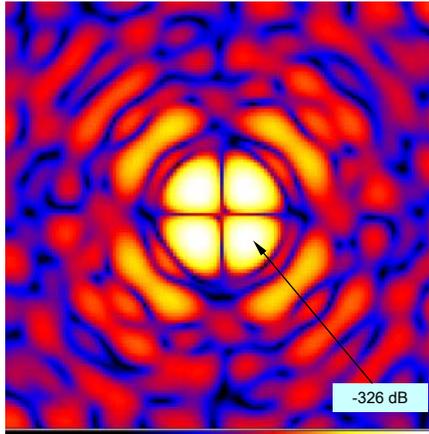
Aperture Phase Distribution and Co-Polar Beam Pattern at Edge of FOV (9' radius at 200 μm , Uniform Illum.)



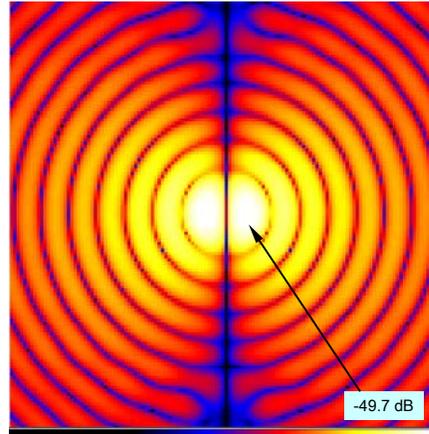
Cross-Polar Beam Patterns (at 200 μm , Uniform Illumination)

Center of FOV

Edge of FOV



51"x51"

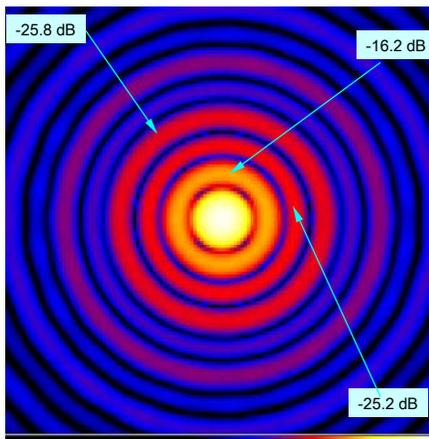


51"x51"

Comparison of Uniform vs. -11 dB Taper Illumination Co-Polar Beam Patterns (at 200 μm)

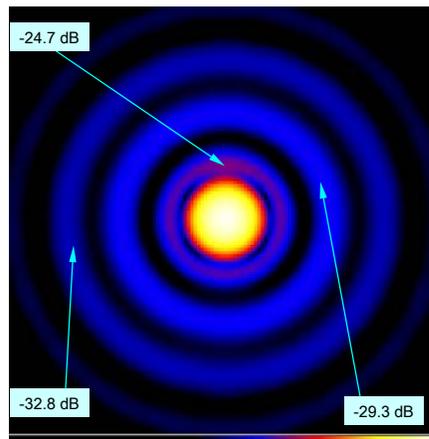
Beam=1.84"

Beam=2.12"



51"x51"

Uniform Illumination



51"x51"

