### **HIFI**

Calibration Group

Core Program topical teams

# HIFI most wanted frequencies

Doc no: SRON/HIFI/TECH/2001-001

Inst no: SRON-G Issue: Draft 0.2 Date: 10 June 2003

Category: 2

## **TITLE:** HIFI most wanted frequencies

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Others			

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### DOCUMENT CHANGE RECORD

Date	Issue / revision	Page	Change
5 May 2003	Draft 0.1		First issue
2 June 2003	Draft 0.2	Tables	Updated table of lines and priorities according to core program topical team inputs

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#### I. Introduction

### I.1 Purpose

This document compiles the lines identified as most requested among the various topical core programs. It is meant at helping the test engineers and calibrators to decide at which frequencies the limited amount of test time available should be preferrably dedicated. A priority rank is attributed to each line, based mostly on scientific return. Most wanted lines and selection criteria are described for each of the core program topical teams.

#### I.2 References

I.2.1 Applicable documents

I.2.2 Reference documents

RD-01 HIFI Science Core Program, http://www.sron.rug.nl/hifiscience/CoreProgramme/

#### II. Most wanted lines

#### II.1 Selection criteria

This section compiles the selection criteria applied by each of the topical teams to build the lists presented in II.2.

### II.1.a Galaxies and galactic nuclei

The tables contain a limited amount of lines corresponding to Appendix A of the core program document related to the galaxies and galactic nuclei topics (RD-01). They are considered with the same priority level and are thus all assigned the highest priority.

### II.1.b Solar system bodies

The table includes the preferred lines for both planetary and cometary core programs. The justifications for the line choices can be found in the corresponding chapters of the core program description (RD-01). In the particular case of the comets, priority 3 samples H<sub>2</sub>O and HDO lines in the most sensitive frequency range of HIFI. Priority 2 contains other H<sub>2</sub>O and HDO (as well as isotopomers) required to study the excitation and o/p ratio. Priority 1 contains other lines of interest for the considered objects, as well as lines obtain at no extra cost in the image band.

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#### II.1.c Insterstellar medium

Priorities are assigned on the bias of scientific merit (3: high, 2: medium, 1: low).

### II.1.d Star formation

Most of the reported lines have priority 3 (highest priority), as listed in the appendixes of the Star Formation Program v.3 (2nd June), available at RD-01. A few lines are ranked priority 1 and are at low priority. The selection criteria are described in detail the proposal for each molecule or water line, for each subsection of the proposal. The general criteria are that the observations of the selected lines are not feasable, or with a lot of difficulty, with the present or future telescopes (including APEX and SOFIA).

#### II.1.e Stellar evolution

The priorities we asign to lines are mainly the same as the priorities we give in our proposal to the subprojects in which those lines are requested. Namely:

#### Water lines:

- o Priority 1 is given to lines included in two subprojects see Part I, Sects. 3.1 (observations of 55 lines in 4 selected sources) and 3.3 (observations of 20 lines in 20 sources). The first of these subprojects has our highest priority among water observations.
- o Priority 2 is given to lines to be observed only in the first project (Sect. 3.1).

#### Other lines

- O Priority 1 is given to lines included in our systematic observations of AGB, PPNe and PNe (Sect. 4, Part I), which has the highest priority among our subprojects. Most of these lines are also included in the systematic molecular observations of hot, massive stars (Sects. 5.4.2 and 5.4.4 Part II).
- o Priority 2: a few lines in Sects. 5.4.2 and 5.4.4 Part II that are not to be observed in evolved star envelopes, plus lines of hydrides in AGBs, PPNe and PNe (a second priority project).
- o Priority 3 is given to lines in our search of C[I] and C[II] emission (our 3rd priority project), except for lines included in Sects. 5.4.2 and 5.4.4 (Part II), which have 2nd priority.

#### II.2 Most wanted frequency tables

In the following, we list the most wanted lines for each of the HIFI bands. Note that some frequencies are uncertain.

# HIFI most wanted frequencies

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Species	Frequency	Priorities						
_	(GHz)	Galaxies	Solar System	ISM	Star	Stellar		
			bodies		formation	evolution		
$O_2$	487.2				3			
CS	489.751					1		
С	492.161	3		2	3	2		
$NH_2D$	494.454			1				
$NH_2D$	494.455			1				
$NH_2D$	494.457			1				
HDO	509.329		3 (comet)					
SiO	520.878					1		
H <sub>2</sub> O	530.343							
СН	532.721			2		2		
СН	536.761		1 (comet)	2	3	2		
H <sub>2</sub> O-para	546.621					1		
H <sub>2</sub> <sup>18</sup> O-ortho	547.676	3	2 (comet)	3	3			
$C^{18}O$	548.831							
<sup>13</sup> CO	550.926							
H <sub>2</sub> <sup>17</sup> O-ortho	552.021		2 (comet)		3			
H <sub>2</sub> O-ortho	556.936	3	1(planet)/3(comet)	3	3	1		
CN	566.942					1		
<sup>15</sup> NH <sub>3</sub>	572.113			1		2		
NH <sub>3</sub> -ortho	572.498		1 (comet)	2	3			
CO	576.268							
HDO	599.927		2 (comet)					
$D_2O$	607.350			1	3			
HCN	620.304					1		
H <sub>2</sub> O-ortho	620.701					2		
HCO <sup>+</sup>	624.205					1		
H <sup>37</sup> Cl	624.964			1				
H <sup>37</sup> Cl	624.978			1				
H <sup>37</sup> Cl	624.988			1				
H <sup>35</sup> Cl	625.901			1	3			
H <sup>35</sup> Cl	625.919			1				
H <sup>35</sup> Cl	625.932			1				

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Species	Frequency	Priorities					
	(GHz)	Galaxies	Solar System Bodies	ISM	Star formation	Stellar evolution	
H <sub>2</sub> O-para	645.834					2	
H <sub>2</sub> O-ortho	651.099					2	
SiH	658.009				3		
DF	660.000			1			
<sup>13</sup> CO	661.067					1	
СО	691.473					1	
$O_2$	715.4				3		
H <sub>2</sub> S-ortho	736.034		1 (comet)	1	1		
H <sub>2</sub> <sup>18</sup> O-ortho	745.320				3		
H <sub>2</sub> O-para	750.501					2	
H <sub>2</sub> O-ortho	750.572					2	
H <sub>2</sub> O-para	752.033		2 (planet & comet)	3	3	2	
H <sub>2</sub> O-para	766.860					2	
<sup>13</sup> CO	771.184						
$O_2$	773.8	_			3		

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Species	Frequency	Priorities					
•	(GHz)	Galaxies	Solar System Bodies	ISM	Star formation	Stellar evolution	
CO	806.652						
<sup>13</sup> C	809.1				3		
С	809.342	3		2	3	3	
CH <sup>+</sup>	835.071		1 (comet)	3	3	2	
H <sub>2</sub> CO	855.151				1		
H <sub>2</sub> O-para	859.859					2	
H <sub>2</sub> O-ortho	863.855					2	
H <sub>2</sub> O-para	863.876					2	
<sup>13</sup> CO	881.273					2	
HD <sup>18</sup> O	883.189			1			
LiH	887.0				3		
HDO	893.639		3 (comet)	3	3		
H <sub>2</sub> CO-o	896.7				1		
H <sub>2</sub> CO	896.805						
$D_2O$	896.947			1			
H <sub>2</sub> O-para	899.410					2	
H <sub>2</sub> O-ortho	902.564					2	
$H_2O$	906.206						
H <sub>2</sub> O-para	916.171				3	2	
СО	921.800					2	
CS	929.723					1	
CH <sub>2</sub>	930.		1 (comet)				
NH <sub>2</sub>	952.542			1			
NH <sub>2</sub>	959.526			1			

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Species	Frequency	Priorities					
	(GHz)	Galaxies	Solar System Bodies	ISM	Star formation	Stellar evolution	
H <sub>2</sub> O-para	970.315				3		
$OH^+$	971.804		1 (comet)	3			
NH	974.462		1 (comet)	3	3		
$OH^+$	984.532					1	
$H_3O^+$	984.6				3		
$H_3O^+$	984.697		1 (comet)	3			
$C^{18}O$	987.6		3				
H <sub>2</sub> O-para	987.927	3	2 (planet & comet)	3	3	2	
<sup>13</sup> CO	991.329		2				
H <sub>2</sub> <sup>18</sup> O-para	994.675	3			3		
NH	999.973			3			
H <sub>2</sub> O-ortho	1000.092					2	
$H_2S$	1002.779				1		
$C^{17}O$	1010.7		3				
NH <sup>+</sup>	1012.524			2			
CO	1036.912		2	2			
NH <sup>+</sup>	1038.1				3		
$H_2O_2$	1047.377		1				
HCN	1062.983					1	
HCO <sup>+</sup>	1069.697					1	
$H_2S$	1072.8		3				
C <sup>18</sup> O	1097.163				3		
H <sub>2</sub> O-ortho	1097.364		2	2	3	1	
H <sub>2</sub> O-para	1100.088					1	
$C^{13}O$	1101.350			3	3	1	
H <sub>2</sub> <sup>18</sup> O-para	1101.698	3		3	3		
H <sub>2</sub> <sup>17</sup> O-para	1107.167				3		
H <sub>2</sub> O-para	1109.586					1	
$H_2D^+$	1111.741				3		
H <sub>2</sub> O-para	1113.343	3	2 (planet & comet)	3	3	1	