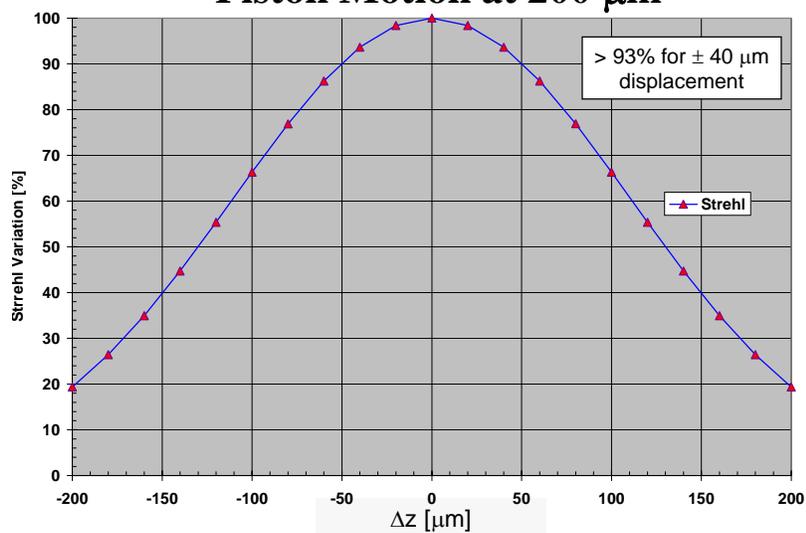
Edge Taper -11.0 dB



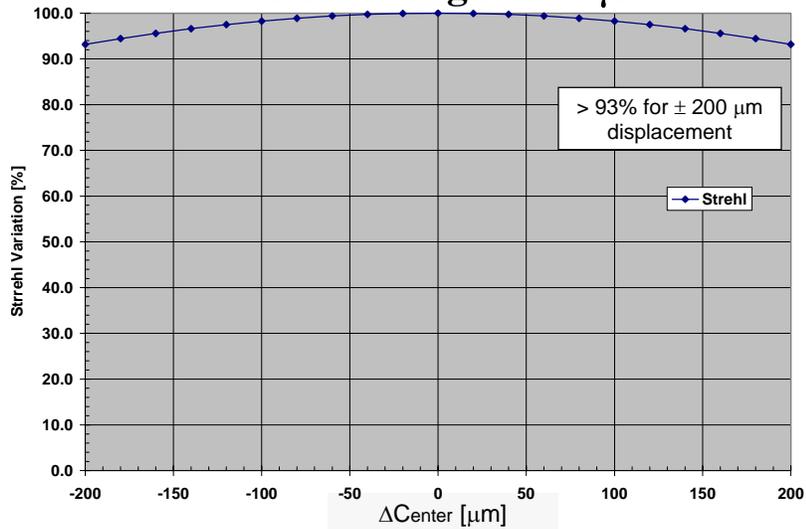
Sub-Reflector Piston, Tip/Tilt and De-centering Sensitivities



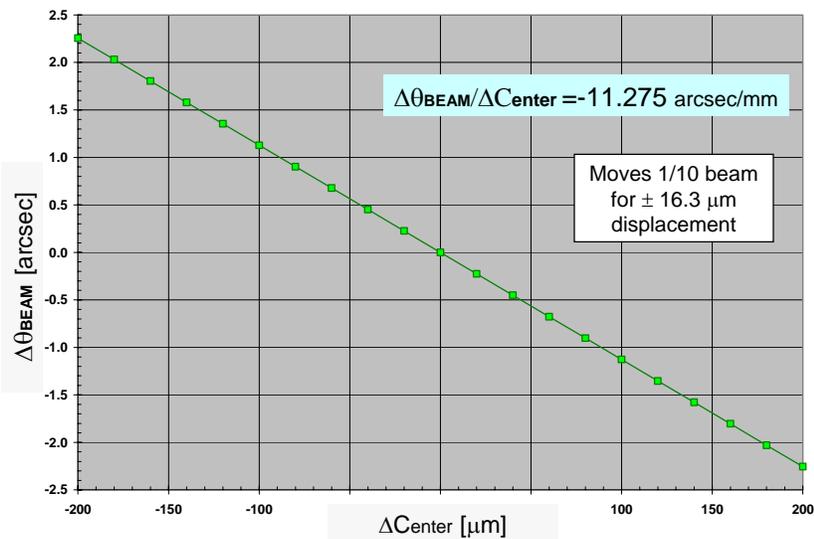
Strehl Ratio vs. Sub-Reflector Piston Motion at 200 μm