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Species	Frequency	Priorities					
-	(GHz)	Galaxies	Solar System	ISM	Star	Stellar	
			Bodies		formation	evolution	
$O_2$	1120.715		1				
CN	1133.199					1	
$H_2^{18}O$	1136.7		2				
$H_2^{17}O$	1149.0		2				
CO	1151.985	3		2	3	1	
NO	1152.9		3				
H <sub>2</sub> O-ortho	1153.118				3	2	
H <sub>2</sub> O-ortho	1153.359					2	
H <sub>2</sub> O-para	1155.095					2	
H <sub>2</sub> O-ortho	1155.098					2	
H <sub>2</sub> O-ortho	1158.324				3		
H <sub>2</sub> O-ortho	1162.931		1	2	3	2	
H <sub>2</sub> O-ortho	1165.0				3		
NH <sub>3</sub>	1168.517			2		2	
SiO	1170.094					1	
H <sub>2</sub> O-para	1172.448					2	
$CO^+$	1178.766				3		
CO <sup>+</sup>	1179.039						
$O_3$	1180.3		1				
H <sub>2</sub> <sup>18</sup> O-ortho	1181.394				3		
H <sub>2</sub> CO	1185.0		1				
H <sub>2</sub> O-ortho	1196.859					1	
H <sub>2</sub> O-para	1205.825					1	
H <sub>2</sub> O-para	1207.666					1	
<sup>13</sup> CO	1211.330						
H <sub>2</sub> O-ortho	1214.615					1	
NH <sub>3</sub>	1214.859					2	
NH <sub>3</sub>	1215.245				1	2	
H <sub>2</sub> O-ortho	1215.830					1	
H <sub>2</sub> O-para	1215.961					1	
HDO	1217.3		1				
H <sub>2</sub> O-para	1228.799			2	3	1	
HF	1232.476			1	3		
HCl	1251.4		3				

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### **HIFI Band 6L**

Species	Frequency	Priorities					
•	(GHz)	Galaxies	Solar System Bodies	ISM	Star formation	Stellar evolution	
FeH	1411.1				3		
H <sub>2</sub> O	1440.782						
$N^+$	1461.134			2			
СН	1470.689				3		
$D_2H^+$	1476.6				3		
СН	1477.292						
CO	1496.923	3		2	3		
Н	1532.6				3		
<sup>13</sup> CO	1540.988						
СО	1611.793						
HDO	1625.408		2		3		
$H_2^{18}O$	1633.484				3		
H <sub>2</sub> O-ortho	1649.00				3		
$CO^+$	1649.309				3		
$CO^+$	1649.582						
<sup>13</sup> CO	1650.767			2			
$H_3O^+$	1655.813		1 (comet)	3	3		
H <sub>2</sub> <sup>18</sup> O-ortho	1655.868			3	3		
СН	1656.961			2			
$H_3O^+$	1657.236			3			
H <sub>2</sub> O-ortho	1661.011	3	2 (comet)	2		2	
СН	1661.107						
H <sub>2</sub> <sup>17</sup> O-ortho	1662.464				3		
$H_3O^+$	1663.584				3		
CH <sup>+</sup>	1669.1590		1 (comet)	3	3	2	
H <sub>2</sub> O-ortho	1669.905	3	2 (planet & comet)	3	3	2	

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### HIFI Band 6H

Species	Frequency	Priorities						
-	(GHz)	Galaxies	Solar System	ISM	Star	Stellar		
			Bodies		formation	evolution		
H <sub>2</sub> O-ortho	1713.940					1		
H <sub>2</sub> O-ortho	1716.765		1(planet)/2(comet)	2		1		
H <sub>2</sub> O-para	1717.037				3	1		
H <sub>2</sub> <sup>18</sup> O	1719.250				3			
CO	1726.602			2				
H <sub>2</sub> O-para	1753.173					2		
H <sub>2</sub> O-ortho	1753.888					2		
<sup>13</sup> CO	1760.486					1		
H <sub>2</sub> O-para	1762.131					2		
NH <sub>3</sub>	1763.525				1			
H <sub>2</sub> O-para	1766.121					2		
HCN	1769.876					1		
H <sub>2</sub> O-para	1794.753					1		
H <sub>2</sub> O-ortho	1794.735					1		
H <sub>2</sub> O-para	1794.830					1		
H <sub>2</sub> O-ortho	1797.238					1		
$O_3$	1808.082							
HDO	1818.530							
ОН	1834.735		2 (planet & comet)		3			
ОН	1837.747		2 (comet)		3			
СО	1841.346			2	3	1		
H <sub>2</sub> O-ortho	1866.301					2		
H <sub>2</sub> O-ortho	1867.825					2		
<sup>13</sup> CO	1870.141							
H <sup>13</sup> Cl	1873.410							
HCl	1876.230		3					
H <sub>2</sub> O-para	1879.720					2		
$C_3$	1890.558			2				
$C_3$	1896.706			2				
<sup>13</sup> C <sup>+</sup>	1900.1				3			
$\mathbf{C}^{+}$	1902.346	3		3	3	2		
H <sub>2</sub> O-ortho	1903.697					2		
H <sub>2</sub> O-para	1903.875					2		
H <sub>2</sub> O-para	1904.495					2		
$C_3$	1906.337			2				

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$C_3$	1914.274		2	
H <sub>2</sub> O-ortho	1918.477			2
H <sub>2</sub> O-para	1919.372			2
CO	1956.018			