Strehl Ratio vs. Sub-Reflector De-centering at 200 μm

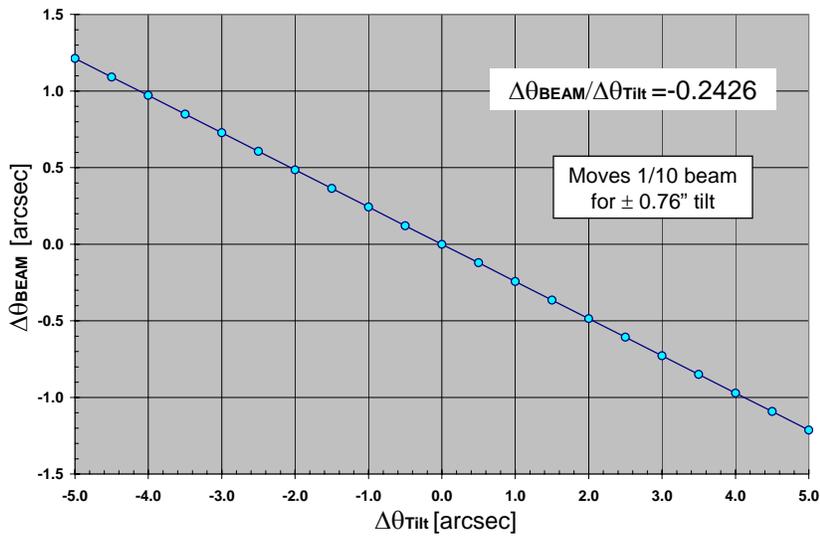


Beam Deviation vs. Sub-Reflector De-Centering





Beam Deviation vs. Sub-Reflector Tilt/Tip



Constraints on Sub-Reflector Motion

Sub-Reflector Motion

Max De-Focusing
 Max De-Centering
 Max Tip/Tilt

$\pm 40.0 \mu m$
 $\pm 16.3 \mu m$
 $\pm 0.76 \text{ arcsec}$

For Strehl < 93%
 $\Delta\theta_{BEAM \text{ MOTION}} < \text{HPBW} / 10$
 $\Delta\theta_{BEAM \text{ MOTION}} < \text{HPBW} / 10$



This translates to $\pm 10 \mu m$
 For a $\varnothing 2.7m$ Sub-Reflector





Future Work

- ◆ Surface roughness – baseline is 12 μm rms
 - ↗ Similar to goal of ALMA, achieved with SMA, CSO, nearly so with APEX
 - ↗ Ruze factor ~ 57% at 200 μm , and >83% at $\lambda > 350 \mu\text{m}$
- ◆ Effects of secondary and spider obstructions
- ◆ Effects of segmentation gaps
 - ↗ Point spread function
 - ↗ noise
- ◆ Off-axis designs...



Conclusions

Strawman design is a 25m Symmetric Reflector Sub-Millimeter telescope in a Nasmyth Ritchey-Chrétien configuration

- ◆ The Best Focal Surface is 1.6 m in diameter for a FOV of 18 arcmin and has a Radius of Curvature of 167 cm \Leftrightarrow 20 cm curvature at edge of field.
- ◆ We looked at a FOV of 18' to accommodate a 1000 \times 1000 (Nyquist Sampled) pixel array at 200 μm .
- ◆ The Strehl ratio is better than 98% over this FOV on the best focal surface, with uniform illumination of the aperture.
- ◆ An issue is detector size: 0.5 mm pixels \Rightarrow **0.5 meter (square) array!** (would be f/5 at 200 μm for Nyquist sampling)





Conclusions Cont...

- ◆ The Maximum Cross-polar level at the edge of FOV are -49 dB and -52 dB for uniform and Gaussian illumination, respectively.
- ◆ The Far Field Side-Lobe Level (SSL) over the FOV is < -16 dB with an uniform illumination, and better than -24 dB with a -11 dB Gaussian illumination taper.
- ◆ We calculated the sub-reflector sensitivities of the present design for focusing of piston motion, de-centering and tilt/tip.
 - ↗ Deviations in focusing of the sub-reflector $\Delta z_{\text{sub}} < 40 \mu\text{m}$ are required in order to maintain a Strehl ratio better than 95% at $200 \mu\text{m}$.
 - ↗ For a beam deviation of $\Delta\theta_{\text{BEAM}} \sim \theta_{\text{HPFW}}/10$ at $200 \mu\text{m}$:
 - ✦ the maximum dynamic de-centering of the sub-reflector should be less than $17 \mu\text{m}$



End

